

¹ INTERNAL GRAVITY WAVES FROM ² ATMOSPHERIC JETS AND FRONTS

³ Riwal Plougonven¹ and Fuqing Zhang²

R. Plougonven, Laboratoire de Météorologie Dynamique, Ecole Normale Supérieure, 24 rue Lhomond, 75005 Paris, France.

F. Zhang, Department of Meteorology, Pennsylvania State University, 601A Walker Building, University Park, PA 16802, U.S.A.

¹Laboratoire de Météorologie Dynamique, Ecole Normale Supérieure, IPSL, Paris, France, riwal.plougonven@polytechnique.org

²Department of Meteorology, Pennsylvania State University, University Park, Pennsylvania, fzhang@psu.edu

4 *For several decades, jets and fronts have been known from observations to*
5 *be significant sources of internal gravity waves in the atmosphere. Motivations*
6 *to investigate these waves have included their impact on tropospheric convec-*
7 *tion, their contribution to local mixing and turbulence in the upper-troposphere,*
8 *their vertical propagation into the middle atmosphere and the forcing of its*
9 *global circulation. While many different studies have consistently highlighted*
10 *jet exit regions as a favored locus for intense gravity waves, the mechanisms*
11 *responsible for their emission had long remained elusive: one reason is the*
12 *complexity of the environment in which the waves appear, another is that the*
13 *waves constitute small deviations from the balanced dynamics of the flow gen-*
14 *erating them, i.e. they arise beyond our fundamental, balanced understand-*
15 *ing of jets and fronts. Over the past two decades, the pressing need for im-*
16 *proving parameterizations of non-orographic gravity waves in climate mod-*
17 *els that include a stratosphere has stimulated renewed investigations. This re-*
18 *view aims at presenting current knowledge and understanding on gravity waves*
19 *near jets and fronts from observations, theory and modelling, and to discuss*
20 *and outline challenges for progress in coming years.*

1. INTRODUCTION

21 Internal gravity waves are waves occurring in the interior of a stratified fluid, with buoy-
 22 ancy providing the restoring force which opposes vertical displacements. Such waves are
 23 ubiquitous in the atmosphere and ocean and are the internal counterpart to the familiar
 24 surface gravity waves. In the atmosphere, they have horizontal scales ranging typically
 25 from 10 to 1000 km, and frequencies bound between the Coriolis parameter and the Brunt-
 26 Väisälä frequency (e.g. *Holton* [1992]). The highest frequencies occur for displacements
 27 that are nearly vertical, and high-frequency waves generally have shorter scales (simply
 28 reflecting that the forcing at high frequencies occurs at shorter scales). Amplitudes of
 29 internal gravity waves (or simply gravity waves, GW) generally are relatively small in the
 30 troposphere and stratosphere, in the sense that the dynamics at large scales (synoptic
 31 and larger) is well described by approximations based on a balance such as geostrophic
 32 balance, which allows to obtain the wind diagnostically from the rest of the flow. These
 33 *balanced* approximations (e.g. quasi-geostrophy) filter out GW by construction, and have
 34 provided much of our fundamental understanding of mid-latitude dynamics (e.g. *Val-*
 35 *lis* [2006]). For example, baroclinic instability was identified with the development of
 36 the quasi-geostrophic approximation [*Charney*, 1948; *Eady*, 1949], and frontogenesis with
 37 that of a higher-order approximation, semi-geostrophy [*Hoskins and Bretherton*, 1972].
 38 Nonetheless, gravity waves can be of importance and reach large amplitudes *locally*, and
 39 their importance grows as we move up into the stratosphere and mesosphere [*Andrews*
 40 *et al.*, 1987]. Indeed, as they propagate vertically and transfer momentum and energy
 41 from their origin (generally in the troposphere) to the level where they dissipate, they
 42 contribute to the circulation and variability in the stratosphere, and force the reversal of

the meridional thermal gradient in the mesosphere [*Fritts and Alexander, 2003*]. General circulation models (GCMs) that include a middle atmosphere generally do not have sufficient resolution to describe gravity waves explicitly, and hence need *parameterizations* to represent their main effects, namely the forcing due to the deposition of momentum where the waves are dissipated [*Kim et al., 2003*]. One major difficulty with present parameterizations of gravity waves is the specification of their *sources*, which can be an arbitrary, tunable parameter due to lack of physical understanding and observational constraints.

The main sources of gravity waves include orography, convection and jet/front systems. Flow over orography has long been known and studied as a source (e.g. *Queney [1948]*, see also references in *Gill [1982]*). Over the past two decades, several mechanisms have been proposed to explain waves generated by moist convection [*Clark et al., 1986; Fovell et al., 1992; Alexander et al., 1995*], paving the way for their parameterizations in General Circulation Models. Atmospheric jets and fronts are known from observations to also be a major source of gravity waves. Studies of gravity wave activity have showed a conspicuous enhancement of gravity wave activity in the vicinity of jets and fronts (e.g. *Fritts and Nastrom [1992]; Eckermann and Vincent [1993]*). In addition, numerous case studies have analyzed the occurrence of strong gravity wave events in the vicinity of a jet/front system. These case studies have isolated specific flow configurations: intense, gravity waves of low frequency have repeatedly been identified in the exit region of jets in the upper troposphere, often upstream of a ridge of geopotential [*Uccellini and Koch, 1987; Guest et al., 2000*]. Such waves of low frequency are often called *inertia-gravity waves*.

However, the exact mechanisms through which the waves are generated near jets and fronts remain an active area of current research and debate. Candidate mechanisms

66 associated with jet-front wave generation have included geostrophic adjustment, Lighthill
 67 radiation, unbalanced instabilities, transient generation, shear instability, convection. . .

68 Several of these can be considered examples of *spontaneous emission* [Ford et al., 2000],
 69 i.e. emission of gravity waves by a flow that initially was well *balanced* (e.g. in geostrophic
 70 balance). This highlights one reason for the slow progress in understanding waves gener-
 71 ated by jets and fronts: the latter are best known precisely in balanced approximations
 72 which by construction filter out gravity waves. Predicting gravity waves that will appear
 73 in flows that otherwise remain close to balance amounts to determining the limitations of
 74 these balanced approximations [Vanneste, 2013].

75 Recent years have brought significant progress in the understanding of mechanisms of
 76 spontaneous emission. Analytical studies have described Lighthill radiation, unbalanced
 77 instabilities and transient generation in simple flows and have provided simple asymptotic
 78 formulae quantifying the emitted waves [Vanneste, 2008]. In more complex flows, emission
 79 in a dipole has been simulated and quantitatively explained, providing a clear paradigm
 80 for emission in jet exit regions [McIntyre, 2009]. These studies have underlined the role
 81 of the background flow on the waves that are generated, and hence the importance of
 82 considering propagation effects.

83 With advances in computational power, it has been possible to complement these stud-
 84 ies with idealized simulations that describe flows of realistic complexity, starting with
 85 O’Sullivan and Dunkerton [1995]. The simulated flows consist in the development and
 86 saturation of the instability of a baroclinic jet in a stratified, rotating fluid. Such baro-
 87 clinic life cycles constitute a fundamental paradigm for our understanding of extratropical
 88 weather systems [Thorncroft et al., 1993]. Gravity waves emitted in such simulations share

89 features common with observational case studies. The background flow in which these
90 waves appear is still quite complex (fully three-dimensional, time-evolving), so that even
91 the origin of the waves is not always clear. A simpler flow has emerged as a paradigm
92 that retains enough complexity (localized wind maximum, i.e. a jet streak) to produce
93 analogous waves yet allow a quantitative explanation of their generation: it consists in a
94 dipole (one cyclone and one anticyclone of similar size and amplitude) that propagates
95 (quasi-)steadily [*Snyder et al.*, 2007; *Wang et al.*, 2009]. The emission of waves is under-
96 stood as the response, within a complex background flow, to residual tendencies, i.e. to
97 the small discrepancy between the full flow and its balanced approximation [*Snyder et al.*,
98 2009; *Wang and Zhang*, 2010].

99 This review will cover recent advances in many aspects of gravity waves from jets and
100 fronts and discuss their impacts and importance. The review will complement the earlier
101 review of *Uccellini and Koch* [1987] on observed gravity wave events associated with jet
102 streaks, and recent reviews of *Fritts and Alexander* [2003] on gravity waves and the middle
103 atmosphere, of *Kim et al.* [2003] on gravity wave parameterizations and of *Richter et al.*
104 [2007] that summarized findings and discussions from a gravity wave retreat held at NCAR
105 in the summer of 2006.

106 The paper is organized as follows: observational evidence for the emission of gravity
107 waves from jets and fronts is reviewed in sections 2 and 3, respectively covering clima-
108 tological studies that establish the general importance of jets and fronts as sources, and
109 case studies which provide insights on favorable flow configurations and characteristics of
110 the emitted waves. Many different generation mechanisms have been proposed in relation
111 to this problem and they are described in sections 4 and 6. Mechanisms that have ini-

112 tially been pinned down through analytical developments, yielding asymptotic results, are
 113 described first, in section 4. These theoretical results however do not connect straight-
 114 forwardly to gravity waves observed in real flows. Understanding the generation and
 115 maintenance of gravity waves in more realistic flows requires a preliminary consideration
 116 of propagation effects (section 5). This allows to consider the emission in laboratory and
 117 numerical experiments (section 6), which occur in more realistic flows and which have led
 118 to a consistent explanation of gravity waves in jet exit regions. Impacts and parameter-
 119 izations of waves generated from jets and fronts are presented in section 7. Our state of
 120 understanding and outstanding issues are discussed in section 8.

2. OBSERVATIONS:

CLIMATOLOGICAL STUDIES

121 Broadly, observational studies of relevance can be separated into two categories: cli-
 122 matological studies, which can describe, for example, the importance of storm tracks as
 123 source regions of gravity waves, and case studies, which provide specific examples of waves
 124 emitted from jets or fronts. The present section describes climatological studies, empha-
 125 sizing those that quantify waves not only geographically, but relative to the flow and in
 126 particular to jets and fronts.

127 The following is organized by observational platform. This is an opportunity to give
 128 an overview of observations available for gravity waves, and to present advantages and
 129 limitation of each type of observations.

2.1. Surface and radiosonde networks

130 Surface observational networks have been available for several decades, and have pro-
 131 vided the first opportunity for systematic climatological documentation and character-

132 ization of gravity waves. For example, *Einaudi et al.* [1989] performed a monthly-long
133 climatological study of gravity waves at the Boulder Atmospheric Observatory with data,
134 from a network of microbarographs and from sensors on the 300 m tower. They found
135 both coherent and incoherent motions within five frequency ranges with periods range
136 1-20 min. A similar but much more extensive climatology of gravity waves was performed
137 in *Grivet-Talocia et al.* [1999] using a mesonet network of barometers over east-central Illinois
138 during 1991-1995. They identified coherent events, for which clear propagation allowed
139 a good estimation of phase velocity, dominant period and horizontal wavelength. Part
140 of these coherent events are attributed to gravity waves, others corresponding to grav-
141 ity currents, solitary waves or bores. Coherent events were found to occur ~ 20 of the
142 total time in fall and winter and 12% in summer with dominant wave speed at at 25-30
143 m/s. They attributed the seasonal dependence of gravity wave occurrences to the stronger
144 baroclinicity of the atmosphere (and the mid-latitude jet streams).

145 The most comprehensive study of gravity waves using surface pressure observations was
146 presented in *Koppel et al.* [2000] who examined the distribution of large hourly pressure
147 changes (> 4.25 hPa) during a 25-yr period over the United States. They found the most
148 frequent occurrences of large-amplitude surface pressure changes are over the Great Plains
149 (which may be related to being in the lee of the Rockies) and over New England (that
150 are in the storm track and jet stream exit region), as illustrated in Figure 1. They also
151 found that the large-amplitude gravity wave activity is more prevalent over winter and
152 spring during the period of strong atmospheric baroclinicity. Their composite analysis
153 shows that the flow patterns are in broad agreement with the gravity wave paradigm of
154 *Uccellini and Koch* [1987] (see section 3).

155 Gravity waves have also been analyzed systematically using the radiosounding network:
 156 these studies contrast with the above by a stronger emphasis on the waves in the upper
 157 troposphere and lower stratosphere¹. Moreover, whereas surface barographs focused on
 158 fast waves in time-series (periods less than an hour), vertical profiles from radiosondes
 159 favor the analysis of low-frequency waves (also called *inertia-gravity waves*) that have an
 160 unambiguous signature in the hodograph [*Hirota and Niki*, 1985].

161 *Wang and Geller* [2003] used the high vertical resolution radiosonde wind and tem-
 162 perature data to examine the gravity wave climatology over the United States during
 163 1998-2001 (see their Figure 6). They found that the tropospheric and lower stratospheric
 164 gravity-wave energies are both stronger in winter than summer, likely owing to the pres-
 165 ence of stronger baroclinic jet-front systems. They found that tropospheric gravity-wave
 166 energy maximizes over the Rocky Mountains while the lower stratospheric energy maxi-
 167 mized over the southeastern United States. An intriguing result from *Wang and Geller*
 168 [2003] is that they found little correlation between the tropospheric and lower strato-
 169 spheric gravity-wave energies. *Gong et al.* [2008] further examined an 8-year climatology
 170 of gravity waves using the high-resolution radiosonde observations with the addition of a
 171 ray-tracing model. They found that the gravity wave sources are anisotropic, with wave
 172 momentum flux directed mostly upstream of the prevailing wind direction. Whereas a
 173 source tied to convection provided the best fit to the observations at low latitudes, a more
 174 general source worked better at middle and high latitudes. Investigations were pursued
 175 with a more sophisticated use of the energies that can be estimated from radiosondes
 176 [*Geller and Gong*, 2010], leading to a better estimation of convective sources [*Gong and*
 177 *Geller*, 2010].

178 Radiosondes have also been analyzed in other regions, in particular when specific cam-
179 paigns have made high-resolution profiles available. *Guest et al.* [2000] have analyzed
180 gravity waves in ozonesonde profiles reaching 30 km altitude over Macquarie Island, South
181 of New Zealand. The location guarantees that observed waves are non-orographic. Cases
182 with strong inertia-gravity waves were analyzed and led to the identification of a common
183 meteorological pattern: intense waves were found downstream of a jet streak, between
184 the inflection point and the ridge of geopotential. Using ray-tracing analysis, they con-
185 firmed that the origin of the waves observed in the lower-stratosphere originated from
186 a tropospheric jet-front system. *Sato and Yoshiki* [2008] examined stratospheric gravity
187 waves from 3-hourly radiosondes launched from Syowa station in Antarctica. Large and
188 sporadic gravity wave activity was observed during the winter months, with some events
189 of gravity waves generated from the Polar Night Jet, and propagating upward and down-
190 ward. *Zhang and Yi* [2007] have analyzed gravity waves in several years of twice-daily
191 radiosondes, from several stations in China, and also found that the upper-tropospheric
192 jet was the main source of waves. More precisely, they suggest the strong wind shear
193 induced by the jet as the source of waves [*Zhang and Yi*, 2005, 2008]. As *Guest et al.*
194 [2000], and in contrast to *Wang and Geller* [2003], *Geller and Gong* [2010] or *Sato and*
195 *Yoshiki* [2008], they found minimum stratospheric gravity wave activity in winter. These
196 differences regarding the seasonal cycle are yet unexplained.

197 *Plougonven et al.* [2003] took advantage of the large number of soundings launched from
198 research vessels over the Atlantic Ocean, far from orographic sources, during the FASTEX
199 campaign (January-February 1997, *Joly et al.* [1997]). Gravity wave activity was found
200 to be maximal in the vicinity of the jet stream. More specific analysis led to identify two

201 flow configurations for which intense gravity waves were present: the vicinity of a strong,
 202 straight jet, and the jet exit region of a strongly curved jet, either in a trough [*Plougonven*
 203 *et al.*, 2003] or in a ridge [*Plougonven and Teitelbaum*, 2003]. Only in the case of jet exit
 204 regions was it possible to carry out case studies of clear, intense inertia-gravity waves
 205 consistently identified in several soundings. The waves had frequencies between f and $2f$,
 206 wavelengths of a few hundred kilometers, wind perturbation of $5\text{-}8\text{ m s}^{-1}$, similar to *Guest*
 207 *et al.* [2000].

2.2. In-situ aircraft measurements

208 Analyzing in-situ measurements aboard commercial aircraft during 1978 and 1979, *Fritts*
 209 *and Nastrom* [1992] and *Nastrom and Fritts* [1992] attributed the mesoscale variance
 210 enhancements of horizontal velocity and temperature (presumably mostly induced by
 211 gravity waves) to four different mechanisms: topography, frontal activity, nonf-rontal
 212 convection, and wind shear. Overall, they found variances of temperature and wind at
 213 horizontal scales less than approximately 100 km to be 6 times larger in the vicinity of
 214 these sources than in a quiescent background, emphasizing the intermittency of gravity
 215 waves sources. The relative importance of the different sources were evaluated as shown
 216 in Figure 2, indicating strong values of variances above jets and fronts, smaller than those
 217 above topography by a factor of ~ 2 for wind speed, and comparable for temperature.

218 The interpretation of the small-scale component of winds and temperature as grav-
 219 ity waves received support from *Bacmeister et al.* [1996], who examined the horizontal
 220 wavenumber power spectra of 3-D wind velocities and potential temperatures measured at
 221 an altitude of about 20 km during 73 NASA ER-2 flights. They argued that the observed
 222 velocity and potential temperature spectra are consistent with gravity waves instead of

223 the inverse energy cascade of either 2-D or 3-D turbulence. However, this study did not
224 attempt to identify the sources of these gravity waves.

225 More recent aircraft based investigations of gravity waves emitted by jets and fronts
226 focus on individual case studies, and are discussed in section 3 and 7.2.

2.3. Balloons and rockets

227 Beyond the reaches of radiosondes and aircrafts, other in-situ measurements of gravity
228 waves in the middle atmosphere include balloons and sounding rockets, which are usually
229 part of coordinated field campaigns.

230 Ultra-long-duration, superpressure balloons drift on isopycnic surfaces and behave as
231 quasi-Lagrangian tracers, yielding a direct measurement of intrinsic frequencies which is
232 very valuable for gravity wave studies [*Hertzog et al.*, 2002b]. Whereas flights in the equa-
233 torial stratosphere during September 1998 showed largest momentum fluxes associated
234 with convection [*Hertzog and Vial*, 2001], campaigns in the winter polar vortices of both
235 hemispheres (2002, 2005) have allowed the investigation of other sources [*Vincent et al.*,
236 2007]. Topography (Greenland, Antarctic Peninsula) comes out strikingly as the source
237 associated with the maximum local values of momentum fluxes, but significant values are
238 also found over oceans and smooth terrain, indicating the importance of non-orographic
239 waves. Measurements from the Vorcore campaign (austral spring of 2005, *Hertzog et al.*
240 [2007]) were analyzed in detail using wavelet analysis [*Boccara et al.*, 2008]. *Hertzog et al.*
241 [2008] confirmed that non-orographic sources, although yielding locally weaker values, had
242 an overall contribution that was comparable to or greater than flow over orography (see
243 Figure 3). Complementary to this data analysis, *Plougonven et al.* [2012] have carried
244 out mesoscale simulations ($\Delta x = 20\text{km}$) of flows above the polar cap over a wide domain

245 and for a long period (2 months). Overall, a satisfactory, quantitative agreement was
 246 found between the simulated and observed gravity wave momentum fluxes, which is very
 247 encouraging [*Plougonven et al.*, 2010, 2012]. Nonetheless, specific biases for orographic
 248 and non-orographic waves were identified, emphasizing that it is preferable to analyze sep-
 249 arately these different types of waves. These simulations confirmed that the contribution
 250 fom non-orographic waves to the momentum fluxes integrated over the polar cap were
 251 comparable or larger than those of orographic waves for this domain and time, consistent
 252 with *Hertzog et al.* [2008] (Figure 3).

253 In-situ measurements above the upper stratosphere are obtained from sounding rock-
 254 ets. They allow investigation of gravity waves in the mesosphere and lower thermosphere,
 255 where their signatures can be very large. The winter MaCWAVE (Mountain and Convec-
 256 tive Caves Ascending Vertically) rocket campaign in January 2003 explicitly focused on
 257 gravity waves, using both ground-based and rocket-borne instruments [*Fritts et al.*, 2004;
 258 *Williams et al.*, 2006]. However, the rocket measurements have not been directly linked
 259 to jet-front gravity wave activity.

2.4. Ground-based remote sensing

260 Besides the aforementioned in-situ measurements, remotely sensed observations from
 261 ground-based radars, lidars and airglow are also widely used to detect atmospheric gravity
 262 wave activity. Among the ground-based remote sensing instruments, specially designed
 263 radars and lidars have the highest temporal and vertical resolution but they are only
 264 available at very limited number of locations around the world.

265 Using observations from an ST (stratosphere-troposphere) radar during four extended
 266 observational campaigns in southern Australia, *Eckermann and Vincent* [1993] examine

267 the generation of gravity waves from cold fronts. They found order of magnitude increases
268 in mesoscale variance of winds attributable to gravity waves during frontal passages. They
269 also found it possible to detect certain waves (in the upper-troposphere, with long hori-
270 zontal wavelengths and large ground-based phase speed) a day before and a day after the
271 fronts' arrival, whereas large amplitude, higher-frequency, shorter horizontal wavelength
272 waves are directly associated with the onset of the frontal circulation at the surface. They
273 speculated that the smaller wave amplitudes observed in the stratosphere may be due to
274 either more oblique propagation of wave energy in a more stable environment or due to
275 the ducting of wave energy below the tropopause.

276 The MU (middle and upper atmospheric) radar located at Shigaraki, Japan has been
277 providing measurements of gravity waves since 1984. *Sato* [1994] examined gravity wave
278 activity using wind data derived from this radar over 1986-1988, and found the dominant
279 waves in the lower stratosphere tend to have short vertical wavelengths (~ 4 km) and
280 long ground-relative periods (~ 10 h). The gravity waves are the strongest in winter
281 which is apparently related to the strong subtropical jet stream over this region [*Sato*,
282 1994]. However, she speculated that topography rather than jet/front systems may be
283 the primary sources of these gravity waves.

284 A mesosphere-stratosphere-troposphere (MST) radar located in Aberystwyth, Wales,
285 has been operated on a quasi-continuous basis since 1997. This VHF wind profiler radar
286 is capable of making continuous measurements of the three-dimensional wind vector at
287 high resolution, and was used for several case studies of inertia-gravity waves excited by
288 jets and fronts [*Pavelin et al.*, 2001; *Pavelin and Whiteway*, 2002], see section 3. *Vaughan*
289 *and Worthington* [2007] analyze inertia-gravity waves with eight years' observations from

290 this MST radar. They found inertia-gravity waves generally propagating upward in the
 291 lower stratosphere and downward in the troposphere, evidence that the source is at the
 292 jet-stream level. Long period waves ($> 12\text{h}$) were not preferentially associated with a jet
 293 stream and showed little seasonal dependence, in marked contrast with shorter period
 294 waves (6-8h) which were clearly associated with the jet and had a winter maximum.

295 The winter maximum in the stratosphere was also found in the 4-year gravity wave
 296 climatology derived from Rayleigh lidars located in two different sites in Southern France
 297 [*Wilson et al.*, 1991] and from a Rayleigh lidar in Japan *Murayama et al.* [1994]. They
 298 found strong correlation between the gravity wave activity and the wind speed in the
 299 stratosphere, and the waves increase in amplitude while propagating from stratosphere to
 300 the mesosphere.

301 Ground-based airglow imagers are often used as an economical means of measuring
 302 gravity waves in the mesosphere and lower thermosphere (e.g., *Taylor and Bishop* [1995];
 303 *Walterscheid et al.* [1999]; *Li et al.* [2011]). Airglows are mostly effective in detecting
 304 gravity waves with short periods ($< 1\text{ h}$), short horizontal wavelengths ($< 100\text{km}$) and
 305 long vertical wavelength ($> 10\text{km}$) (*Liu and Swenson* [2003]). For example, most recently,
 306 *Li et al.* [2011] documented a year-long climatology of gravity waves observed by an airglow
 307 imager over northern China. They found the gravity waves occurs more frequently over
 308 the summer and winter than the other two seasons. These waves have typical horizontal
 309 wavelengths of 10-35 km and phase speeds of 30-60 m/s.

2.5. Satellite observations

310 There have been a large number of observational studies of gravity waves from satellites
 311 since *Fetzer and Gille* [1994] using LIMS (Limb Infrared Monitor of the Stratosphere). On

312 one hand, the satellite-based remote sensing measurements a priori have relatively poor
313 spatial and temporal resolutions, but on the other they provide the most complete cov-
314 erage of global gravity waves and as such constitute an invaluable source of information.
315 *Wu and Waters* [1996] are one of the first to estimate global activity of gravity waves,
316 using MLS (Microwave Limb Sounder) observations. Further developments led to deriv-
317 ing global gravity wave *momentum fluxes* from satellite observations (*Ern et al.* [2004],
318 using temperature measurements from CRISTA (Cryogenic Infrared Spectrometers and
319 Telescopes for the Atmosphere)).

320 Overall synthesis of such observations can be found in the recent review paper of *Alexan-*
321 *der and et. al.* [2010]. Figure 8 of *Alexander and et. al.* [2010], adapted from *Preusse*
322 *et al.* [2008], summarizes the spatial and temporal resolution of different satellite-based
323 instruments which include infrared limb sounders, microwave sub-limb instruments, and
324 infrared nadir sounders². *Wu et al.* [2006] compared gravity waves measured by differ-
325 ent satellite instruments and found similar gravity wave characteristics from the nadir
326 techniques³. The limb-sounding instruments are complementary, having better vertical
327 but poorer horizontal resolution. Higher resolution and accurate measurements of grav-
328 ity waves can be achieved with the more recently launched HIRDLS instruments (e.g.
329 *Alexander and co authors* [2008]; *Yan et al.* [2010]).

330 These satellite studies provide global distributions, and hence some information on
331 gravity waves generated from jets and fronts. One study which specifically tied gravity
332 wave activity to the tropospheric baroclinic jet front systems is presented in *Wu and Zhang*
333 [2004] using the AMSU-A microwave data (see Figure 6). They particularly focused on
334 the gravity wave properties and variabilities over the northeastern United States and the

335 North Atlantic in the December-January periods. It is found that gravity waves in this
 336 storm-track exit region, found in many winters, can reach the stratopause with growing
 337 amplitude. More importantly, this is one of the first studies that directly linked the
 338 satellite-derived gravity wave activity with the intensity and location of the tropospheric
 339 baroclinic jet front systems. *Wu and Eckermann* [2008] further show strong seasonal
 340 fluctuations of the global gravity wave variance derived from Aura MLS for each month
 341 of 2006 (their Figure 8).

342 Estimates of GW momentum fluxes or temperature variances have shown consistently, in
 343 different studies, enhanced values in the stratospheric winter polar night jet (e.g. *Wu and*
 344 *Eckermann* [2008]; *Alexander and co authors* [2008]; *Yan et al.* [2010]; *Ern and Preusse*
 345 [2011], see Figure 4). This can be interpreted as a signature of significant sources (oro-
 346 graphic and non-orographic) in the winter mid-latitudes, but also as the signature of
 347 favored propagation within the positive shear of the strong westerlies [*Dunkerton*, 1984;
 348 *Ern and Preusse*, 2011]. The zonal asymmetries are indications of enhanced sources, and
 349 emphasize orography as a source at mid and high latitudes [*Wu and Eckermann*, 2008].
 350 Interestingly, *Wu and Eckermann* [2008] use the different sensitivity of their instrument
 351 between ascending and descending orbits to show that waves in the mid-latitudes have a
 352 preferred horizontal orientation, with phaselines extending from south-west to north-east
 353 in the Northern Hemisphere, and from north-west to south-east in the Southern Hemi-
 354 sphere. This is consistent with the momentum fluxes estimated over the Southern ocean
 355 from balloon observations [*Hertzog et al.*, 2008] and numerical simulations [*Plougonven*
 356 *et al.*, 2012].

2.6. Analyses and forecasts from meteorological models

357 With increased model resolution and advanced model physics, along with enhanced ob-
358 servations and improved data assimilation methods, numerical weather prediction (NWP)
359 models are increasingly capable of resolving at least part of the gravity wave spectra in the
360 troposphere and stratosphere in both their analyses and forecasts. Even at moderate reso-
361 lutions, relevant information about the location and intrinsic frequency of gravity waves
362 can be obtained [*Plougonven and Teitelbaum*, 2003]. *Wu and Eckermann* [2008] showed
363 the monthly-mean GW-induced temperature variances at 44 km pressure altitude derived
364 from operational global analysis fields of the European Center for Medium-Range Weather
365 Forecast (ECMWF) Integrated Forecast System in August 2006 (Figure 4). They found
366 qualitative agreement between the gravity wave variances in terms of latitude bands and
367 propagation directions derived from the global analysis and those derived from satellite
368 observations (Aura MLS). At least part of the enhanced gravity wave activity over the
369 Southern Oceans is likely related to strong baroclinic jet-front systems during this winter
370 month.

371 Consistency of the satellite-derived jet-stream-related gravity waves with those from
372 a NWP model was also shown in *Wu and Zhang* [2004], but with the AMSU-A and a
373 higher-resolution mesoscale model (see section 3 and Figure 6). *Schroeder et al.* [2009]
374 also found good agreement between gravity wave-induced temperature fluctuations de-
375 rived from satellite observations (SABER, Sounding of the Atmosphere using Broadband
376 Emission Radiometry) and the ECMWF analysis, including those at the edge of the winter
377 polar vortex or the midlatitude jet streams.

378 More recently, *Shutts and Vosper* [2011] presented an indepth comparison of the gravity
379 wave fluxes derived from both the Met Office and ECMWF forecast models for August

380 2006 with those from HIRDLS. They concluded that the state-of-the-art NWP models are
 381 capable of capturing the correct overall strength and distribution of gravity wave activity.
 382 In the Southern mid and high latitudes, they note that waves tend to have phaselines
 383 oriented from North-West to South-East, consistent with *Wu and Eckermann* [2008].
 384 *Plougonven et al.* [2012] used observations from the Vorcore balloons *Hertzog et al.* [2007]
 385 to systematically assess the realism of the gravity wave field in a mesoscale meteorological
 386 model. Relative to the observations⁴, the simulations overestimated orographic waves by
 387 a factor $\sim 2-3$, whereas non-orographic waves were slightly underestimated (factor ~ 0.8
 388 for the time-averaged value).

389 These recent investigations advocate NWP models as a relevant means to document
 390 the global variations and impacts of gravity wave activity and fluxes [*Alexander and et.*
 391 *al.*, 2010]. Combining NWP output with observations and a careful assessment of biases
 392 and limitations of each promises to lead, in coming years, to a converging estimation of
 393 gravity wave momentum fluxes.

3. OBSERVATIONS AND MESOSCALE MODELLING: CASE STUDIES

394 In contrast to the climatological studies above, individual case studies isolate specific
 395 configurations in which intense gravity waves are unambiguously identified. They are
 396 described below in section 3.1. Finally, an overview of the observational studies is given
 397 in section 3.2, discussing the limitations and biases of the different observational platforms,
 398 and the needs for future observations.

3.1. Case studies

399 Tropospheric jets and fronts were long hypothesized to be responsible for numerous
400 observed gravity wave events, both in the troposphere [*Tepper*, 1951] and in the up-
401 per atmosphere above the tropopause [*Hines*, 1968]. However, given the limitation in
402 the observing techniques, there were inherent uncertainties in the source attribution of
403 these earlier observations [*Hines*, 1968; *Gossard and Hooke*, 1975]. Below we review case
404 studies starting from the review of *Uccellini and Koch* [1987]. Whereas early studies
405 emphasized tropospheric (ducted) waves, the focus over the last decade has shifted to
406 upper-tropospheric waves propagating into the stratosphere.

407 *Uccellini and Koch* [1987] (hereafter UK87) reviewed 13 long-lived observed lower-
408 tropospheric gravity wave events in literature (see refs therein). These mesoscale distur-
409 bances have wave periods of 1-4h, horizontal wavelengths of 50-500km and surface pressure
410 perturbations of 0.2-7 mb, all of which have been shown to influence the mesoscale struc-
411 ture of precipitation systems. They found a common synoptic environment for the gener-
412 ation and maintenance of these waves as being in the exit region of upper-level jet streaks
413 and cold-air side of a surface frontal boundary (Figure 5). They further hypothesized that
414 these gravity waves are likely to be generated by the unbalanced upper-tropospheric jet-
415 front systems through geostrophic adjustment [*Rossby*, 1938; *Cahn*, 1945; *Blumen*, 1972],
416 and to be maintained through wave ducting (*Lindzen and Tung* [1976], and section 5.1).

417 Case studies in the years following UK87 increasingly involved mesoscale numerical
418 modelling. The earliest simulations of mesoscale gravity waves using numerical weather
419 prediction models were first conducted by *Zhang and Fritsch* [1988]; *Schmidt and Cotton*
420 [1990] and *Cram et al.* [1992]. Gravity waves in these studies were generated by the simu-
421 lated mesoscale convective systems. However, detailed verification of these waves against

422 mesoscale observations was not performed due to the unavailability of the mesoscale data
 423 sets. Mesoscale numerical models have subsequently been developed into powerful tools
 424 for the detailed study of gravity wave structure, energy sources, and maintenance mech-
 425 anisms, all of which are difficult to detect with standard observations.

426 The first published attempt to use a mesoscale model for the sole purpose of simulating
 427 and studying an observed gravity wave event, and for which verification was performed
 428 against detailed mesoanalysis, was provided by *Powers and Reed* [1993]. The case simu-
 429 lated was the long-lived, large-amplitude gravity wave event on 15 December 1987 over the
 430 Midwest of the US which is believed to have created life-threatening blizzard conditions
 431 with peak pressure falls up to 11mb in 11 min as documented in *Schneider* [1990]. *Powers*
 432 *and Reed* [1993] concluded that the mesoscale NWP model used can successfully simulate
 433 mesoscale gravity waves and can capture many aspects of the observed waves in terms of
 434 both timing and magnitudes. Although this event had characteristics of mesoscale gravity
 435 waves under typical synoptic settings conceptualized by Uccellini and Koch (1987), the au-
 436 thors suggested the model waves were maintained and amplified by wave-CISK processes.
 437 *Powers* [1997] further concluded that elevated convection above a stable wave duct was
 438 the forcing mechanism in the model. *Pokrandt et al.* [1996], who studied the same case
 439 also with numerical simulations, on the other hand hypothesized that a transverse circu-
 440 lation about the approaching jet streak produced a mesoscale potential vorticity anomaly
 441 at midlevels that subsequently forced the mesoscale waves.

442 One of the cases reviewed in UK87 is the 11-12 July 1981 gravity wave event that is be-
 443 lieved to be responsible for triggering and organizing mesoscale convection over southeast
 444 Wyoming into the Dakotas during CCOPE [*Koch and Golus*, 1988; *Koch and Dorian*,

1988; *Koch et al.*, 1988, 1993]. There are at least two distinct wave episodes detected by the CCOPE high-resolution surface mesonet [*Koch and Golus*, 1988]. The synoptic-scale analysis in *Koch and Dorian* [1988] showed that the waves are confined to the region between the axis of inflection and the ridge in the 300 hPa height field, downstream of a jet streak and to the cold air side of a surface quasi-stationary front. There is also evidence of strong flow imbalance associated with the upper-level jet from observational analysis [*Koch and Dorian*, 1988] and from mesoscale modeling [*Kaplan et al.*, 1997]. Subsequent numerical simulations by *Zhang and Koch* [2000] and *Koch et al.* [2001] did simulate reasonably well the observed gravity waves. However, these latter studies concluded that, despite the proximity of the wave generation with the jet streaks, the thermally-driven mountain-plains circulation (MPS) is responsible for the generation of both wave episodes: the first through an orographic density current relegated from a remnant daytime MPS circulation [*Zhang and Koch*, 2000] and the second by convection triggered by the developing MPS [*Koch et al.*, 2001].

The relevance of the UK87 paradigm has been highlighted in a number of case studies and shown to be robust for the presence of waves (e.g. *Ramamurthy et al.* [1993]). Often it is found that the waves have an impact on convection and precipitation [*Trexler and Koch*, 2000; *Richiardone and Manfrin*, 2003], although the relation varies. This impact has been one motivation for the development of an automated system for predicting and detecting mesoscale gravity waves using surface observations [*Koch and O’Handley*, 1997; *Koch and Saleeby*, 2001]. Both studies suggest the hypothesis that the unbalanced flow in the jet streak exit region or near frontal boundaries is associated to mesoscale gravity wave generation.

468 Another well-studied case is the 1992 St. Valentine’s Day mesoscale gravity wave event
 469 observed during STORM-FEST [*Trexler and Koch, 2000; Rauber et al., 2001*]. High-
 470 resolution mesoscale NWP models had been used to simulate the event with varying
 471 degrees of success, while the mechanisms derived from different simulations differ greatly.
 472 Through unbalanced flow diagnosis of the model simulations, *Jin [1997]* and *Koch and*
 473 *O’Handley [1997]* believe this event followed closely the jet-gravity wave paradigm of
 474 UK87, though as in previous studies, *Jin [1997]* also finds convection is important for
 475 maintaining and amplifying the mesoscale waves. Through numerical experiments with
 476 and without evaporative processes, *Jewett et al. [2003]*, on the other hand, singled out
 477 the importance of the evaporatively driven downdrafts that impinges upon the surface
 478 warm-frontal inversion on the wave genesis.

479 Whereas observations alone have recurrently been insufficient to support conclusions
 480 on the relation of gravity waves and convection (e.g. *Ralph et al. [1993]*), high-resolution
 481 mesoscale simulations in complement to observations can provide key insights. A large-
 482 amplitude gravity wave event over the northeastern United States on 4 January 1994
 483 was documented in *Bosart et al. [1998]* that showed wavelengths of 100-200km and peak
 484 crest-to-trough pressure falls exceeding 13 hPa within 30 min associated with short-term
 485 blizzard conditions. The synoptic-scale pattern of this wave event is again consistent
 486 with the UK87 paradigm from the observational analysis. Through successful simulation
 487 of this event with a high-resolution mesoscale model, *Zhang et al. [2001, 2003]* demon-
 488 strated the radiation of the gravity waves to the lower-troposphere from an unbalanced
 489 upper-tropospheric jet streak. The wave packet emitted from the upper-level jet streak
 490 subsequently merged with a mid-tropospheric cold-front aloft and triggered moist convec-

491 tion. A ducted wave-CISK mode was responsible for the subsequent wave maintenance
492 and amplification. Hence, although moist processes were not at the origin of the wave,
493 they played a crucial role to amplify it, as shown by dry simulations.

494 It is worth noting that a number of case studies fall outside the flow configuration of the
495 UK87 paradigm. For example, *Ralph et al.* [1999] described gravity waves found ahead of a
496 cold front, suggesting that the cold front plays the role of an obstacle to the flow impinging
497 on it. These waves are very similar to some of the waves simulated in idealized studies of
498 frontogenesis (see section 6). The flow pattern in this case was significantly constrained by
499 the presence of mountains to the West of the cold front, and further investigations would
500 be necessary to determine whether this 'obstacle effect' of cold fronts was exceptional, or
501 commonly occurs.

502 The above case studies have focused on tropospheric waves, their interactions with
503 convection and their effects near the surface. The flow configuration identified by UK87
504 has also been found to be relevant for emission into the lower stratosphere. *Guest et al.*
505 [2000] have highlighted the jet exit region of a jet streak approaching the inflection point
506 between the base of a trough and a ridge as a configuration conducive to intense gravity
507 waves in the lower stratosphere. Ray-tracing was used to identify the origin of clear,
508 intense inertia-gravity waves observed in the lower stratosphere, and has highlighted the
509 upper-level jet as the region of emission [*Guest et al.*, 2000; *Hertzog et al.*, 2001]. Case
510 studies based on FASTEX radiosoundings also highlighted jet exit regions, either upstream
511 of a ridge *Plougonven and Teitelbaum* [2003] or upstream of a deep trough *Plougonven*
512 *et al.* [2003]. Instances of generation from jets in a region a priori dominated by orographic
513 waves were documented by *Spiga et al.* [2008]. They combined global reanalysis, satellite

514 and radiosoundings data along with mesoscale model simulations in the Andes Cordillera
 515 region to identify the cases where, respectively, the jet-stream source, the convective source
 516 and the topography source are predominantly in action.

517 Case studies focusing on upper-tropospheric and lower-stratospheric observations have
 518 often emphasized the presence of both upward and downward waves from the jet as
 519 a distinctive signature of emission by the jet *Thomas et al.* [1999]; *Plougonven et al.*
 520 [2003]; *Wu and Zhang* [2004]; *Spiga et al.* [2008]. From 17 radiosoundings launched at 3-
 521 hour intervals over Northern Germany, *Peters et al.* [2003] clearly identified inertia-gravity
 522 waves propagating upward and downward from the jet which amplified downstream of the
 523 jet streak. Complementing similar radiosonde observations with mesoscale simulations,
 524 *Zülicke and Peters* [2006] investigated the spontaneous generation of waves from the
 525 upper-level jet streak in a poleward-breaking Rossby wave. They identified subsynoptic
 526 (horizontal wavelength $\lambda_h \sim 500\text{km}$) and mesoscale waves ($\lambda_h \sim 500\text{km}$), and showed
 527 the waves to propagate upward and downward from the level of the jet stream. Their
 528 study provides further evidence that the jet exit region is hereby the key feature of the
 529 background flow. Numerical simulations have also been carried out in complement to
 530 satellite observations by *Wu and Zhang* [2004]. A good level of agreement was found
 531 between the waves interpreted from radiance perturbations east of Newfoundland, and the
 532 simulated waves (see Figure 6). Such comparison serves both to validate the interpretation
 533 of the observations and to assess the realism of the model.

534 First systematic measurements of upper-tropospheric and lower-stratospheric gravity
 535 waves with a dedicated research aircraft conducted during the 2008 field experiment of
 536 Stratosphere-Troposphere Analyses of Regional Transport (START08; *Pan et al.* [2010]).

537 During one of the research flights, accompanied with a strong baroclinic jet-front across
538 the continental United States, apparent activity of gravity waves at different scales near
539 or just above the tropopause region were sampled during nearly the entire flight mission
540 that covered a distance of a few thousand kilometers. While research is still ongoing
541 to examine the sources of these gravity waves observed during START08, it is apparent
542 the tropospheric jet-front systems, in interaction with the local topography and moist
543 convection, were playing essential roles in the forcing and characteristics of these gravity
544 waves [Zhang *et al.*, 2009].

545 Regarding generation mechanisms, case studies have often referred to geostrophic ad-
546 justment (e.g. Pavelein *et al.* [2001]). The justification is that observed and simulated
547 GW are often found in the vicinity or just downstream of regions of imbalance, with La-
548 grangian Rossby numbers serving as an indicator of imbalance [Koch and Dorian, 1988;
549 Ramamurthy *et al.*, 1993; Spiga *et al.*, 2008]. A more sophisticated indicator is provided
550 by the residual of the nonlinear balance equation Zhang *et al.* [2000, 2001], and has been
551 used efficiently (e.g. Hertzog *et al.* [2001]). However, the relation is merely a collocation
552 (the waves are found where or near maxima of indicator of imbalance), but it is not sys-
553 tematic (e.g. there are other maxima that are not associated to waves) and a quantitative
554 relationship is still lacking.

3.2. Limitations and challenges

555 Observational estimates of gravity wave activity or momentum fluxes face several dif-
556 ficulties: first, each observational platform has its own limited resolution (spatial and/or
557 temporal), making it sensitive only to a certain portion of the gravity wave spectrum (see
558 Preusse *et al.* [2008]; Alexander and *et al.* [2010]). Second, each observational platform

559 has limitations in terms of spatial and/or temporal coverage. For instance, high-resolution
 560 radiosondes describe *in situ* gravity waves with low frequencies and with vertical wave-
 561 lengths of a few kilometers, but each station only samples one location and with a limited
 562 frequency. In contrast, satellite observations can provide a nearly global coverage, but
 563 with limited spatial resolution and significant assumptions used in the process of convert-
 564 ing, say, radiance anomalies to momentum fluxes.

565 Numerical simulations have also been used to explore the uncertainties in the current
 566 gravity-wave observing techniques. For example, *Zhang et al.* [2004] examined the uncer-
 567 tainties in the commonly used hodograph method in retrieving inertio-gravity wave char-
 568 acteristics from individual vertical profiles of the winds. Analysis of mesoscale numerical
 569 simulations of a gravity wave event in which a seemingly coherent quasi-monochromatic
 570 inertia-gravity wave packet showed that important uncertainties were found to exist for all
 571 the wave characteristics derived from single vertical profiles using the hodograph method.
 572 Large uncertainties were found in particular in estimating derived quantities such as hor-
 573 izontal wavelengths. Similar approaches can be performed to assess the uncertainties in
 574 the gravity wave observations by in-situ or remotely sensing instruments (reviewed in
 575 section 2). One such example is presented in *Wu and Zhang* [2004] which compared
 576 the mesoscale simulations of gravity waves with those derived from space-borne sense on
 577 AMSU (see Figure 6).

578 Observations have provided substantial evidence for the importance of jets and fronts
 579 as sources of gravity waves and case studies have identified flow configurations favorable
 580 to the presence of significant waves. Two limitations need to be mentionned: one is that
 581 observations identify where gravity waves are found, not necessarily where they are gen-

erated. Second, case studies may introduce a bias towards cases that lend themselves well
to case studies, i.e. where conspicuous gravity waves (large amplitude, large enough scale
and time span that the wave can be identified, say, in several radiosondes...) that can
be well identified and interpreted. Generally, perturbations that occur on smaller scales,
and in particular those that are tied to moist convection, prove more difficult to interpret
beyond statistical approaches (e.g. *Fritts and Nastrom* [1992]; *Eckermann and Vincent*
[1993]). Now, as described above a number of case studies have emphasized the possible
role of moist processes in generating or amplifying waves near fronts. Gravity waves
directly generated by convective cells will have clearly higher intrinsic frequencies (and
shorter horizontal scales) than waves excited dynamically by jets and fronts. Nonetheless,
clarifying the contribution of moist processes to waves in the vicinity of jets and fronts
calls for dedicated research efforts.

A first challenge, that is presently being addressed given the maturity of observational
gravity wave studies (in particular from satellites), will be to make the different analyses
of the gravity wave field converge [*Alexander and et. al.*, 2010]. Comparisons of estimates
from different satellites [*Ern and Preusse*, 2011], between satellites and analyses *Shutts*
and Vosper [2011], or between mesoscale simulations and balloon observations *Plougonven*
et al. [2012] provide indications on the biases of these different sources of information, and
suggest that these different estimates may soon converge. A second challenge is to define
and obtain a complete description of the useful characteristics of the gravity wave field:
whereas mean momentum fluxes have very much been emphasized, they are not the only
relevant quantity. For example, the intermittency of the wave field also matters, and this
may be described through the probability distribution function of the momentum fluxes

605 [*Hertzog et al.*, 2012]. A final challenge will consist in extracting information on the wave
 606 sources from a combination of observations and simulations. Investigation of the gravity
 607 wave field relative to the flow (both the tropospheric flow which may act as a source, and
 608 the stratospheric flow which acts as a background) will be a path to help identify sources,
 609 going beyond geographical and seasonal variations.

610 Despite the availability of near continuous 4-dimensional model output, difference be-
 611 tween different modeling studies of the same events highlight the difficulties in pinpointing
 612 the forcing and generation mechanisms. These difficulties have at least partially driven
 613 the need for more idealized simulations with simpler flow patterns, which will be described
 614 in section 6.

4. GENERATION MECHANISMS:

ANALYTICAL RESULTS

615 This section and section 6 review theoretical studies of generation mechanisms that have
 616 been invoked to explain gravity waves in the vicinity of jets and fronts. The present sec-
 617 tion restricts mainly to analytical studies⁵ and hence simple flow configurations, allowing
 618 asymptotic results. This section is complemented, in Section 6, by a review of studies for
 619 which laboratory or numerical experiments have been a necessary component, providing
 620 an examination of emission in more realistic flows.

621 The observational evidence for a strong enhancement of gravity waves in the vicinity
 622 of jet/front systems has been one motivation for investigations of dynamical mechanisms
 623 generating gravity waves from predominantly *balanced* features of the flow. Another fun-
 624 damental motivation has been to identify the limitations of balanced approximations, i.e.

625 to determine when the evolution of the flow, while remaining predominantly balanced,
626 includes the *spontaneous* generation of gravity waves.

The fundamental difficulty for the emission is the scale separation between the slow balanced motions and the fast gravity waves, making it difficult for both types of motions to interact. The Rossby number measures this separation of the time scales: balanced motions evolve on the advective time-scale L/U , whereas the longest time scale for gravity wave motions is $1/f$. Their ratio yields the Rossby number

$$Ro = \frac{U}{fL}, \quad (1)$$

627 which is typically small for mid-latitude flows. To a very good approximation, atmospheric
628 and oceanic motions at small Rossby numbers are balanced, i.e. a diagnostic relation can
629 be established between the wind and other variables. The simplest balance relation is
630 geostrophic balance, but there are more accurate relations (e.g. *Hoskins et al.* [1985];
631 *Zhang et al.* [2000]). Additionally, the flows considered in this review nearly all have
632 aspect ratios justifying hydrostatic balance in the vertical (e.g. *Vallis* [2006]). These bal-
633 ances provide diagnostic relations which can reduce the number of time-derivatives in the
634 system: balanced approximations such as quasi-geostrophy provide a simple description
635 of the balanced flow, consisting of an inversion relation and a *single* prognostic equation,
636 the advection of the materially conserved potential vorticity *Hoskins et al.* [1985]. Other
637 motions, such as gravity waves, have been filtered out. Balanced models have provided
638 much of our understanding of mid-latitude dynamics and are helpful for initialization
639 issues in numerical weather forecasting (e.g. *Kalnay* [2003]). The occurrence of gravity
640 waves in the vicinity of jets and fronts constitutes a deviation from balance.

641 First we describe geostrophic adjustment, because it has very regularly been invoked
 642 (section 4.1). Studies of geostrophic adjustment address how an initial imbalance projects
 643 onto gravity waves, but not the origin of the imbalance. The discussion on the relevance of
 644 geostrophic adjustment in the present context is deferred to section 8.1. Next we describe
 645 explicit examples of spontaneous emission (or spontaneous adjustment emission, SAE),
 646 mechanisms explicitly addressing how balanced motions excite, in the course of their evo-
 647 lution, gravity waves: Lighthill radiation (section 4.2.2), unbalanced instabilities (section
 648 4.3) and transient generation (section 4.4). Further studies of spontaneous emission, in
 649 more realistic flows, are discussed in section 6. Finally, generation mechanisms involving
 650 shear instability are discussed in section 4.5.

4.1. Geostrophic adjustment

651 Geostrophic adjustment occurs when a rotating fluid is forced away from a balanced
 652 state on timescales that are short relative to the inertial timescale. The process forcing
 653 the fluid away from balance need not be specified: for example a wind burst forcing the
 654 upper ocean [*Rossby*, 1938], heating due to convection [*Schubert et al.*, 1980], an absorbed
 655 gravity wave [*Zhu and Holton*, 1987], or mixing due to shear instabilities [*Bühler et al.*,
 656 1999]. It only matters that this forcing be fast relative to the inertial timescale, so that
 657 it can be considered instantaneous, yielding the classical initial value problem. More
 658 generally, this is only a special case of the adjustment to a time-dependent local body
 659 forcing [*Weglarz and Lin*, 1997; *Chagnon and Bannon*, 2005a]. Below, we reserve the
 660 term ‘geostrophic adjustment’ for the classical initial value problem with geostrophy as
 661 the underlying balance.

662 The classical problem of geostrophic adjustment describes how an arbitrary initial con-
 663 dition, in a rotating fluid subject to gravity, splits into a geostrophically balanced part
 664 that remains and inertia-gravity waves which propagate away [*Rossby*, 1938; *Cahn*, 1945;
 665 *Obukhov*, 1949]. *Rossby* [1938] considered as an initial condition a rectilinear current in
 666 the upper layer of the ocean, with limited horizontal extent and with no surface height
 667 anomaly. Hence the initial current is out of balance and the fluid adjusts so as to find a
 668 state in which velocity and pressure (here surface height) are in geostrophic balance and
 669 which preserve the potential vorticity and mass relative to the initial state. The excess
 670 energy contained in the initial condition is shed off, in the form of inertia-gravity waves
 671 that propagate away.

672 Studies on geostrophic adjustment have focused on configurations for which the problem
 673 is well-posed:

674 1. if all motions are small perturbations to a state of rest, the adjustment can be de-
 675 scribed asymptotically in Rossby number [*Blumen*, 1972]. To leading order, the balanced
 676 part of the flow is described by quasi-geostrophic dynamics for Burger number of order
 677 unity [*Reznik et al.*, 2001].

678 2. if the flow is rectilinear or axisymmetric, the separation is again unambiguous be-
 679 cause the balanced part of the flow, even for large Rossby numbers, has a trivial time
 680 evolution: it is stationary. Adjustment has been investigated for purely zonal flows (e.g.
 681 *Rossby* [1938]; *Yeh* [1949]; *Ou* [1984]; *Kuo and Polvani* [1997]; *Kuo* [1997]; *Zeitlin et al.*
 682 [2003]) and axisymmetric flows (e.g. *Paegle* [1978]; *Schubert et al.* [1980]; *Kuo and Polvani*
 683 [2000]). In both cases, the unambiguous separation made it possible to describe analyt-

684 ically nonlinear adjustment (e.g. *Glendening* [1993]; *Blumen and Wu* [1995]; *Wu and*
 685 *Blumen* [1995]; *Plougonven and Zeitlin* [2005]).

686 Note that in both cases, the initial imbalance is prescribed. The origin of this imbalance
 687 lies outside the scope of these studies. They only describe the response of the fluid, in
 688 certain limited configurations (small perturbations to a state of rest (1), or symmetric
 689 flows (2)).

690 Numerous aspects of the geostrophic adjustment problem have been studied, e.g. the
 691 dependence of the response to the scale of the initial perturbation (e.g. *Matsumoto* [1961];
 692 *Blumen and Wu* [1995]; *Kuo* [1997]), or the interpretation of geostrophic adjustment as
 693 a minimization of energy for a given potential vorticity distribution [*Vallis*, 1992]. With
 694 the emission from jets in mind, *Fritts and Luo* [1992] have considered, in a stratified fluid
 695 at rest, initial imbalances having dimensions comparable with those of a jet stream. They
 696 found emitted waves that have low frequencies, consistent with the dispersion relation
 697 and the spatial scales of the prescribed imbalance. Their first, two-dimensional study was
 698 complemented by consideration of three-dimensional imbalances having long scales in the
 699 along-jet direction *Luo and Fritts* [1993].

700 In all of the examples above, the gravity waves originate from the initial, prescribed
 701 imbalance, and hence these examples provide little insight into generation from balanced
 702 motions. The geostrophic adjustment problem was in fact used to investigate the interac-
 703 tions of gravity waves and balanced motions: in the first several orders of the asymptotic
 704 theory, *Reznik et al.* [2001] showed a complete decoupling of the balanced motions and
 705 gravity waves (see also *Dewar and Killworth* [1995]), yielding an unambiguous separation,
 706 and hence no spontaneous emission [*Zeitlin*, 2008].

707 Now, various diagnostics of flow imbalance, as surveyed in *Zhang et al.* [2000], have
708 been widely and successfully used to identify the sources of gravity waves with respect to
709 the balanced flow (e.g., *O'Sullivan and Dunkerton* [1995]; *Jin* [1997]; *Zhang et al.* [2001]).
710 In consequence, 'geostrophic adjustment' has very often been referred to explain emitted
711 waves near jets and fronts (e.g. *O'Sullivan and Dunkerton* [1995]). In a related study of
712 an idealized baroclinic life cycle, and in order to emphasize the differences with classical
713 geostrophic adjustment, *Zhang* [2004] proposed the term *balanced adjustment*⁶ to describe
714 the spontaneous generation of gravity waves from a predominantly balanced flow that
715 continuously produces imbalance (as can be diagnosed from the residual of the nonlinear
716 balance equation for example), with an associated, continuous emission of gravity waves.
717 The investigation of this mechanism relies heavily on numerical simulations and will be
718 described in section 6.

4.2. Lighthill radiation

719 It is preferable to briefly recall the context in order to understand the change in paradigm
720 between the previous section and the present one.

721 4.2.1. A foreword on slow manifolds

722 The atmosphere and oceans are and remain so close to a balanced state on synoptic
723 scales that the existence of a *slow manifold* [*Lorenz*, 1980; *Leith*, 1980] was suggested
724 and investigated: within the phase space of the full equations, this would be an invariant
725 subspace of reduced dimensionality containing only balanced dynamics (for more general
726 definitions, see discussions in *Warn et al.* [1995] and *Ford et al.* [2000]). Investigating
727 whether such a manifold exists is equivalent to investigating whether motions that are at

728 one initial time purely balanced (or more precisely on the slow manifold) can produce, in
 729 the course of their evolution, unbalanced motions, i.e. gravity waves.

730 Several lines of evidence have progressively shown that such emission is inevitable,
 731 i.e. that an exactly invariant slow manifold in fact does not exist and that one should
 732 rather think slow manifolds of various accuracies (*MacKay* [2004]; *Vanneste* [2013] and
 733 references therein). One line of evidence came from low-order models such as the Lorenz-
 734 Krishnamurty model [*Lorenz*, 1986; *Lorenz and Krishnamurty*, 1987] describing with 5
 735 Ordinary Differential Equations (ODEs) the interactions of 3 slow vortical modes and 2
 736 fast gravity wave modes⁷. The divergence of perturbative procedures [*Vautard and Legras*,
 737 1986; *Warn and Ménard*, 1986], numerical simulations [*Lorenz and Krishnamurty*, 1987;
 738 *Camassa*, 1995; *Bokhove and Shepherd*, 1996], and exponential asymptotics *Vanneste*
 739 [2004] have demonstrated the spontaneous generation of fast motions is inevitable(Figure
 740 7). *Vanneste* [2004] has explicitly quantified the emission in this model as exponentially
 741 small in Rossby number, i.e. of a form involving $e^{-\alpha/Ro}$, with a prefactor that involves
 742 algebraic powers of the Rossby number Ro .

743 A second line of evidence comes from mechanisms of spontaneous emission identified in
 744 full flows, i.e. described by a system of Partial Differential Equations. The first is Lighthill
 745 radiation, and constitutes an explicit example of spontaneous generation (section 4.2.2).
 746 Two other mechanisms of SAE are unbalanced instabilities (section 4.3) and transient
 747 generation in shear (section 4.4).

748 **4.2.2. Lighthill radiation**

749 'Lighthill radiation' of gravity wave motions by balanced vortical motions [*Ford*,
 750 1994a, b, c] is analogous to the radiation of acoustic waves by turbulent vortical motions

described by *Lighthill* [1952]. The analogy is straightforward for the non-rotating shallow water equations which are equivalent to the two-dimensional equations for gas dynamics, with gravity waves replacing acoustic waves, and the Froude number $F = U/\sqrt{gH}$ replacing the Mach number $M = U/c_s$, where c_s is the sound speed. The inclusion of rotation inhibits the emission of waves, as frequency matching between the vortical motion and the inertia-gravity waves only occurs for $Ro > 1$ [Ford, 1994a]. The smallness F allows asymptotic investigation of the problem, and has an essential implication regarding the scale of the waves: the excited gravity waves having frequencies matching those of the balanced motions, of order U/L , the dispersion relation for shallow water waves imposes that they have spatial scales $\lambda \sim L/F \gg L$. Hence there is a scale separation between the small balanced motions and the large-scale gravity waves that are emitted.

Many aspects of the emission can be summarized by rearranging the equations of motions in such a way as to obtain a wave equation on the left-hand side (lhs), forced by nonlinear terms on the right-hand side (rhs) Ford [1994c]; Ford *et al.* [2000]:

$$\left(\frac{\partial^2}{\partial t^2} + f^2 - g h_0 \nabla^2 \right) \frac{\partial h}{\partial t} = \frac{\partial^2}{\partial x_i \partial x_j} T_{ij}, \quad (2)$$

where h is the height of the surface, h_0 is the height at rest, f is the Coriolis parameter and T_{ij} result from the combination of the nonlinear terms of the equations. In itself, this rearrangement does not prove anything [Snyder *et al.*, 1993; Plougonven *et al.*, 2009]. When one adds assumptions on the regime parameter, as above ($Ro > 1, F \ll 1$), one deals with Lighthill radiation: as the waves are large-scale and as the small-scale balanced motions are supposed to occur only in a compact region, it is appropriate to consider that the waves are propagating on the background of a fluid at rest and that the forcing is a point, quadrupolar source. The quadrupolar nature of the forcing implies,

770 in this setting, significant destructive interferences and hence weak emission (order F^2 ,
 771 [*Ford et al.*, 2000]).

772 As *Ford et al.* [2000, 2002] emphasized, one key feature of Lighthill radiation were
 773 that the emission is weak enough that the source can be described without taking the
 774 emission into account, e.g. from a balanced model. The lhs of (2) being the standard
 775 equation for gravity waves for a fluid at rest, standard intuitions apply: for example,
 776 Fourier transforms [*Ford*, 1994c] can be used to isolate the part of the rhs forcing that
 777 produce gravity waves (frequencies larger than f). Matched asymptotic expansions or
 778 Green's functions can be used to solve the forced problem [*Ford*, 1994a, b; *Ford et al.*,
 779 2000]. Gravity wave emission by balanced motions was investigated in rotating shallow
 780 water for unstable modes of axisymmetric vortices [*Ford*, 1994a], for the emission by an
 781 elliptic vortex [*Ford*, 1994b], for arbitrary localized balanced motions [*Ford et al.*, 2000]
 782 and for the roll-up of an unstable shear layer [*Ford*, 1994c]. In the latter case, numerical
 783 simulations were used to describe the small-scale vortical motions, and knowledge of the
 784 resulting forcing, averaged in the streamwise direction, was successfully used to predict
 785 the large-scale inertia-gravity waves in the far field (see Figure 8).

786 The analysis of Lighthill radiation was extended to a continuously stratified fluid for the
 787 emission by an ellipsoidal vortex [*Plougonven and Zeitlin*, 2002]. The radiative instability
 788 of an axisymmetric vortex [*Ford*, 1994a] and the evolution of the elliptic vortex [*Ford*,
 789 1994b; *Plougonven and Zeitlin*, 2002] can be interpreted as a coupling of Rossby waves on
 790 the PV gradient on the edge of the vortex [*Brunet and Montgomery*, 2002] with inertia-
 791 gravity waves in the far-field. The emitted waves are found to scale as F^2 , and hence the
 792 backreaction on the vortical motions only occur on very slow timescales (F^{-4}).

Rankine vortices were used for the above studies, for analytical tractability. In more realistic cases, when the vortices have a continuous distribution of PV, mixing at a critical level in the skirt of the vortex may inhibit the growth of these radiative instabilities [Schecter and Montgomery, 2006]. The regime of parameters for Lighthill radiation make it relevant for strong supercell mesocyclones and hurricanes (Schecter [2008] and ref. therein).

The study of Lighthill radiation was recently extended with numerical experiments to carry out a systematic parameter sweep [Sugimoto *et al.*, 2008], and also to spherical geometry Sugimoto and Ishii [2012].

4.3. Unbalanced instabilities

Unbalanced instabilities (also called non-geostrophic or ageostrophic instabilities) are instabilities of a balanced flow that involve unbalanced motions, typically gravity waves. These constitute a mechanism for spontaneous emission, provided that there is an initial deviation, however small, from the balanced flow under consideration [Vanneste, 2008].

A flow for which unbalanced instabilities have received considerable attention is an unbounded vertical shear above a flat surface. The quasi-geostrophic solution of Eady [1949] was extended beyond the Eady cutoff by Stone [1970] and Tokioka [1970], independently. The spatial structure of the modes they obtained was elucidated by Nakamura [1988], who showed that the modes changed character through the inertial-critical level (ICL) present in the flow. At the ICL, the Doppler shifted wave period is equal to the inertial period. The stability analysis was extended to nonzero meridional wavenumber l by Yamazaki and Peltier [2001a, b]. The growth rates of these modes [Molemaker *et al.*, 2005] and the spatial structure [Plougonven *et al.*, 2005] were both revisited recently. The

815 unstable modes consist of an Eady edge wave between the ground and the ICL, and of
 816 sheared gravity waves above (see Figure 9). A WKBJ approximation can give an accurate
 817 description of the normal mode, including its exponentially small growth rate (Vanneste,
 818 personal communication).

819 Unbalanced instabilities can involve different types of waves, from IGW (e.g.
 820 *Plougonven et al.* [2005]) to Kelvin waves (e.g. *Kushner et al.* [1998]), and have been
 821 identified in different flows: two layer sheared flow [*Sakai*, 1989], sheared flow over a slope
 822 [*Sutyrin*, 2007, 2008], horizontal shear [*Vanneste and Yavneh*, 2007], stratified Taylor-
 823 Couette flow [*Yavneh et al.*, 2001; *Molemaker et al.*, 2001], vortices [*LeDizès and Billant*,
 824 2009], a front of potential vorticity [*Dritschel and Vanneste*, 2006], elliptical instability
 825 [*McWilliams and Yavneh* [1998]; *Aspden and Vanneste* [2009]].

826 A strong motivation for the study of these various unbalanced instabilities has come
 827 from the suspicion that they play a significant role in the ocean interior (see section 7.4),
 828 in the forward cascade necessary to transfer energy from the anisotropic, balanced large-
 829 scale flow down to more nearly isotropic flows leading to dissipation [*McWilliams et al.*,
 830 2001].

831 Another motivation has been to better understand the dynamics of two-layer systems
 832 encountered in laboratory experiments (see section 6.1). Instabilities coupling a Rossby
 833 wave and a Kelvin wave in a two-layer rotating fluid were recently revisited with an
 834 emphasis on their nonlinear development [*Gula et al.*, 2009a]. Simulations suggested
 835 that the instability saturated early on, leaving behind only a limited signature of gravity
 836 waves. Recently, careful laboratory experiments (see Section 6.1) have provided the first
 837 evidence of these instabilities in real flows, and confirmed the weakness of their growth.

838 The stability study of realistic (continuous) frontal states [*Snyder, 1995*] also provides
 839 evidence for the weakness of unbalanced instabilities.

4.4. Transient generation by sheared disturbances

The evolution of potential vorticity anomalies in a horizontal shear [*Vanneste and Yavneh, 2004*] leads to a transient generation of gravity waves. This differs from the unbalanced instabilities described above in several respects: **1)** the generation of gravity waves occurs at a specific time; **2)** the final amplitude of the waves can be predicted within the linear theory⁸, and **3)** it is not necessary to include, in the initial condition, a small perturbation *Vanneste* [2008]. *Vanneste and Yavneh* [2004] quantified the emission of gravity waves for a sheared disturbance at one along-shear wavenumber, and demonstrated that the final amplitude of the waves is proportional to

$$\varepsilon^{-1/2} \exp(-\alpha/\varepsilon)$$

840 . As for the spontaneous generation in the Lorenz-Krishnamurty model *Vanneste* [2004],
 841 exponential asymptotics were necessary to describe this exponentially weak emission. The
 842 solutions obtained for one wavenumber can be combined (as a Fourier decomposition) to
 843 describe the emission by localized features of the flow such as a sheared vortex [*Olafsdottir*
 844 *et al.*, 2008].

845 The transient generation of sheared potential vorticity anomalies in a vertical shear
 846 was calculated by *Lott et al.* [2010] for 2D and *Lott et al.* [2012b] for 3D anomalies.
 847 These two studies illustrate transient emission in a vertical shear, and hence can be
 848 read as a vertical counterpart of *Olafsdottir et al.* [2008]. However, they are based on a
 849 modal (Fourier) decomposition, and hence can be read as a counterpart of *Plougonven*
 850 *et al.* [2005] without a surface. The same equation for the vertical structure of modes

851 is solved in both cases, what differs are the boundary conditions, leading to unstable
 852 modes when a lower boundary is present, and neutral modes when no boundary is present
 853 but a Dirac- δ PV anomaly is included. Physically, the key process in both cases is the
 854 coupling, by differential advection, of balanced motions and gravity waves, on one and
 855 the other side of an inertial critical level (ICL). Despite this commonality, the results
 856 on shear disturbances [*Vanneste and Yavneh*, 2004] appear very different from those on
 857 unbalanced instabilities in the same flow [*Vanneste and Yavneh*, 2007]. This is in part due
 858 to the different approaches used, i.e. nonmodal versus modal. The relationships between
 859 the different approaches, in a vertical shear, are discussed by *Mamatsashvili et al.* [2010].

860 This highlights the connections between the different mechanisms of spontaneous emis-
 861 sion. Both unbalanced instabilities and transient emission fundamentally rely on shear to
 862 connect motions that have different intrinsic timescales.

863 The transient emission of gravity waves by sheared regions have been investigated also
 864 in different contexts, to determine what gravity wave response could be expected from a
 865 stochastically perturbed shear layer or jet [*Lott*, 1997; *Bakas and Farrell*, 2008, 2009a, b].
 866 Investigation of momentum transport by gravity waves in a stochastically forced jet has
 867 shown for instance that the jet not only passively filters waves, but also amplifies portions
 868 of the spectrum, leading to possibly significant decelerations.

4.5. Shear Instability

869 Another possible route for the excitation of GWs from jets and fronts involves shear
 870 instabilities on small scales⁹. In the course of frontogenesis, both near the surface and at
 871 upper-levels, very intense shear layers are produced, potentially leading to shear instability
 872 (e.g. *Snyder* [1995]; *Esler and Polvani* [2004]). As such, this constitutes a mechanism for

873 spontaneous emission; however, the scales of shear instability are short enough that it has
874 generally been considered in non-rotating flow, and is not discussed in the literature on
875 spontaneous emission.

876 Over the past four decades, several candidate mechanisms have been investigated by
877 which shear instabilities excite gravity waves, in a direct or indirect way, in a linear or
878 nonlinear framework. One essential difficulty here lies in the range of scales involved,
879 from tens of meters for the turbulence initiated from the instability of a shear layer to
880 thousands of kilometers for the baroclinic instability setting the environmental shear and
881 modulating the background stratification.

882 The first investigations of possible mechanisms focused on the linear stability analysis
883 of an atmospheric shear layer. The aim was to determine whether unstable modes exist
884 that comprise a radiating GW above the shear layer or a jet [*Lalas and Einaudi, 1976;*
885 *Lalas et al., 1976; Mastrantonio et al., 1976*]. Although such unstable modes do exist,
886 their growth rates are always considerably smaller than those of KH instability [*Fritts,*
887 *1980*]. The latter always occurs on small scales such that their signature above and below
888 the shear layer is evanescent. *McIntyre and Weissman [1978]* point out a fundamental
889 difficulty for shear instabilities to generate gravity waves: to couple propagating gravity
890 waves above the shear layer, it is necessary that the (real part of the) phase speed, c ,
891 and the horizontal wavenumber, k , to verify the Phase Speed Condition: $U - N/k < c <$
892 $U + N/k$, where U is the wind velocity, N is the Brunt-Väisälä frequency. For large values
893 of k , the interval becomes very narrow and only evanescent responses are found in the
894 layer above the shear.

895 Hence, both the findings of the linear studies and the Phase Speed Condition strongly
 896 suggested that generation from shear instabilities likely involved nonlinear mechanisms.

897 The first nonlinear mechanism to be investigated as a route to larger scales was vortex
 898 pairing [*Davis and Peltier*, 1979]. To obtain significantly larger scales, *Fritts* [1982, 1984]
 899 and *Chimonas and Grant* [1984] described the interaction of two KH modes having nearby
 900 wavenumbers, k and $k + \delta k$. These weak nonlinear interactions produce scales $2\pi/\delta k$,
 901 large enough to radiate gravity waves. This mechanism, called 'envelope radiation', has
 902 been further investigated by *Scinocca and Ford* [2000]. using direct numerical simulations
 903 of the 2D evolution of a region of unstable shear. They focused on the early stages of the
 904 instability (when the two-dimensionality is relevant) and on quantifying the momentum
 905 fluxes associated to envelope radiation. Going beyond the two-dimensional approximation
 906 *Tse et al.* [2003] simulated the three-dimensional turbulence in a forced, unstable jet. In
 907 a subsequent study, *Mahalov et al.* [2007] focused on the emission of gravity waves and
 908 confirmed their capacity to exert a significant drag on the flow emitting them.

909 The end effect of the shear instability will be to mix the fluid over the region where it
 910 developed. This mixing occurs over a short timescale relative to the inertial period, so
 911 the fluid is forced out of balance and will then undergo geostrophic adjustment to recover
 912 a balanced state, and emit inertia-gravity waves in the process *Bühler et al.* [1999]. *Bühler*
 913 *and McIntyre* [1999], who calculated the subsequent propagation of the emitted waves in
 914 a mean wind profile representative of the summer stratosphere. They concluded that the
 915 contribution of this source could not safely be neglected in the global angular momentum
 916 budget.

917 The above studies focused on shear layers in a fluid having constant Brunt-Väisälä
 918 frequency. Another possibility consists in having variations of the stratification leading to
 919 either propagating wave instabilities [*Lott et al.*, 1992; *Sutherland*, 2006] or to a coupling
 920 of the shear instability to upward propagating waves [*Sutherland et al.*, 1994; *Sutherland*
 921 *and Peltier*, 1995]. This may be relevant as the upper-tropospheric jet-stream is indeed
 922 just below the tropopause and its sharp jump in stratification [*Gettelman et al.*, 2011].

923 In summary, theoretical and numerical studies support the notion that gravity waves
 924 generated from shear instabilities need to be considered for middle atmospheric dynamics,
 925 but the complexity of the flows considered has hindered theoretical progress in quantifying
 926 them, while their small scales have made observations difficult.

5. PROPAGATION AND MAINTENANCE

927 The framework of parameterizations and the resulting demand encourages one to think
 928 separately of the gravity wave sources and of their subsequent propagation (in a vertical
 929 column for parameterizations). Now, several mechanisms described above (unbalanced
 930 instabilities and transient generation, sections 4.3 and 4.4) precisely emphasize the key
 931 role played by a varying background wind for the appearance of the waves. In more
 932 complex flows (sections 3 and 6), studies of wave emission emphasize the importance of
 933 propagation effects. This motivates a pause in the review of generation mechanisms to
 934 briefly describe wave ducting, ray-tracing and wave-capture.

5.1. Ducted gravity waves

Ducting of gravity waves between the ground and a layer acting as a partial reflector
 has been modelled by *Lindzen and Tung* [1976]. It occurs when a stable layer is present

near the ground, capped by a layer which efficiently reflects waves (e.g. of low stability, or conditionnally unstable, possibly beneath a critical level). The stable layer needs to be thick enough, and not to contain a critical level. Ducted waves, reflecting off the ground and (partially) at the top of the layer, may travel significant distances in the horizontal, with energy leaking only slowly through the top of the duct. In consequence, such '*almost free*' waves [Lindzen and Tung, 1976] need only a weak forcing to be present, and the geometry and stability of the duct selects some of their characteristics. One characteristic selected by the duct is the phase speed

$$C_D \sim \frac{N_D \mathcal{H}}{\pi \left(\frac{1}{2} + n\right)}, \quad n = 0, 1, 2, \dots \quad (3)$$

where N_D is the Brunt-Väisälä frequency in the duct, and \mathcal{H} its height. The tallest wave ($n = 0$) will be least damped, and is hence of greatest interest. This is a clear example of how the environment in which gravity waves are forced selects certain characteristics of the waves, making it in practice more important to know the duct rather than the details of the forcing.

The relevance of ducting has been shown by numerous case studies focusing on lower-tropospheric waves in the vicinity of surface fronts (e.g. Eom [1975]; Bluestein and Jain [1985]; Parsons and Hobbs [1983]; Uccellini and Koch [1987]; Nicholls et al. [1991]; Powers and Reed [1993]; Zhang and Koch [2000]; Zhang et al. [2003]). Ducted gravity waves are found propagating ahead of cold fronts, and on smaller-scales ahead of gust fronts Knupp [2006], and can play a significant role in triggering convection. The complex interaction between ducted gravity waves and moist convection that maintains and amplifies the mesoscale waves is also referred to "ducted wave-CISK" model [Powers, 1997; Zhang et al., 2001]. Other mechanisms leading to maintenance of gravity waves, e.g. solitary

949 wave dynamics [*Lin and Goff*, 1988], lie beyond the scope of the present paper and will
 950 not be discussed.

5.2. Ray-tracing

951 A common approach to investigate the propagation of gravity waves in complex flows
 952 has been the use of ray-tracing, which we briefly recall below (see *Lighthill* [1978] or
 953 *Bühler* [2009] for a complete discussion, and *Aspden and Vanneste* [2010] for an alternative
 954 derivation). It has typically been used in case studies to identify the origin of observed
 955 waves [*Guest et al.*, 2000; *Hertzog et al.*, 2001], and in idealized simulations to identify
 956 sources and follow emitted waves [*Lin and Zhang*, 2008; *Wang and Zhang*, 2010]. Many
 957 of these studies use the ray-tracing software package developed in *Eckermann and Marks*
 958 [1996, 1997] with various complex background flows.

Consider a wave-packet described by

$$u(\mathbf{x}, t) = A(\mathbf{x}, t) e^{i\theta(\mathbf{x}, t)} \quad (4)$$

for the x -component of the velocity, with A a slowly changing amplitude and θ a fast-varying phase. The local wavevector and frequency are defined by $\mathbf{k}(\mathbf{x}, t) = \nabla\theta$ and $\omega(\mathbf{x}, t) = -\theta_t$, where the subscript is used to denote partial derivation. They vary slowly (as A and the background flow), and are assumed to locally satisfy the dispersion relation:

$$\omega = \Omega(\mathbf{k}(\mathbf{x}, t), \mathbf{x}, t) = \hat{\Omega} + \mathbf{U} \mathbf{k}, \quad (5)$$

959 with ω the absolute frequency and $\hat{\Omega}(\mathbf{k}, \mathbf{x}, t)$ the appropriate dispersion relation.

Now, cross-differentiating the definitions of \mathbf{k} and ω we can obtain $\mathbf{k}_t + \nabla\omega = 0$.

Substitution into (5), using the chain rule and the fact that $\nabla \times \mathbf{k} = 0$ yields:

$$\frac{d\mathbf{x}}{dt} = \frac{\partial\Omega}{\partial\mathbf{k}} \quad \text{and} \quad \frac{d\mathbf{k}}{dt} = -\frac{\partial\Omega}{\partial\mathbf{x}} \quad (6)$$

where

$$\frac{d}{dt} = \frac{\partial}{\partial t} + (\mathbf{U} + \hat{\mathbf{c}}_g) \cdot \nabla \quad \text{and} \quad \hat{\mathbf{c}}_g = \frac{\partial\hat{\Omega}}{\partial\mathbf{k}}.$$

960 An additional equation, generally for the conservation of wave action $A = E/\hat{\omega}$, with
 961 E the energy of the wave, is necessary to follow the evolution of the amplitude of the
 962 wave-packet *Bühler* [2009].

5.3. Wave capture

963 Ray-tracing allows to investigate, with simple considerations, how jet exit region may
 964 have a specific effect on inertia-gravity waves. In studies that have emphasized jet exit
 965 regions as particularly favorable to the occurrence of large-amplitude gravity waves, it
 966 has often been assumed, implicitly, that waves were large because they were generated
 967 there. This overlooks another possibility of interest: that jet exit regions have a particular
 968 significance for gravity waves not only for generation, but also for propagation.

969 Case studies have highlighted a specific region within the jet, where the flow decelerates
 970 and the streamlines are diffluent. The effect of such a background flow on wave packets
 971 propagating through them has been emphasized in theoretical studies as 'wave-capture'
 972 [*Badulin and Shrira*, 1993; *Bühler and McIntyre*, 2005]). The combination of strong
 973 deformation and vertical shear can lead to the contraction of the wave packet to smaller
 974 and smaller scales, until dissipation occurs, without having the intrinsic frequency tending
 975 to either bound of the GW frequency spectrum.

976 Quantifying this effect introduces new possible interactions between waves and mean
 977 flows [Bühler and McIntyre, 2003, 2005], but requires to take into consideration horizontal
 978 variations of the background flow, i.e. to consider propagation in $\mathbf{U}(x, y, z)$. This is in
 979 contrast to the columnar approximation made for parameterizations (where only $\mathbf{U}(z)$ is
 980 considered), and which is encouraged by parallel computing.

For a low-frequency wave packet of sub-synoptic scale, the group velocity is small¹⁰. This warrants an analogy [Bühler, 2009] between the evolution equation for the wavevector and for the evolution of the gradient of a passive, conserved tracer ϕ , respectively:

$$\frac{d\mathbf{k}}{dt} = -(\nabla\mathbf{U}) \cdot \mathbf{k} \quad \text{and} \quad \frac{D\nabla\phi}{Dt} = -(\nabla\mathbf{U}) \cdot \nabla\phi, \quad (7)$$

the two equations differing in their operators on the lhs by

$$\frac{d}{dt} - \frac{D}{Dt} = \hat{\mathbf{c}}_g \cdot \nabla.$$

981 Now, assuming the background flow to be layerwise non-divergent, $\mathbf{U} = (U, V, 0)$, with
 982 $U_x + V_y = 0$, which is relevant as a leading-order description of the background balanced
 983 flow, the evolution of the advected tracer gradient is governed by the sign of

$$D = -U_x V_y + V_x U_y \quad (8)$$

$$= \frac{1}{4} ((U_x - V_y)^2 + (V_x + U_y)^2 - (V_x - U_y)^2). \quad (9)$$

984 The first two terms on the rhs of (9) constitute the strain [Batchelor, 1967], and the last is
 985 the vertical component of the relative vorticity. D is also referred to as the Okubo-Weiss
 986 parameter and extensively discussed in studies of tracer advection (e.g. Lapeyre et al.
 987 [1999] and refs. therein).

If the wavepacket remains where the strain dominates ($D > 0$), the wavenumber experiences exponential growth (see Bühler [2009], section 14.3). As a simple example,

consider a pure deformation flow with extension along the y -axis, with vertical shear: $U = -\alpha x + \beta z$ and $V = \alpha y + \gamma z$. Equation 7 then yields $k = k_0 e^{\alpha t}$, $l = l_0 e^{-\alpha t}$ and $m \rightarrow -\alpha^{-1} \beta k(t)$ as $t \rightarrow +\infty$. Asymptotically, the wavevector will tend to

$$(k, l, m) \rightarrow k_0 e^{\alpha t} \left(1, 0, \frac{U_z}{U_x} \right)$$

988 for $t \rightarrow \infty$, and with k_0 the initial value of wavenumber k . More generally, the above
 989 considers the action of only one region of strain on a wavepacket. As a packet moves
 990 within the flow (by advection and by its own propagation), it may encounter different
 991 regions of strain, and *Aspden and Vanneste* [2010] show that this will lead to growth of
 992 the wavenumber, as for tracer gradients [*Haynes and Anglade*, 1997].

993 We emphasize two implications: first, in jet exit regions, deformation and shear are
 994 large. For wave packets that have a long enough residence time in such regions, propaga-
 995 tion effects will favor certain orientation and intrinsic frequency, with little sensitivity to
 996 the initial condition, and contraction of the wavelength. Second, this is only an asymp-
 997 totic result, neglecting spatial variations of the background shear and strain. Its efficiency
 998 will depend on the residence time of the wave packet in the jet exit region. This effect
 999 has been named 'wave-capture', because the asymptotic calculation suggests contraction
 1000 of the wavelength down to dissipation. In practice, it may be that capture is only par-
 1001 tially realized, but this effect will nonetheless constrain wave characteristics, and the term
 1002 wave-capture will be used to designate this influence.

6. GENERATION MECHANISMS: LABORATORY AND MODELLING EXPERIMENTS

1003 There is a certain discrepancy between the simplicity necessary for analytical stud-
1004 ies, e.g. plane-parallel unbounded shears (sections 4.3-4.4), and the complex, three-
1005 dimensional flow patterns highlighted in observations, e.g. jet exit regions (section 3).
1006 Laboratory experiments (section 6.1) and idealized simulations (sections 6.2-6.5) have
1007 provided a realm for exploring spontaneous emission in flows of intermediate complex-
1008 ity, bridging the two, and establishing a convincing sketch of the generation mechanism
1009 involved near jet exit regions.

6.1. Laboratory experiments

1010 Laboratory experiments provide valuable examples of *real* flows, in which a fundamental
1011 dynamical mechanism may be identified, and to some extent isolated. Understanding
1012 these experiments can greatly enhance our understanding of the atmosphere and ocean,
1013 provided the mechanisms at play are the same.

1014 Several experiments have been reported as exhibiting spontaneous generation of gravity
1015 waves in stratified fluids, mainly in a rotating annulus, either thermally or shear-driven,
1016 but also in other configurations.

1017 A classical laboratory experiment of baroclinic instability has focused on a shear-driven
1018 fluid in a rotating annulus [*Hart*, 1972]. For such a configuration, with two immiscible
1019 fluid layers having each an aspect ratio of 2 (height / width), *Lovegrove et al.* [2000]
1020 and *Williams et al.* [2005] reported the appearance of inertia-gravity waves in a flow
1021 dominated by baroclinic instability. The flow is investigated from measurements of the
1022 interface height. In a regime dominated by large scale baroclinic waves (wavenumber
1023 2), small-scale features (wavenumber between 30 and 40, *Williams et al.* [2008]) occur
1024 which are interpreted as inertia-gravity waves. The amplitude is estimated, in the range

1025 $0.05 < Ro < 0.14$, to vary linearly with Rossby number [*Williams et al.*, 2008]. The
 1026 generation mechanism was argued by *Williams et al.* [2005] to be Lighthill radiation,
 1027 because the 'forcing terms' (as in equation (2) [*Ford et al.*, 2000] and assuming shallow
 1028 water) are colocated with the gravity waves. However, given that the flow regime ($Ro \ll 1$,
 1029 and not shallow water) and the scale separation (small-scale waves) differ so completely
 1030 from those for Lighthill radiation, and that the amplitude varies linearly although the
 1031 hypothesized forcing is quadratic, one may say that the generation mechanism remains to
 1032 be explained.

1033 A similar experiment has recently been carried out by *Scolan et al.* [2011], but with
 1034 a salt stratification including a sharp transition rather than immiscible fluids, and with
 1035 an aspect ratio (~ 0.2) compatible with a shallow water interpretation. Interpretation is
 1036 supported by the complete stability analysis for two-layer shallow water sheared flows in an
 1037 annulus obtained by *Gula et al.* [2009b], which includes an *unbalanced instability* (Rossby-
 1038 Kelvin, see Section 4.3). *Scolan et al.* [2011] identify this unbalanced instability, for the
 1039 first time in laboratory experiments. They also find that small-scale perturbations are
 1040 present in many regimes of parameters. These small-scale features are argued in many
 1041 cases to result from Hölmböe instability (e.g. *Lawrence et al.* [1991]). This instability
 1042 occurs when a sharp density interface is colocated with a thicker shear layer, and is hence
 1043 particularly relevant for the experiments with immiscible fluids of *Williams et al.* [2005].

1044 Thermally-driven annulus experiments have also reported small-scale features [*Read*,
 1045 1992] which could be gravity waves. Numerical simulations have proved necessary to
 1046 confirm this [*Jacoby et al.*, 2011], and have further identified an instability of the lateral
 1047 boundary layer as the generation mechanism. Its location in azimuth remains unexplained,

1048 but is likely tied to be the separation of the large-scale geostrophic jet from the inner
1049 boundary. This example, and the reinterpretation of the 'waves' investigated by *Williams*
1050 *et al.* [2005] as Hölmboë instability *Scolan et al.* [2011], emphasizes the importance of
1051 boundary or interfacial layers in such laboratory experiments, making it more difficult to
1052 relate these results to atmospheric or oceanic flows.

1053 Another unbalanced instability has been identified in laboratory experiments: *Riedinger*
1054 *et al.* [2010a] have analyzed the radiative instabilities of axisymmetric, columnar vor-
1055 tices in non-rotating, stratified fluid. The radiative instability of the flow around a
1056 rotating cylinder has been described theoretically and very clearly displayed in experi-
1057 ments *Riedinger et al.* [2011]. The robust agreement between theory and experiments
1058 in this somewhat contrived configuration makes the (difficult) experimental identification
1059 [*Riedinger et al.*, 2010b] of the radiative instability of a columnar vortex all the more
1060 convincing. Remarkably, this is the first laboratory evidence of an unbalanced, radiative
1061 instability.

1062 Spontaneous emission was also investigated during the collision and rearrangement of
1063 two dipoles in the interior of a two-layer, non-rotating fluid [*Afanasyev*, 2003]. The
1064 experiments confirmed the radiation of interfacial gravity waves, occurring when fluid
1065 parcels underwent strong accelerations, such that the spatial scale and the Lagrangian
1066 timescale matched the dispersion relation.

1067 Perhaps the clearest experimental evidence of spontaneous emission was provided by
1068 study of an unstable coastal jet in a two-layer fluid [*Afanasyev et al.*, 2008]. A clever
1069 visualisation technique (Altimetric Imaging Velocimetry, *Rhines et al.* [2006]) allowed
1070 to detect and quantify precisely the waves emitted, and to describe with a very high

1071 resolution the vortical flow emitting the waves. A notable difference relative to other
 1072 studies on spontaneous generation is that the emitted waves are inertial waves in the
 1073 unstratified lower layer, hence not constrained by $\hat{\omega} \geq f$. Waves were radiated away
 1074 from the meanders of the baroclinic instability when the deformation radius was short
 1075 enough that the characteristics of the meanders matched the dispersion relationship for
 1076 the inertial waves, see Fig. 10. In experiments with larger deformation radius, single
 1077 events of emission could be isolated, emphasizing regions of strong curvatures and large
 1078 accelerations. Emitted waves represented only a small fraction, about 0.5%, of the total
 1079 energy of the flow.

6.2. Early Simulations

1080 The numerical study of geostrophic adjustment of a jet streak by *Tuyl and Young*
 1081 [1982] deserves to be highlighted because they identified several essential issues which,
 1082 although simple, have sometimes been overlooked thereafter. They simulated, in a two-
 1083 layer model, the adjustment of perturbations added to a jet streak and emphasize how the
 1084 background flow crucially changes the adjustment and the wave dynamics. They give
 1085 three reasons why traditional approaches (more specifically, the normal mode techniques of
 1086 *Machenhauer* [1977]; *Baer and Tribbia* [1977]) fail to separate gravity waves and balanced
 1087 motions in the vicinity of jet streaks: 1) *the gravity-inertia modes are eigenfunctions for*
 1088 *a base state of rest, rather than a sheared, time-dependent jet; 2) the methods may not*
 1089 *work for strong accelerations (Rossby number of order unity (...)); and 3) the frequency*
 1090 *separation has been based upon Eulerian (fixed frame) frequencies, rather than Lagrangian*
 1091 *(Doppler shifted) ones,'* (*Tuyl and Young* [1982], p 2039). Indeed, points 1 and 3 underlie

1092 the spontaneous generation of gravity waves in a shear (sections 4.3 and 4.4), and point
 1093 2 is an ingredient of Lighthill radiation.

1094 The simulations of *Tuyl and Young* [1982] may be regarded as an early prototypes of
 1095 the recent dipole experiments (section 6.5). With anticipation, they suggest that gravity
 1096 wave modes near jet streaks, although usually discarded as meteorological noise, *'may*
 1097 *eventually show their more persistent members to be a complex part of the jet streak signal*
 1098 (p 2038).

6.3. Two-dimensional frontogenesis

1099 Early numerical experiments of spontaneous generation described two-dimensional fron-
 1100 togenesis. This is understandable for two reasons: physically, fronts are regions of the flow
 1101 where short scales are produced (collapse to a near-discontinuity in a finite time [*Hoskins*
 1102 *and Bretherton*, 1972] and large velocities are encountered. Practically, major features of
 1103 frontogenesis can be understood in a two-dimensional framework [*Hoskins*, 1982], which
 1104 greatly simplifies the problem and made it possible to attain higher resolutions. Al-
 1105 though frontogenesis has sometimes been considered as an *adjustment* (e.g. *Kalashnik*
 1106 [1998, 2000]), it is a specific process, central to mid-latitude dynamics, and deserves its
 1107 own discussion, distinct from that of geostrophic adjustment (section 4.1).

1108 A first study of gravity waves emitted by fronts was carried out with a mostly analytical
 1109 approach by *Ley and Peltier* [1978]. They calculated the far-field gravity wave response
 1110 to a frontogenesis event modeled by SG, assuming the background to be at rest when
 1111 calculating the gravity wave response. Subsequent studies explicitly simulated the frontal
 1112 collapse with different numerical methods [*Gall et al.*, 1987, 1988], including a Lagrangian
 1113 description [*Garner*, 1989], with contradictory results regarding gravity waves. *Snyder*

1114 *et al.* [1993] showed that some of the excited waves were spurious, due to poor initialization
 1115 and an inconsistency between the aspect ratio of the grid ($\Delta z/\Delta x$) and of the frontal slope
 1116 yielding spurious waves [*Lindzen and Fox-Rabinowitz*, 1989].

1117 *Snyder et al.* [1993] simulated both inviscid frontogenesis prior to frontal collapse, and
 1118 postcollapse frontogenesis with horizontal diffusion, with frontogenesis forced by either
 1119 deformation or shear. They used a nonhydrostatic model and their domain was bounded in
 1120 the vertical by a flat surface and a rigid lid. Significant gravity waves, i.e. dominating other
 1121 corrections to semi-geostrophy, are emitted when the frontogenesis is sufficiently intense,
 1122 and are most prominent in the postcollapse solutions, above the surface front. Emission
 1123 occurred when the advective time-scale, which decreases as frontogenesis proceeds and the
 1124 cross-frontal scale shrinks, became comparable to or shorter than the inertial period. This
 1125 emission was explained as the linear response, in the frontogenetical background flow, to
 1126 the cross-front accelerations neglected by semi-geostrophy.

1127 More realistic simulations focused on gravity waves generation were carried out by *Grif-*
 1128 *fiths and Reeder* [1996], who considered a domain including a stratosphere. Three cases
 1129 of deformation frontogenesis were simulated: without, with negative and with positive
 1130 vertical shear in the transverse direction (this transverse shear makes the large-scale fron-
 1131 togenetical forcing time-dependent). Emission of large-scale, low-frequency waves from
 1132 the upper-level front and propagating up into the stratosphere was found in all three
 1133 cases. Their comparison revealed that a determining factor for the amplitude of the emit-
 1134 ted waves was the rapidity of the frontogenesis rather than its intensity (estimated by
 1135 the maximum cyclonic vorticity). In other words, the emission is limited by the fact that
 1136 the forcing (the transverse, frontogenetic circulation) poorly 'projects' on gravity wave

1137 modes. *Reeder and Griffiths* [1996] used ray-tracing to confirm the origin of the waves
 1138 from the upper-level front, and its initial near-inertial frequency ($\hat{\omega} \sim 1.3 f$). The emission
 1139 was analyzed, with reference to Lighthill radiation *Ford* [1994c], as the linear response,
 1140 *in a background flow consisting of the imposed deformation and transverse shear*, to the
 1141 nonlinear terms from the frontal circulation. Linear forced simulations reproduced satis-
 1142 factorily the emitted waves away from the fronts, whether using the full simulation or a
 1143 balanced approximation to estimate these forcing terms. Crucially, the linear simulations
 1144 include the background deformation and time-dependent shear, leading respectively to
 1145 contracting wavelengths (from 1000 km to 500 km) and increasing vertical wavelength
 1146 (from 3 km to 10 km). Inclusion of this background flow profoundly modifies the problem
 1147 relative to Lighthill radiation (see section 6.5).

6.4. Idealized baroclinic life cycles

1148 Idealized life cycles of baroclinic instability provide more realistic flows to investigate
 1149 spontaneous emission, but requires significant computational resources as an additional
 1150 spatial dimension is needed. *O'Sullivan and Dunkerton* [1995] simulated a baroclinic
 1151 life cycle on the sphere (wavenumber 6, following *Simmons and Hoskins* [1978]) with a
 1152 spectral truncation at wavenumber 126 (T126, approximately equivalent to a horizontal
 1153 grid spacing of 1°). Inertia-gravity waves with intrinsic frequencies between f and $2f$ arose
 1154 during the nonlinear stage of the development of the baroclinic wave, principally in the jet-
 1155 stream exit region in the upper troposphere (see Figure 12). Surface fronts were shown not
 1156 to be the source of these waves. They subsequently propagated horizontally within the jet,
 1157 but only few IGWs penetrated the lower stratosphere. *O'Sullivan and Dunkerton* [1995]
 1158 showed maps of the Lagrangian Rossby number with a large-scale maximum roughly

1159 coincident with the waves and put forward geostrophic adjustment as the generation
 1160 mechanism.

1161 The simulations and interpretations of *O'Sullivan and Dunkerton* [1995] have become a
 1162 milestone for several reasons: they explicitly showed IGWs generated by jets, with more
 1163 realism than 2D frontogenesis simulations, allowing essential features emphasized from
 1164 observations (low frequency, jet exit region) to be reproduced. As a consequence, their
 1165 interpretation in terms of geostrophic adjustment, and the confirmation of the relevance of
 1166 the Lagrangian Rossby number as a diagnostic, have guided interpretations in subsequent
 1167 studies, in particular for observations (e.g. *Pavelin et al.* [2001]; *Plougonven et al.* [2003]).

1168 As shown by sensitivity tests, the simulations of *O'Sullivan and Dunkerton* [1995] did
 1169 not converge numerically (see their Fig. 9), which was somewhat controversial at the
 1170 time¹¹. In fact, a contemporaneous study by *Bush et al.* [1995] used very similar idealized
 1171 baroclinic life cycles (with $\Delta x \sim 60km$) to analyze the degree of balance of the flow.
 1172 Gravity waves were found to be more intense near the cold fronts than in the upper-
 1173 troposphere, but the analysis of these frontal waves strongly suggested that they were a
 1174 numerical artifact, again due to the shallow slope of the front near the surface, shallower
 1175 than $\Delta z/\Delta x$. These numerical issues raise two questions: **1)** at what resolution would the
 1176 gravity waves converge, and what small-scale gravity waves would then be obtained? **2)**
 1177 how should one interpret such gravity waves from simulations that have not converged?

1178 Regarding the first question, the resolution used by *O'Sullivan and Dunkerton* [1995]
 1179 only allowed subsynoptic scale inertigravity waves with horizontal wavelengths of 600-
 1180 1000 km to be described. However, as reviewed in section 3, it is mesoscale gravity waves
 1181 with horizontal wavelengths of 50-500 km that are found to be prevalent in the vicinity

1182 of the unbalanced upper-level jet streaks. This has been demonstrated repeatedly from
1183 observational studies of gravity waves (e.g., *Uccellini and Koch* [1987]; *Bosart et al.* [1998];
1184 *Thomas et al.* [1999]) and the corresponding numerical investigations (e.g., *Powers and*
1185 *Reed* [1993]; *Zhang and Koch* [2000]; *Zhang et al.* [2001]). These mesoscale waves may
1186 have a greater impact on the transport of momentum than the subsynoptic waves [*Fritts*
1187 *and Nastrom*, 1992].

1188 Consequently, *Zhang* [2004] performed multiply nested mesoscale numerical simulations
1189 with horizontal resolution up to 3.3 km to study the generation of mesoscale gravity waves
1190 during the life cycle of idealized baroclinic jetfront systems. Long-lived vertically propa-
1191 gating mesoscale gravity waves with horizontal wavelengths ~ 100 -200 km are simulated
1192 originating from the exit region of the upper-tropospheric jet streak, in a manner con-
1193 sistent with past observational studies. The residual of the nonlinear balance equation
1194 is found to be a useful index in diagnosing flow imbalance and predicting the location
1195 of wave generation. *Zhang* [2004] proposed the term *balanced adjustment* to describe the
1196 continuous radiation of waves within the developing baroclinic wave. A framework to de-
1197 scribe this emission was proposed by *Plougonven and Zhang* [2007] through scale analysis
1198 and analytical derivation of a wave equation linearized on the balanced background flow
1199 that is forced by synoptic-scale flow imbalance. This was implemented and expanded to
1200 explain gravity waves emitted in dipoles [*Wang and Zhang*, 2010], and has recently been
1201 used to explain at least some of the jet-exit region gravity waves found in baroclinic life
1202 cycles *Wang* [2008].

1203 To further investigate the sources and propagation of gravity waves in the baroclinic
1204 jet-front systems, *Lin and Zhang* [2008] carried out ray-tracing from the four groups of

1205 waves they identified in the lower stratosphere: a northward-propagating short-scale wave
 1206 packet (horizontal wavelength $\lambda_H \sim 150$ km), and a northeastward-propagating medium-
 1207 scale wave packet ($\lambda_H \sim 350$ km) in the exit region of the upper-tropospheric jet, a third
 1208 packet in the deep trough region above (and nearly perpendicular to) the jet ($\lambda_H \sim 100$ -
 1209 150 km), and a fourth group far to the south of the jet right above the surface cold front
 1210 ($\lambda_H \sim 100$ -150 km). The ray-tracing analysis suggests that the medium-scale gravity
 1211 waves originate from the upper-tropospheric jet-front system where there is maximum
 1212 imbalance, though contributions from the surface fronts cannot be completely ruled out.
 1213 The shorter scale wave packets, on the other hand, possibly originate from the surface
 1214 front: this is a possibility for the northward-propagating gravity waves in the jet-exit
 1215 region, and a certainty for the other two packets. Ray-tracing analysis also reveals a very
 1216 strong influence of the spatial and temporal variability of the complex background flow
 1217 on the characteristics of gravity waves as they propagate.

1218 *Wang and Zhang* [2007] investigated the sensitivity of mesoscale gravity waves to the
 1219 baroclinicity of the background jet-front systems by simulating different life cycles of
 1220 baroclinic waves with a high-resolution mesoscale model. In all experiments, vertically
 1221 propagating mesoscale gravity waves are found in the exit region of upper-tropospheric
 1222 jet streaks. The intrinsic frequencies of these gravity waves tend to increase with the
 1223 growth rate of the baroclinic waves. They further found that the growth rate of flow
 1224 imbalance also correlates well to the growth rate of baroclinic waves and thus correlates
 1225 to the frequency of gravity waves.

1226 Regarding the second question, *Plougonven and Snyder* [2005] have shown that simu-
 1227 lations that did not converge numerically nevertheless could carry relevant information

1228 regarding the location, the horizontal orientation and the intrinsic frequency $\tilde{\omega}$ of the
1229 waves in jet exit regions. The reason is that those characteristics are largely determined
1230 by propagation through the large-scale flow (see section 5.3), which is well described.
1231 The large-scale strain causes a wavepacket's wavenumber to increase exponentially along
1232 a ray, so that details of the waves will always be sensitive to resolution. Nonetheless,
1233 even at low resolution the location, orientation and $\tilde{\omega}$ of waves may be relevant because
1234 they are constrained by the large-scale deformation and shear. Evidence for this effect
1235 also comes from comparison of simulated waves with observations [*Plougonven and Teit-*
1236 *elbaum, 2003*] and from the remarkable insensitivity to resolution of $\tilde{\omega}$ for waves in jet
1237 exit regions [*Plougonven and Snyder, 2007*].

1238 Now, the above studies focused each on one idealized life cycle, emphasizing gravity
1239 waves emanating from the upper-level jet¹². One issue then concerns the relation of these
1240 three-dimensional simulations to the evidence from 2D frontogenesis simulations and from
1241 some observations (*Eckermann and Vincent [1993]*; *Fritts and Nastrom [1992]*) indicating
1242 wave generation from the surface fronts.

1243 In order to test the sensitivity of the wave generation to the background flow, *Plougonven*
1244 *and Snyder [2007]* ran two very different baroclinic life cycles, following the paradigm of
1245 *Thorncroft et al. [1993]* who highlighted two types of nonlinear Rossby wave breaking,
1246 cyclonic and anti-cyclonic. In the cyclonic run, gravity waves were found in the jet exit
1247 region, clearly emitted by the jet (both an upward and a downward wavepacket are found),
1248 have sub-synoptic scale, similar to those described by *O'Sullivan and Dunkerton [1995]*.
1249 In the anticyclonic run, the most conspicuous waves are found ahead of the surface cold
1250 front (see Figure 14), reminiscent of those found in 2D frontogenesis studies, and have a

1251 different sensitivity to resolution: as resolution increases, their vertical wavelength remains
 1252 unchanged while the horizontal one decreases, yielding higher frequencies (up to $3f$): these
 1253 waves are *not* undergoing wave capture. Their generation seems tied to an obstacle effect
 1254 (strong surface winds impinging on the cold front), as in the case study of *Ralph et al.*
 1255 [1999].

1256 Baroclinic life cycles in a very different configuration (triply periodic domain, initial jet
 1257 specified by strong interior PV anomalies) have been carried out by *Viúdez and Dritschel*
 1258 [2006] to study spontaneous emission with a sophisticated code and inversion for the
 1259 balanced flow. Waves with intrinsic frequencies close to inertial ($N/f \sim m/k$) were
 1260 produced in very localized bursts where the flow has strong curvature, on the anticyclonic
 1261 side of the jet. One packet remains trapped within the vortices, while another propagates
 1262 significantly outward.

1263 *Waite and Snyder* [2009] carried out baroclinic life cycle experiments at high resolution
 1264 ($\Delta x = 10\text{km}$, $\Delta z = 60\text{m}$), which revealed three types of waves spontaneously generated
 1265 (a long packet tied to the cold front [*Snyder et al.*, 1993], a compact one east of the
 1266 ridge, turning into the cyclone, east of the ridge [*Zhang*, 2004], and a long packet from
 1267 the jet exit region in the ridge down into the trough [*Plougonven and Snyder*, 2007]). At
 1268 later times, these localized packets give way to more disordered wave signatures filling the
 1269 whole region of the baroclinic jet and vortices. *Waite and Snyder* [2009] investigated the
 1270 contribution to the mesoscale energy spectrum of the spontaneously emitted IGW (both
 1271 by a direct cascade and by vertical propagation), and showed that they could yield a $-5/3$
 1272 spectrum, but only in the lower stratosphere and with too low amplitude. This suggests
 1273 that other contributions to the mesoscale spectrum (convection, topography) are crucial.

6.5. Dipoles

Both observations and idealized baroclinic life cycles have stressed jet exit regions as favored sites for the appearance of conspicuous inertia-gravity waves. Now, a simple model of jet exit regions is provided by dipoles (e.g. *Cunningham and Keyser* [2000]). Numerical simulations of dipoles have been carried out by several different groups, using different configurations and very different models: *Snyder et al.* [2007] simulated a surface dipole, from an initial dipole that is an exact solution in the quasi-geostrophic approximation [*Muraki and Snyder*, 2007]. *Viudez* [2007, 2008] simulated a dipole in the interior of the fluid with constant stratification, from an initial condition with PV anomalies of opposite magnitudes but slightly different structures. Their model uses potential vorticity (PV) as one of the prognostic variables [*Dritschel and Viúdez*, 2003] and evolves it using contour advection to ensure a good conservation. They invert the initial PV distribution with a unique method which iteratively finds an optimal balanced state which minimizes the unbalance in the full dynamics [*Viúdez and Dritschel*, 2003]. *Wang et al.* [2009] simulated both surface and interior dipoles. Antisymmetric initial PV anomalies are inverted using the Nonlinear Balance Equations [*Davis and Emanuel*, 1991], producing asymmetric dipoles. *Snyder et al.* [2007] and *Wang et al.* [2009] use different models based on finite differences, and *Viudez* [2008] has checked his results with a pseudospectral code. All these simulations were carried out on the f -plane, with domains that are doubly periodic in the horizontal, or bounded [*Wang et al.*, 2009].

In all cases, the dipoles proved to be robust structures: after an initial adjustment, the dipoles propagated steadily and for long periods (tens of days), along trajectories that curve with a radius of curvature very large relative to the dipole size. Hence they have

1296 the great advantage of providing a background flow that retains a jet exit region, but that
 1297 is nearly stationary in the appropriate frame of reference.

1298 A robust phenomenology emerged from these simulations: a gravity wave packet was
 1299 systematically found in the front of the dipole, in the jet exit region, with phase lines
 1300 rather normal to the jet and wavelengths contracting to smaller scales in the front of the
 1301 wavepacket (see Fig. 15). The phase lines extend into the anticyclone. The intrinsic
 1302 frequencies are close to the inertial frequency ($f < \hat{\omega} < 2f$). In these diverse simulations,
 1303 the presence, orientation and relation to the background flow is strikingly robust, making
 1304 these Jet Exit Region Emitted (JEREmi) waves a paradigm to understand similar wave
 1305 packets found in baroclinic life cycles. Some minor aspects differ between the simulations,
 1306 such as the importance of the bias toward the anticyclone and the rather weak amplitude
 1307 of the simulated waves.

1308 The origin of the waves has been carefully examined and discussed, demonstrating un-
 1309 ambiguously that they are not remnants of the adjustment of the initial condition but
 1310 truly result from spontaneous generation [*Snyder et al.*, 2007; *Wang and Zhang*, 2010].
 1311 Vertical cross-sections through the dipole axis clearly suggest that the waves originate in
 1312 the jet core, where fluid parcels undergo significant acceleration then deceleration, accom-
 1313 panied with vertical displacements *McIntyre* [2009]. The waves appear as a conspicuous
 1314 component of the flow downstream, in the jet exit region, and are there consistent with
 1315 wave-capture [*Bühler and McIntyre*, 2005]. This influence, due to the background defor-
 1316 mation and shear, can be seen graphically from the tendency of the phase lines to align
 1317 with the isolines of along-jet velocity (Fig. 15, or Figure 4.b of *Wang et al.* [2009]), and
 1318 was verified using ray-tracing by *Wang et al.* [2010]. This was highlighted in other simu-

1319 lations and discussed independently, but yielding the same conclusion *Viudez* [2008]. The
 1320 waves were found not to be detectable when the Rossby number was too small (less than
 1321 0.15 *Snyder et al.* [2007] or less than 0.05 *Wang et al.* [2009]), and showed an algebraic
 1322 dependence above that (exponents between 2 and 6). The dependence on the Rossby
 1323 number however is very sensitive to resolution *Wang et al.* [2009], and is obtained only for
 1324 a narrow range of Rossby numbers (e.g. 0.15-0.30 in *Snyder et al.* [2007]). Hence it could
 1325 not be conclusive to compare this dependence with theoretical predictions, in particular
 1326 a non-algebraic one such as an exponential dependence in Rossby number.

1327 The waves have been explained as a linear response to a forcing which is akin to the
 1328 imbalance produced by the balanced flow. The idea of such a linearization goes back, in the
 1329 context of frontogenesis, at least to *Ley and Peltier* [1978], *Snyder et al.* [1993] and *Reeder*
 1330 *and Griffiths* [1996], and more generally to *Lighthill* [1952]. Building on the (spontaneous)
 1331 balance adjustment hypothesis proposed in *Zhang* [2004], a general framework for such
 1332 linearization has been discussed by *Plougonven and Zhang* [2007], emphasizing the need to
 1333 linearize around the large-scale background flow. The basic assumption is that the waves
 1334 are small enough to be described by linearized dynamics around a balanced approximation
 1335 of the flow.

As a crude sketch of this linearization, we consider the equation for the velocity in the x direction, u , in the Boussinesq approximation on the f -plane (e.g. *McWilliams and Gent* [1980]):

$$\frac{\partial u}{\partial t} + \mathbf{u} \nabla u - fv + \frac{\partial \Phi}{\partial x} = 0, \quad (10)$$

where f is the Coriolis parameter and Φ is geopotential. Now, the flow can always be decomposed into two components $u = \bar{u} + u'$, where \bar{u} is a balanced approximation of the

flow (or its large-scale part) and u' the residual, including gravity waves and higher-order balanced corrections. For Rossby numbers smaller than one but finite, it is expected that the background flow will well be approximated by a balanced relation ($|u'| \ll |\bar{u}|$), but that the emission of gravity waves will dominate u' . Injecting the decomposition into (10), three types of terms appear: terms involving only the balanced flow are moved to the right hand side (*rhs*), terms linear in the perturbations are kept on the *lhs*, and terms that are quadratic in perturbations are neglected. This yields forced equations for the perturbations u' , linearized on the background balanced flow $\bar{\mathbf{u}}$:

$$\frac{\partial u'}{\partial t} + \bar{\mathbf{u}}\nabla u' + \mathbf{u}'\nabla\bar{u} - fv' + \frac{\partial\Phi'}{\partial x} = \mathcal{F}_u, \tag{11}$$

where

$$\mathcal{F}_u = \frac{\partial\bar{u}}{\partial t} + \bar{\mathbf{u}}\nabla\bar{u} - f\bar{v} + \frac{\partial\bar{\Phi}}{\partial x} \tag{12}$$

1336 is the residual tendency, i.e. the residual when the balanced solution is injected into the
 1337 primitive equations. If used in a systematic asymptotic approach with $Ro \ll 1$, the above
 1338 approach yields no emission [*Reznik et al.*, 2001; *Vanneste*, 2008; *Plougonven et al.*, 2009].
 1339 More precisely, there may be an emission that is exponentially small in Ro , and hence
 1340 not described by the above heuristic approach. Emission appears at finite Ro , when the
 1341 advection on the *lhs* and the forcing on the *rhs* are both strong enough.

1342 Several studies have investigated variations on this approach, using different sets of
 1343 equations (for momentum and potential temperature in *Snyder et al.* [2009], for horizontal
 1344 divergence and vorticity in *Wang et al.* [2010]; *Wang and Zhang* [2010]). They consistently
 1345 show that the structure (location, orientation, intrinsic frequency) of the wave-packet is
 1346 mainly determined by the background flow, (i.e. the *lhs* operator), not by the forcing:

1347 the latter is large-scale and bears no resemblance with the produced waves [*Snyder et al.*,
 1348 2009]. In other words, and as confirmed using ray-tracing, *'the effects of propagation*
 1349 *dominate over the source'* [*Wang et al.*, 2010]. The latter does matter to determine the
 1350 amplitude of the emitted waves, i.e. they do not arise merely as an instability but require
 1351 a forcing to be present in (11). Given the importance of propagation effects, diagnostics
 1352 of the large-scale flow such as the Okubo-Weiss parameter (e.g. *Lapeyre et al.* [1999]),
 1353 which appears in the description of wave capture, are likely as important as traditionnaly
 1354 used diagnostics of imbalance.

1355 In summary, different dipole experiments have shown the robustness of Jet Exit Region
 1356 Emission (JEREmission). The crucial ingredients are strong velocities in the jet core,
 1357 combined with along-jet variations: the first leads to strong advection ($\bar{\mathbf{u}}\nabla u'$ in equation
 1358 (11)), the second produces a forcing, e.g. as in equation (12) (a zonally symmetric jet
 1359 does not by itself produce waves). The advection allows this forcing to project onto fast
 1360 Largangian timescales (shorter than $1/f$).

1361 This linearization is in part inspired by Ford's work on Lighthill radiation [*Reeder and*
 1362 *Griffiths*, 1996; *Plougonven and Zhang*, 2007]. Essential differences need to be emphasized
 1363 to avoid confusion: in the case of Lighthill radiation, the scale separation between the
 1364 vortical flow and the GW implies that the *lhs* operator is that for GW on a background
 1365 of fluid at rest [*Plougonven and Zhang*, 2007; *Plougonven et al.*, 2009]. One important
 1366 consequence is that the quadrupolar form of the forcing partly determines the weakness
 1367 emission [*Ford et al.*, 2000, 2002]. For JEREmission waves, the scale separation is the opposite,
 1368 and advection plays a crucial role to allow projection of the forcing onto fast intrinsic
 1369 timescales. The higher-order derivatives of the large-scale forcing enhances the small-

1370 scale part of the forcing, and hence this projection. Fundamental conclusions concerning
 1371 regarding Lighthill radiation [*Ford et al.*, 2002] no longer hold, and this motivates a sharp
 1372 distinction between the generation mechanism at play in stratified dipoles or baroclinic
 1373 life cycles and Lighthill radiation [*Zhang*, 2004; *McIntyre*, 2009].

7. IMPACTS AND PARAMETERIZATIONS

1374 A major motivation driving recent research on atmospheric gravity waves is their role
 1375 in transferring momentum towards the middle atmosphere (e.g. *Fritts and Alexander*
 1376 [2003]). Constraints from observations and simulations, along with a better physical un-
 1377 derstanding, are needed to improve parameterizations in Atmospheric General Circulation
 1378 Models (GCMs) (section 7.1). Yet gravity waves emitted from atmospheric jets and fronts
 1379 also matter for other impacts, such as their local contributions to mixing and turbulence
 1380 (section 7.2), and also to temperature-dependent phenomena (section 7.3).

1381 While all the studies described above focus on the atmosphere, the same dynamical
 1382 mechanisms that have been discussed in sections 4 and 6 are also active in the ocean, as
 1383 discussed in section 7.4.

7.1. Momentum fluxes and parameterizations

1384 Gravity waves are crucial to the general circulation of the stratosphere and mesosphere
 1385 because they transfer momentum upward [*Andrews et al.*, 1987]. Atmospheric General
 1386 Circulation Models (GCMs) typically include two parameterizations, one for orographic
 1387 gravity waves and one for non-orographic gravity waves. The latter generally have an arbi-
 1388 trarily fixed source at a given level, tuned in order to produce a reasonable stratospheric
 1389 circulation [*Kim et al.*, 2003]. While parameterizations of convective sources of gravity

1390 waves have been elaborated and implemented in the last decade [*Beres et al.*, 2004, 2005;
1391 *Song and Chun*, 2005], parameterizations of waves produced by jets and fronts remain
1392 exceptionnal: *Rind et al.* [1988] included waves generated by wind shear at the level of
1393 the tropospheric jet stream. *Charron and Manzini* [2002] and *Richter et al.* [2010] have
1394 used the frontogenesis function [*Miller*, 1948; *Hoskins*, 1982] in the mid-troposphere (600
1395 hPa) as a diagnostic to identify active source regions. *Richter et al.* [2010] prescribed the
1396 emitted waves with a Gaussian phase speed spectrum centred on the local wind, and kept
1397 the amplitudes as a tunable parameter. Improvements included a reduction of the cold
1398 pole bias and a better variability of the stratospheric circulation (frequency of Strato-
1399 spheric Sudden Warmings), although they are not solely due to the changes in the GW
1400 parameterization (the addition of Turbulent Mountain Stress also contributed).

1401 Implementing successfully a new parameterization with variable source, without degrad-
1402 ing other features of the GCM's circulation, already is a significant achievement. Yet, the
1403 parameterizations described above remain heuristic, and progress is needed to include
1404 more physical understanding. Pathways to improve parameterizations of jets and fronts
1405 as sources include the systematic use of observational datasets (e.g. *Gong et al.* [2008]
1406 for radiosonde observations), numerical modeling (e.g. *Zülicke and Peters* [2006]) and
1407 theoretical developments (e.g. *Lott et al.* [2010]). *Zülicke and Peters* [2008] have elabo-
1408 rated a parameterization of inertia-gravity wave generation in poleward-breaking rossby
1409 waves, using the cross-stream Lagrangian Rossby number as a central quantity to diagnose
1410 emission, and describing the upward propagation with a WKB approximation. Mesoscale
1411 simulations and observations of ten cases were used to validate their approach.

1412 Motivation to render the non-orographic sources more realistic (e.g. variable in time
 1413 and space) includes evidence from studies of GW sources and needs from GCM modelling:
 1414 different lines of evidence (idealized simulations of *Sato et al.* [2009], balloon observations
 1415 *Hertzog et al.* [2008] and real-case simulations *Plougonven et al.* [2012]) point to oceanic
 1416 regions in the mid-latitudes (i.e. to non-orographic GW sources) as significant sources.
 1417 Regarding modelling, it is evidently unsatisfactory and unphysical not to link emitted
 1418 waves to the flow that is exciting them. In practice, the poor representation of gravity waves
 1419 has been emphasized as a likely cause of important biases in GCMs *Pawson et al.* [2000];
 1420 *Austin et al.* [2003]; *Eyring and Co-Authors* [2007]. Yet more fundamentally, *Haynes*
 1421 [2005] concludes his review of stratospheric dynamics by emphasizing that '*further (and*
 1422 *potential greater) potential uncertainty enters through the extreme difficulty in simulating*
 1423 *potential changes in gravity wave sources in the troposphere.*'

7.2. Transport, mixing and turbulence

1424 Gravity waves contribute in several ways to transport and mixing. *Danielsen et al.* [1991]
 1425 proposed, based on the analysis of airborne measurements, that the differential advection
 1426 due to a low-frequency, large scale wave can induce laminar structures, favoring cross-
 1427 jet transport and mixing. Irreversible mixing is then achieved by smaller-scale gravity
 1428 waves when they break. *Pierce and Fairlie* [1993] thus suggested that inertia-gravity
 1429 waves contribute to transport across the edge of the polar vortex, but called for further
 1430 investigation for this effect to be quantified. Observational evidence for the production of
 1431 laminae by inertia-gravity waves has been described by *Teitelbaum et al.* [1996] and *Pierce*
 1432 *and Grant* [1998]. In another case study explicitly addressing this process, *Tomikawa et al.*
 1433 [2002] found the contribution of inertia-gravity waves to be negligible.

1434 In the numerical simulations of *O'Sullivan and Dunkerton* [1995], significant displace-
1435 ments due to inertia-gravity waves appeared in plots of the potential vorticity near the
1436 tropopause, which were interpreted as a signature of transport. *Moustaoui et al.* [1999]
1437 argued, based on observations and the numerical results of *O'Sullivan and Dunkerton*
1438 [1995], that gravity waves could promote cross-tropopause mixing.

1439 In summary, there is evidence that inertia-gravity waves can produce laminae, and
1440 strong arguments that this will promote mixing. However, quantifying such contribution
1441 of gravity waves to mixing remains an issue.

1442 The breaking of gravity waves will produce small-scale mixing and turbulence (e.g.
1443 *Fritts et al.* [2003]). The latter is of importance for aviation and forecasting of turbulence
1444 *Sharman et al.* [2006, 2012]. It is of particular importance to predict occurrences of clear-
1445 air turbulence (CAT), and the tropopause region near the jet stream is a major source of
1446 CAT events (e.g. [*Kim and Chun*, 2011]). Now, case studies have proven inertia-gravity
1447 waves in the vicinity of the jet-stream to be one mechanism leading to CAT [*Lane et al.*,
1448 2004; *Koch et al.*, 2005] by locally enhancing shear. *Knox et al.* [2008] claimed to predict
1449 CAT from jet-generated IGWs as an application of Lighthill radiation, yet for several
1450 reasons Lighthill radiation here does not apply [*Plougonven et al.*, 2009; *Knox et al.*,
1451 2009]. In fact, further investigation of this case *Trier et al.* [2012] has recently showed
1452 that gravity waves due to convection were at least partly responsible for the turbulence
1453 events analyzed by *Knox et al.* [2008].

7.3. Temperature dependent phenomena

1454 Propagating gravity waves induce reversible temperature fluctuations. These can be of
1455 importance for phenomena that depend on temperature, and particularly those with a

1456 threshold. High frequency waves, as generated from convection and orography, will be
 1457 most efficient in producing substantial temperature fluctuations, yet inertia-gravity waves
 1458 have also been found to contribute.

1459 At high latitudes, gravity waves contribute in this way to polar stratospheric clouds
 1460 (PSCs). Orographic waves are a priori the main source of waves involved [*Carshaw et al.*,
 1461 1998] and for which clear and systematic effects have been documented and the impact
 1462 on PSCs discussed (e.g. *Dörnbrack et al.* [2002]; *Hertzog et al.* [2002a]; *Mann et al.* [2005];
 1463 *Eckermann et al.* [2009]). The contribution from orographic waves is well established
 1464 (e.g. *McDonald et al.* [2009]; *Alexander et al.* [2011]) and is more emphasized than that
 1465 of non-orographic waves. Yet observational case studies have shown that gravity waves
 1466 generated by jets and fronts can also produce PSCs, both in the Antarctic [*Shibata et al.*,
 1467 2003] and in the Arctic *Hitchman et al.* [2003]; *Buss et al.* [2004]; *Eckermann et al.* [2006].

1468 Another example is the freeze-drying of air entering the stratosphere in the Tropical
 1469 Tropopause Layer [*Fueglistaler et al.*, 2009]. Gravity waves contribute to temperature
 1470 fluctuations that will affect the freeze-drying process [*Potter and Holton*, 1995; *Jensen*
 1471 *et al.*, 1996], but it likely does not modify significantly the final water vapor mixing ratios
 1472 [*Jensen and Pfister*, 2004]. In any case, convection is here the relevant source for the
 1473 gravity waves involved.

7.4. In the ocean

1474 One motivation for many of the studies on the limitations of balance (section 4.3)
 1475 comes from the need to understand dissipation in the ocean [*Wunsch and Ferrari*, 2004].
 1476 The prevalent balances (hydorstasy and geostrophy, or some forms of gradient-wind bal-
 1477 ance) and the implied energy cascade to large scales implies a conundrum: what are the

1478 pathways for energy, injected by the wind forcing into geostrophic motions, toward the
1479 small-scales, where it can be dissipated [*McWilliams*, 2003]? Interaction of balanced mo-
1480 tions with internal gravity waves and inertial oscillations constitutes one possible route
1481 [*Müller et al.*, 2005]. Several studies of unbalanced instabilities have been undertaken
1482 to quantify the efficiency of this route (e.g. *Molemaker et al.* [2005] and refs. therein).
1483 Recent high-resolution numerical simulations, both idealized [*Molemaker et al.*, 2010] and
1484 realistic [*Capet et al.*, 2008a, b], have rather emphasized the appearance, at short scales,
1485 of frontal instabilities. Such instabilities are however absent from other high-resolution
1486 simulations of upper-ocean geostrophic turbulence *Klein et al.* [2008], calling for further
1487 investigation. Now, while internal waves or inertial oscillations may play a role in the
1488 forward energy cascade leading to dissipation, it is those forced by their mechanisms, par-
1489 ticularly winds, that are likely involved [*Gertz and Straub*, 2009]. In both cases, the focus
1490 has moved away from spontaneously generated gravity waves.

1491 *Danioux et al.* [2012] have recently investigated specifically the spontaneous generation
1492 of waves from upper-ocean turbulence in an idealized setting. Surface quasi-geostrophy
1493 (SQG) captures well the dynamics of the baroclinically unstable current and the turbulent
1494 mesoscale and submesoscale eddy field. In particular, SQG leads to large Rossby numbers
1495 at small scales [*Juckes*, 1994], and hence spontaneous generation. The generation is hence
1496 very localized (i.e. very intermittent spatially), which is consistent with an exponential
1497 dependence on Rossby number. Once generated however, the waves contribute to a more
1498 homogeneously distributed gravity wave field at depth, where the flow is much weaker.
1499 This generation is small (the energy in the gravity waves is 10^5 times weaker than the
1500 energy in the balanced flow) in comparison to inertia-gravity waves generated by winds

1501 (e.g. *D’Asaro et al.* [1995]). The generation occurs near the grid-scale, and further
 1502 investigations will be necessary to assess more firmly the intensity and realism of such
 1503 generation.

1504 *Polzin* [2008, 2010]It has argued that wave-capture was playing a role in the ocean,
 1505 and more generally that the consideration of horizontally varying background flows fun-
 1506 damentally modifies interactions between waves and the mean flows. However, detailed
 1507 evidence for the occurrence of wave capture in the ocean is still lacking. Observation and
 1508 simulations of this faces one major difficulty in the ocean: near the surface, the major
 1509 source of near-inertial motions are the surface winds, forcing large-scale motions (several
 1510 hundreds to a thousand of kilometers) that are then distorted by the mesoscale, balanced
 1511 vortices (scales of ten to a few hundred kilometers) to finer and finer scales [*Young and*
 1512 *Jelloul, 1997; Klein and Smith, 2001*]. If waves undergoing capture are present, it will be
 1513 at scales smaller than those of the mesoscale vortices, with amplitudes weaker than the
 1514 wind-forced near-inertial oscillations, making them difficult to observe and simulate.

8. DISCUSSION AND PERSPECTIVES

1515 Current knowledge from observations, theory and modelling studies on internal gravity
 1516 waves emanating from jets and fronts has been reviewed. Below we discuss to what extent
 1517 the different threads of investigation tie up together to provide a comprehensive under-
 1518 standing. Focusing on generation mechanisms, we summarize salient points, emphasize
 1519 limitations so as to determine, critically, what should be preserved as robust conclusions,
 1520 and identify what open questions constitute essential challenges.

8.1. On generation mechanisms

1521 The generation mechanism that has most often been invoked is geostrophic adjust-
1522 ment (section 4.1), not only in observations (*Kaplan et al.* [1997]; *Pavelin et al.* [2001];
1523 *Plougonven et al.* [2003]), but also in numerical simulations [*O’Sullivan and Dunkerton,*
1524 1995] and sometimes in analytical studies [*Fritts and Luo,* 1992].

1525 We wish to emphasize that the recurrent reference to geostrophic adjustment turns out
1526 to be unhelpful and argue that it should be avoided. It gives the misleading impression
1527 that there is, readily available, a theoretical paradigm for understanding the emission
1528 of gravity waves by jets and fronts, with foundations going back several decades to the
1529 work of *Rosby* [1938]. We argue that studies of geostrophic adjustment are in fact
1530 unhelpful for three reasons: **1-** they take the imbalance as part of a given initial condition,
1531 hence circumventing the essential difficulty, i.e. to understand how, why and where this
1532 imbalance is produced. **2-** The background flows for which the adjustment problem is
1533 well-posed theoretically, and for which results are available, are simple flows: axially or
1534 zonally symmetric, or with small Rossby number Ro . Relevant flows in practice are more
1535 complex (with spatial and temporal variations, locally large Ro). **3-** The classical scenario
1536 (imbalance propagating away as IGW, leaving a balanced flow behind) is valid only for the
1537 simple configurations afore-mentioned. This does not describe the phenomena observed
1538 and simulated near jets and fronts, where the emission is continuous and no simple, final
1539 adjusted state can be identified.

1540 Now, it is true also that the notion of geostrophic adjustment can be extended, e.g.
1541 to include adjustment of perturbations on a background flow [*Tuyl and Young,* 1982;
1542 *Plougonven and Zeitlin,* 2005]. It can be stretched to describe the response to arbitrary,
1543 time-dependent injection of imbalance [*Weglarz and Lin,* 1997; *Chagnon and Bannon,*

1544 2005a, b]. The traditional initial condition problem is then a particular case, with a
 1545 forcing that is a Dirac δ function of time. With such a generalized definition however,
 1546 geostrophic adjustment loses its precise meaning and encompasses all linear responses to
 1547 a prescribed forcing, for instance, convectively generated waves (diabatic forcing). Hence
 1548 we prefer to preserve a precise meaning for 'geostrophic adjustment' and continue below
 1549 to use it in its traditional acceptation (section 4.1).

1550 In summary, geostrophic adjustment has been repeatedly invoked as the mechanism
 1551 responsible for emission near jets and fronts, partly through lack of a better explanation
 1552 and partly because of the presence of a strong, large-scale imbalance in the vicinity of the
 1553 waves.

1554 The following picture, generalizing the notion of adjustment, has guided intuition: the
 1555 nonlinear evolution of a balanced flow leads to the appearance and growth of localized
 1556 regions of imbalance. This imbalance partly projects onto gravity waves. The 'production'
 1557 of imbalance may persist, so that the flow does not appear to adjust, i.e. the imbalance
 1558 does not decrease and disappear (at least not on timescales of a few inertial periods) and
 1559 gravity waves are continuously emitted. Now this phenomenology, as found in case studies
 1560 (section 3) or in idealized experiments (section 6), differs from that described by classical
 1561 geostrophic adjustment: first, the emission takes place continuously in time, not just in
 1562 a short initial period. Second, the imbalance is not found to decay after the appearance
 1563 of waves: for instance, it is stationary in the dipole. Concomitantly, the flow does not
 1564 evolve simply to a balanced state that can be predicted in advance, e.g. in baroclinic life
 1565 cycles the flow continues its complex, non-linear evolution which comprises imbalance.
 1566 Third, the waves do not necessarily propagate away: for example, waves emitted in the

1567 dipole remain as an inherent part of the dipole. In baroclinic life cycles, only part of
1568 the waves generated near the upper-level jet leak away into the stratosphere. Hence we
1569 believe it is preferable to distinguish the emission by jets and fronts from geostrophic
1570 adjustment (*McIntyre* [2001], p1723 and 1731). Keeping the term 'adjustment' (because
1571 of the guiding image sketched above, which generalizes adjustment to a situation where
1572 the imbalance is continuously forced), we advise to use the terms spontaneous balanced
1573 adjustment [*Zhang*, 2004; *Wang and Zhang*, 2010] or spontaneous adjustment emission
1574 [*Ford et al.*, 2000; *Viúdez and Dritschel*, 2006], or simply spontaneous emission.

1575 Over the past two decades, substantial progress has been achieved in understanding and
1576 quantifying how balanced motions may create imbalance and gravity waves spontaneously
1577 [*Vanneste*, 2013]. We first summarized mechanisms for spontaneous emission that have
1578 been identified analytically (section 4). Lighthill radiation (section 4.2.2), which has
1579 been very inspiring as the first clear mechanism of gravity wave emission from balanced
1580 motions, explains waves that have spatial scales larger than the balanced flow (with $Ro >$
1581 1) generating them. It is useful to explain waves generated from intense vortices such
1582 as cyclones and mesocyclones *Schecter* [2008]. Unbalanced instabilities and transient
1583 generation (sections 4.3 and 4.4) describe how shear couples gravity waves and balanced
1584 motions, leading to emission in the form of unstable modes or transient bursts. These have
1585 scales comparable to or somewhat larger than the Potential Vorticity (PV) anomalies that
1586 are sheared. The range of applicability of these mechanisms remains to be evaluated, but
1587 two points are worth noting: first, unbalanced instabilities have been difficult to exhibit in
1588 dedicated laboratory studies because of their weakness (weak growth rates and/or low level
1589 of saturation, see Section 6.1). Second, the coupling of gravity waves and PV anomalies

1590 in shear may be more relevant for other flow configurations, where other processes such
 1591 as wave-breaking [*Plougonven et al.*, 2010] produce small-scale PV anomalies that are
 1592 subsequently sheared. In other words, these theoretical mechanisms for the spontaneous
 1593 generation of gravity waves from balanced motions have not, so far, been found to apply
 1594 and explain the emission of waves from jets and fronts in real cases.

1595 In all three mechanisms, emission occurs when and where the appropriate scales
 1596 (timescales and spatial scales) match: the scales of the balanced flow and the scales
 1597 of potential inertia-gravity waves, i.e. consistent with the dispersion relation. In the con-
 1598 figurations most relevant to jets and fronts (plane parallel sheared flows), studies have
 1599 emphasized the importance of differential advection (i.e. shear) for coupling balanced
 1600 motions and gravity waves: the slow, balanced motions connect to fast gravity wave mo-
 1601 tions thanks to Doppler shifting. Finally, note that there are many connections between
 1602 these different mechanisms (sections 4.2, 4.3 and 4.4), some unbalanced instabilities being
 1603 described as Lighthill radiation for instance. The fact that these mechanisms do not apply
 1604 easily to cases found in real flows makes it necessary to consider more complex flows.

8.2. Jet Exit Region Emitted (JEREmi) waves

1605 One remarkable outcome from observations and numerical modelling has been the ro-
 1606 bustness of the paradigm put forward by *Uccellini and Koch* [1987], and the dynamical
 1607 understanding obtained since. Observational case studies (sections 3) and idealized sim-
 1608 ulations (6.4 and 6.5) have emphasized jet exit regions, upstream of a ridge and also, less
 1609 frequently, of a trough, as a favored location for large-amplitude, sub-synoptic inertia-
 1610 gravity waves (see section 3 and figure 5). The convergence of different approaches and

1611 the recurrence of this configuration in numerous studies are indications of the robustness
1612 of this result.

1613 Theory has highlighted propagation effects, namely '*wave-capture*', as a mechanism
1614 enhancing IGW in such a region of the flow (section 5.3), the large-scale strain and
1615 vertical shear determining certain of the wave characteristics. Simulations of idealized
1616 baroclinic life cycles (section 6.4) have also highlighted jet exit regions (see Figure 12). A
1617 further simplification of the flow has consisted in restricting to dipoles that have a nearly
1618 steady propagation. Several different modelling studies have robustly identified a low-
1619 frequency wave packet in the front of the dipole, with characteristics consistent with wave
1620 capture, as an inherent part of the dipoles, steadily propagating with them (see Figure
1621 15). The emission mechanism has been explained as the linear response to the differences
1622 between the balanced and the full tendencies (see Section 6.5). The key point is that the
1623 dynamics are linearized on the background of a balanced approximation of the dipole¹³.
1624 The response is not very sensitive to the specific shape of the forcing but rather to the
1625 background flow used in the linearization.

1626 The explanation of waves found in dipoles is an encouraging result, because of the
1627 similarity of these JEREmi (Jet Exit Region Emitted) waves with waves identified in more
1628 complex, idealized flows, and of the similarity of these latter waves with those described in
1629 observational studies. Nonetheless, revisiting observations with the understanding gained
1630 from theory and idealized simulations remains largely to be done in order to assess: **1)**
1631 what proportion of the wave field can be said to be affected by wave capture?, **2)** how
1632 systematic is the presence of such waves in jet exit regions?, **3)** why are amplitudes
1633 found in idealized simulations weaker than those observed?, and **4)** how much IGWs leak

1634 upward or propagate out of the region of strong strain? A further, fundamental issue is 5)
 1635 to understand the impact, for the interaction of waves with the mean flow, of this effect
 1636 due to horizontal variations of the background flow, which is traditionnally ignored.

1637 Other issues, beyond the case of IGW in jet exit regions, remain: how can this under-
 1638 standing guide the elaboration of parameterizations of jets and fronts as gravity waves
 1639 sources? An essential issue that remains to be addressed concerns the role of moisture,
 1640 idealized studies having until now focused on dry dynamics.

8.3. Waves from other processes

1641 JEREmi waves are not the only waves present in the vicinity of jets and fronts, there are
 1642 other potential sources of gravity waves near jets and fronts: first, extant idealized mod-
 1643 elling studies have simulated a richer array of gravity waves, e.g. with waves emanating
 1644 from surface fronts (sections 6.3 and 6.4). Second, these simulations have limitations such
 1645 as the absence of moist processes or of a boundary layer. The parameterizations of these
 1646 small-scale processes have their own uncertainties, yet these processes are of great impor-
 1647 tance: for instance, diabatic heating acts directly on the buoyancy and at small-scales,
 1648 and is therefore a very efficient forcing for gravity waves. Case studies have recurrently
 1649 mentionned the possible important role of moisture (see Section 3). Addition of moisture
 1650 in idealized baroclinic life cycles will have a priori two implications: one is to accelerate
 1651 and intensify the development of baroclinic instability (e.g. *Waite and Snyder* [2012]),
 1652 which should enhance the excitation of gravity waves through spontaneous generation
 1653 [*Reeder and Griffiths*, 1996; *Wang and Zhang*, 2006]. The other is to excite, through
 1654 moist convection, additional waves. Those produced on small-scales from convective cells
 1655 should have strikingly different characteristics (short horitzontal wavelengths (tens of km),

1656 long vertical wavelengths (5-10 km), and correspondingly high intrinsic frequencies). On
1657 the other hand, the large scale envelope of convection may contribute to the gravity wave
1658 field on larger scale, and this contribution will be more difficult to isolate.

1659 Idealized moist simulations will contribute to guide our understanding, as for the im-
1660 pacts of moisture on the predictability of mesoscale weather [*Zhang et al.*, 2007], but they
1661 necessarily involve parameterizations convective and boundary layer processes, which are
1662 themselves quite uncertain. The implication is that further studies of moist generation of
1663 gravity waves from fronts will call strongly for observational constraints. Combined stud-
1664 ies involving both simulations and observations should be an important step to provide a
1665 complete description of waves near moist fronts [*Zhang et al.*, 2011].

1666 In a similar vein, additional complexity relative to idealized baroclinic life cycles may
1667 come from the generation of gravity waves from small scale turbulent motions, e.g. emis-
1668 sion from shear instability. Previous studies on the subject have conclusively ruled out a
1669 straightforward, linear connection, but studies of the nonlinear development of the shear
1670 instability have shown that this mechanism should be considered as a source of grav-
1671 ity waves (section 4.5). Yet, the numerical configurations used remained quite idealized.
1672 Here again, observations will play a key role in constraining the realism of numerical sim-
1673 ulations. A fundamental difficulty here again is the complexity of the background flow,
1674 involving a wide range of scales from the synoptic motions to the small-scale turbulence.

8.4. Perspectives

1675 Now, both points above have emphasized the complexity that will be encountered in
1676 exploring gravity waves generated by jets and fronts as one explores finer scales. Moist
1677 convection and small-scale turbulence are themselves challenges for modelling and ob-

1678 servation. It will likely be impossible to draw a simple, deterministic and convincing
 1679 picture of the way gravity waves are generated from these processes in a complex flow
 1680 environment such as a cold front within a baroclinic wave. Yet, the demand from appli-
 1681 cations (parameterizations for GCMs, forecasting of turbulence) may not call for such a
 1682 deterministic picture. Observations should play a key role (see also challenges discussed
 1683 in section 3.2). Global high-resolution datasets have been obtained, and the combined
 1684 use of different observational platforms along with modeling promises to provide global
 1685 descriptions of the gravity wave field in coming years. We believe one way forward will
 1686 be to analyze such high-resolution datasets to produce flow-dependent characterizations
 1687 of gravity waves (e.g. rather than quantify the mean GW activity at a given location,
 1688 quantify it relative to flow configuration). This can bring practical answers to the needs
 1689 of climate and forecast models. Presently, GCMs that include a parameterization of non-
 1690 orographic waves are the exception, and there is much room for improving on the heuristic
 1691 relation used to connect the emitted waves to the tropospheric flow. The trend towards
 1692 stochastic parameterizations (*Palmer* [2001], and *Eckermann* [2011]; *Lott et al.* [2012a]
 1693 for gravity waves specifically) is in phase with new descriptions of the gravity wave field
 1694 [*Hertzog et al.*, 2012].

1695 The perspective of quantifying jets and fronts as sources of gravity waves, and hence
 1696 of measuring and parameterizing their variability, will make GCMs more physical, and
 1697 should improve their internal variability. It will also set the stage for investigations of the
 1698 variability of this forcing, of its evolution in a changing climate and of the implications,
 1699 as questioned by *Haynes* [2005] (see Section 7.1).

GLOSSARY

1700 **Balanced models:** approximate model that relies on balance relations which diag-
 1701 nostically relates several variables (e.g. velocity and pressure in geostrophic balance) to
 1702 simplify the dynamics. Evolution of the flow typically reduces to one equation (conserva-
 1703 tion of Potential Vorticity), and the balance relations (e.g. hydrostasy and geostrophy for
 1704 the quasi-geostrophic approximation) make it possible to *invert* the Potential Vorticity to
 1705 recover all fields, and in particular the velocity (see *Hoskins et al.* [1985], and section 4.2).

1706 **Baroclinicity:** measure of how the isolines of the density field and of the pres-
 1707 sure field are misaligned. In the atmosphere, baroclinicity is strongest where there are
 1708 strong horizontal thermal gradients, as in mid-latitudes, and is associated to vertical shear
 1709 through thermal wind balance (e.g. *Holton* [1992]).

1710 **Inertia-gravity wave:** gravity wave having a low frequency (close to the lower
 1711 bound of the gravity wave spectrum, i.e. f the Coriolis parameter). See section 1.

1712 **Intrinsic frequency:** : frequency in the frame moving with the fluid. The intrinsic
 1713 frequency $\hat{\omega}$ is related to the ground based frequency ω by $\hat{\omega} = \omega - \mathbf{k} \cdot \mathbf{U}$, where
 1714 \mathbf{k} is the wavenumber and \mathbf{U} is the background wind (see section 5).

1715 **Polar Night Jet:** intense westerly jet that forms in the winter stratosphere, at high
 1716 latitudes (typically 60°) and altitudes higher than 20 km. It encloses the polar vortex,
 1717 and isolates it from mid-latitude air.

1718 **Rossby number:** ratio U/fL , where U is a typical order of magnitude for wind
 1719 velocity, L is a typical horizontal scale, and f is the Coriolis parameter. This compares
 1720 the advective timescale L/U with the inertial timescale $1/f$, and is typically small at
 1721 mid-latitudes at synoptic scales.

1722 **Superpressure balloons:** balloons used for atmospheric measurements, with an
 1723 envelope that is not extendable. At the level where the balloons drift, the gas inside has
 1724 a pressure larger than the environment, so that the balloon remains fully inflated and the
 1725 full device has a constant density. It therefore drifts along an isopycnic surface, and may
 1726 be considered a quasi-Lagrangian tracer (see *Hertzog et al.* [2007] and section 2.3).

1727 **Unbalanced instabilities:** instabilities in a rotating fluid that involves unbalanced
 1728 motions. These are of interest in regimes where balance is expected or even dominant
 1729 (e.g. weak Rossby number), and hence the term preferentially refers to instabilities that
 1730 couple balanced and unbalanced motions (section 4.3).

1731 **ACKNOWLEDGMENTS.** The authors are grateful to C. Snyder and J. Vanneste for careful reading of the
 1732 manuscript and judicious remarks, to D. Durran for precious advice, to S. Wang for proofreading and to A. Kara for
 1733 providing time to advance this project. They also wish to thank C. Snyder, J. Vanneste, O. Bühler, M.E. McIntyre,
 1734 R. Rotunno, M. Reeder, T. Lane, C. Epifanio, T. Dunkerton, A. Medvedev, F. Lott and A. Hertzog for instructive
 1735 and stimulating discussions on the subject. FZ acknowledges funding support from US National Science Foundation
 1736 under grants: 0904635 and 1114849. RP acknowledges support from the European EMBRACE project.

NOTES

1. *Trexler and Koch* [2000] have compared observations from wind profilers and from a surface mesonet, and concluded
 1737 that the latter may be limited to detecting waves that affect primarily the lower atmosphere.
2. The observational filter corresponding to superpressure balloons and radiosondes is also displayed
3. MLS, AMSU-A, AIRS, GPS, and CLAES; AIRS has much better horizontal resolution
4. The observations have their own biases, and in particular underestimate gravity waves with high intrinsic frequencies,
 which primarily affects orographic waves in this region *Plougonven et al.* [2008].
5. Some numerical simulations are also included in the present section because they are closely tied to the analytical results.
6. which has been refined as spontaneous balance adjustment in *Wang and Zhang* [2010], so as to avoid any confusion with a
 generalization of geostrophic adjustment that would simply include adjustment to higher-order balances than geostrophy
 (e.g. cyclo-geostrophic balance, see *Holton* [1992])
7. These low order models can be interpreted as describing the motions of a swinging spring [*Lynch*, 2002], or of a spring
 tied to a pendulum [*Vanneste*, 2006, 2008]. The small parameter equivalent to the Rossby number is the ratio of the
 (slow) pendulum to the (fast) spring oscillation frequencies.
8. This is in contrast with unbalanced instabilities, which require an initial deviation (however small) from the unstable
 balanced state *Vanneste* [2008]. The growing amplitudes of the waves will depend on this initial condition, linear theory
 providing only the growth rate, and the final amplitude of the waves will depend on the nonlinear saturation of the
 instability (e.g. *Gula et al.* [2009a]).
9. Shear instabilities are here treated separately from the other unbalanced instabilities because they occur on small scales
 such that the background rotation is generally not considered. In other terms, they occur at large Rossby numbers, such
 that balance (and imbalance) are not relevant for their development.
10. The dispersion relationships for waves in stratified fluid and in shallow water differ crucially here: in shallow water,
 short-scale waves necessarily have large frequencies and fast group velocities (recall $\omega^2 = f^2 + gH(k^2 + l^2)$). Wave
 capture can not occur in shallow water flows with Froude numbers smaller than unity.

- 11.The example of two-dimensional frontogenesis simulations had shown how spurious gravity waves could easily be produced and mistaken for spontaneously generated waves (see discussion of *Snyder et al.* [1993] on *Gall et al.* [1988]).
- 12.the study of *Bush et al.* [1995], which identified waves coming from the surface cold front, discarded them as a numerical artefact.
- 13.The forcing is also deduced from this balanced dipole.

REFERENCES

- 1738 Afanasyev, Y. (2003), Spontaneous emission of gravity waves by interacting vortex dipoles
 1739 in a stratified fluid: laboratory experiments, *Geophys. Astrophys. Fluid Dyn.*, *97*(2),
 1740 79–95.
- 1741 Afanasyev, Y., P. Rhines, and E. Lindhal (2008), Emission of inertial waves by baroclin-
 1742 ically unstable flows: Laboratory experiments with altimetric imaging velocimetry, *J.*
 1743 *Atmos. Sci.*, *65*(doi:10.1175/2007JAS2336.1), 250–262.
- 1744 Alexander, M., and co authors (2008), Global estimates of gravity wave momentum
 1745 flux from High Resolution Dynamics Limb Sounder Observations, *J. Geophys. Res.*,
 1746 *113*(D15S18), doi:10.1029/2007JD008,807.
- 1747 Alexander, M., and et. al. (2010), Recent developments in gravity-wave effects in climate
 1748 models and the global distribution of gravity-wave momentum flux from observations
 1749 and models, *Q.J.R. Meteorol. Soc.*, *136*, 1103–1124.
- 1750 Alexander, M., J. Holton, and D. Durran (1995), The gravity wave response above deep
 1751 convection in a squall line simulation, *J. Atmos. Sci.*, *52*, 2212–2226.
- 1752 Alexander, S., A. Klecociuk, M. Pitts, A. McDonald, and A. Arevalo-Torres (2011), The
 1753 effect of orographic gravity waves on Antarctic polar stratospheric cloud occurrences
 1754 and composition, *J. Geophys. Res.*, *116*(D06109), doi:10.1029/2010JD015,184.

- 1755 Andrews, D., J. Holton, and C. Leovy (1987), *Middle atmosphere dynamics*, Academic
1756 Press.
- 1757 Aspden, J., and J. Vanneste (2009), Elliptical instability of a rapidly rotating, strongly
1758 stratified fluid, *Phys. Fluids*, *21*, 074,104.
- 1759 Aspden, J., and J. Vanneste (2010), *IUTAM Symposium on Turbulence in the Atmo-*
1760 *sphere and Oceans*, chap. Inertia-gravity-wave generation: a geometric-optic approach,
1761 Springer.
- 1762 Austin, J., et al. (2003), Uncertainties and assessment of chemistry-climate models of the
1763 stratosphere, *Atm. Chem. Phys.*, *3*, 1–27.
- 1764 Bacmeister, J., S. D. Eckermann, P. Newman, L. Lait, K. Chan, M. Loewenstein, M. Prof-
1765 fitt, and B. L. Gary (1996), Stratospheric horizontal wavenumber spectra of winds,
1766 potential temperature and atmospheric tracers observed by high-altitude aircraft, *J.*
1767 *Geophys. Res.*, *101*, 9441–9470.
- 1768 Badulin, S., and V. Shrira (1993), On the irreversibility of internal wave dynamics due
1769 to wave trapping by mean flow inhomogenities. Part 1: Local analysis, *J. Fluid Mech.*,
1770 *251*, 21–53.
- 1771 Baer, F., and J. Tribbia (1977), On complete filtering of gravity modes through nonlinear
1772 initialization, *Mon. Weath. Rev.*, *105*, 1536–1539.
- 1773 Bakas, N., and B. Farrell (2008), Momentum and energy transport by gravity waves in
1774 stochastically driven stratified flows. Part II: radiation of gravity waves from a Gaussian
1775 jet, *J. Atmos. Sci.*, *65*(7), 2308–2325.
- 1776 Bakas, N., and B. Farrell (2009a), Gravity waves in a horizontal shear flow. Part I: Growth
1777 mechanisms in the absence of potential vorticity perturbations., *J. Phys. Oceanogr.*, *39*,

1778 481–496.

1779 Bakas, N., and B. Farrell (2009b), Gravity waves in a horizontal shear flow. Part II: Inter-
 1780 action between gravity waves and potential vorticity perturbations., *J. Phys. Oceanogr.*,
 1781 *39*, 497–511.

1782 Batchelor, G. (1967), *An introduction to fluid dynamics*, Cambridge University Press.

1783 Beres, J., M. Alexander, and J. Holton (2004), A method of specifying the gravity wave
 1784 spectrum above convection based on latent heating properties and background wind, *J.*
 1785 *Atmos. Sci.*, *61*, 324–337.

1786 Beres, J., R.R.Garcia, B. Boville, and F. Sassi (2005), Implementation of a gravity
 1787 wave source spectrum parameterization dependent on the properties of convection
 1788 in the Whole Atmosphere Community Climate Model (WACCM), *J. Geophys. Res.*,
 1789 *110*(D10108), doi:10.1029/2004JD005,504.

1790 Bluestein, H., and M. Jain (1985), Formation of mesoscale lines of precipitation - severe
 1791 squall lines in Oklahoma during the spring, *J. Atmos. Sci.*, *42*(16), 1711–1732.

1792 Blumen, W. (1972), Geostrophic adjustment, *Reviews of Geophysics and Space Physics*,
 1793 *10*(2), 485–528.

1794 Blumen, W., and R. Wu (1995), Geostrophic adjustment: frontogenesis and energy con-
 1795 version, *J. Phys. Oceanogr.*, *25*, 428–438.

1796 Boccara, G., A. Hertzog, R. Vincent, and F. Vial (2008), Estimation of gravity-wave
 1797 momentum fluxes and phase speeds from long-duration stratospheric balloon flights. 1.
 1798 Theory and simulations, *J. Atmos. Sci.*, *65*, 3042–3055.

1799 Bokhove, O., and T. Shepherd (1996), On hamiltonian balanced dynamics and the slowest
 1800 invariant manifold, *J. Atmos. Sci.*, *53*, 276–297.

- 1801 Bosart, L., W. Bracken, and A. Seimon (1998), A study of cyclone mesoscale structure
1802 with emphasis on a large-amplitude inertia-gravity wave, *Mon. Weath. Rev.*, *126*, 1497–
1803 1527.
- 1804 Brunet, G., and M. Montgomery (2002), Vortex Rossby waves on smooth circular vortices.
1805 Part I. Theory, *Dyn. Atmos. Ocean*, *35*, 153–177.
- 1806 Bühler, O. (2009), *Waves and mean flows*, 341pp pp., Cambridge University Press.
- 1807 Bühler, O., and M. McIntyre (1999), On shear-generated gravity waves that reach the
1808 mesosphere. Part II: wave propagation, *J. Atmos. Sci.*, *56*, 3764–3773.
- 1809 Bühler, O., and M. McIntyre (2003), Remote recoil: a new wave-mean interaction effect,
1810 *J. Fluid Mech.*, *492*, 207–230.
- 1811 Bühler, O., and M. McIntyre (2005), Wave capture and wave-vortex duality, *J. Fluid*
1812 *Mech.*, *534*, 67–95.
- 1813 Bühler, O., M. McIntyre, and J. Scinocca (1999), On shear-generated gravity waves that
1814 reach the mesosphere. Part I: wave generation, *J. Atmos. Sci.*, *56*, 3749–3763.
- 1815 Bush, A., J. McWilliams, and W. Peltier (1995), Origins and evolution of imbalance in
1816 synoptic-scale baroclinic wave life cycles, *J. Atmos. Sci.*, *52*, 1051–1069.
- 1817 Buss, S., A. Hertzog, C. Hostettler, T. Bui, D. Lüthi, and H. Wernli (2004), Analysis of
1818 a jet stream induced gravity wave associated with an observed stratospheric ice cloud
1819 over Greenland, *Atmos. Chem. Phys.*, *4*, 1680–7324/acp/2004–4–1183.
- 1820 Cahn, A. (1945), An investigation of the free oscillations of a simple current system, *J.*
1821 *Atmos. Sci.*, *2*(2), 113–119.
- 1822 Camassa, R. (1995), On the geometry of an atmospheric slow manifold, *Physica D*, *84*,
1823 357–397.

- 1824 Capet, X., J. McWilliams, M. Molemaker, and A. Schepetkin (2008a), Mesoscale to sub-
 1825 mesoscale transition in the california current system: Part I: Flow structure, eddy flux
 1826 and observational tests, *J. Phys. Oceanogr.*, *38*, 44–69.
- 1827 Capet, X., J. McWilliams, M. Molemaker, and A. Schepetkin (2008b), Mesoscale to sub-
 1828 mesoscale transition in the california current system: Part III: energy balance and flux,
 1829 *J. Phys. Oceanogr.*, *38*, 2256–2269.
- 1830 Carslaw, K., et al. (1998), Increased stratospheric ozone depletion due to mountain-
 1831 induced atmospheric waves, *Nature*, *391*, 675–678.
- 1832 Chagnon, J., and P. Bannon (2005a), Wave response during hydrostatic and geostrophic
 1833 adjustment. Part I: Transient dynamics, *J. Atmos. Sci.*, *62*, 1311–1329.
- 1834 Chagnon, J., and P. Bannon (2005b), Wave response during hydrostatic and geostrophic
 1835 adjustment. Part I: Potential vorticity conservation and energy partitioning, *J. Atmos.*
 1836 *Sci.*, *62*, 1330–1345.
- 1837 Charney, J. (1948), On the scale of atmospheric motions, *Geophys. Publ. Oslo*, *17*(2),
 1838 1–17.
- 1839 Charron, M., and E. Manzini (2002), Gravity waves from fronts: parameterization and
 1840 middle atmosphere response in a general circulation model, *J. Atmos. Sci.*, *59*, 923–941.
- 1841 Chimonas, G., and J. Grant (1984), Shear excitation of gravity waves. Part II: upscale
 1842 scattering from Kelvin-Helmholtz waves, *J. Atmos. Sci.*, *41*, 2278–2288.
- 1843 Clark, T., T. Hauf, and J. Kuettner (1986), Convectively forced internal gravity waves:
 1844 results from two-dimensional numerical experiments, *Q.J.R. Meteorol. Soc.*, *112*, 899–
 1845 925.

- 1846 Cram, J., R. Pielke, and W. Cotton (1992), Numerical simulation and analysis of a pre-
1847 frontal squall line. Part II: Propagation of the squall line as an internal gravity wave,
1848 *J. Atmos. Sci.*, *49*, 209–225.
- 1849 Cunningham, P., and D. Keyser (2000), Analytical and numerical modelling of jet streaks:
1850 barotropic dynamics, *Q.J.R. Meteorol. Soc.*, *126*, 3187–3217.
- 1851 Danielsen, E., R. S. Hipskind, W. Starr, J. Vedder, S. Gaines, D. Kley, and K. Kelly
1852 (1991), Irreversible transport in the stratosphere by internal waves of short vertical
1853 wavelength, *J. Geophys. Res.*, *96*(D9), 17,433–17,452.
- 1854 Danioux, E., J. Vanneste, P. Klein, and H. Sasaki (2012), Spontaneous inertia-gravity
1855 wave generation by surface-intensified turbulence, *J. Fluid Mech.*, *in press*.
- 1856 D’Asaro, E., C. Eriksen, M. Levine, P. Niiler, C. Paulson, and P. V. Meurs (1995), Upper-
1857 ocean inertial currents forced by a strong storm. Part I: data and comparison with linear
1858 theory, *J. Phys. Oceanogr.*, *25*, 2909–2936.
- 1859 Davis, C., and K. Emanuel (1991), Potential vorticity diagnostics of cyclogenesis, *J. At-*
1860 *mos. Sci.*, *119*, 1929–1953.
- 1861 Davis, P., and W. Peltier (1979), Some Characteristics of the Kelvin-Helmholtz and Res-
1862 onant Overreflection Modes of Shear Flow Instability and of Their Interaction through
1863 Vortex Pairing, *J. Atmos. Sci.*, *36*(12), 2394–2412.
- 1864 Dewar, W., and P. Killworth (1995), Do fast gravity waves interact with geostrophic
1865 motions?, *Deep-Sea Research*, *42*(7), 1063–1081.
- 1866 Dörnbrack, A., T. Birner, A. Fix, H. Flentje, A. Meister, H. Schmid, E. V. Browell, and
1867 M. J. Mahoney (2002), Evidence for inertia-gravity waves forming polar stratospheric
1868 clouds over Scandinavia, *J. Geophys. Res.*, *107*(D20), 8287, 10.1029/2001JD000,452.

- 1869 Dritschel, D., and J. Vanneste (2006), Instability of a shallow-water potential-vorticity
 1870 front, *J. Fluid Mech.*, *561*, 237–254.
- 1871 Dritschel, D., and A. Viúdez (2003), A balanced approach to modelling rotating stably
 1872 stratified geophysical flows, *J. Fluid Mech.*, *488*, 213–150.
- 1873 Dunkerton, T. (1984), Inertia-gravity waves in the stratosphere, *J. Atmos. Sci.*, *41*, 3396–
 1874 3404.
- 1875 Eady, E. (1949), Long waves and cyclone waves, *Tellus*, *1*, 33–52.
- 1876 Eckermann, S. (2011), Explicitly Stochastic Parameterization of Nonorographic Gravity
 1877 Wave Drag, *J. Atmos. Sci.*, *68*(8), 1749–1765.
- 1878 Eckermann, S., and C. Marks (1996), An idealized ray model of gravity wave tidal inter-
 1879 actions, *J. Geophys. Res.*, *101*, 21,195–21,212.
- 1880 Eckermann, S., and C. Marks (1997), GROGRAT: a new model of the global propagation
 1881 and dissipation of atmospheric gravity waves, *Adv. Space. Res.*, *20*(6), 1253–1256.
- 1882 Eckermann, S., and R. Vincent (1993), VHF radar observations of gravity-wave production
 1883 by cold fronts over Southern Australia, *J. Atmos. Sci.*, *50*, 785–806.
- 1884 Eckermann, S., A. Dörnbrack, S. Vosper, H. Flentje, M. Mahoney, T. P. Bui, and
 1885 K. Carslaw (2006), Mountain wave-induced polar stratospheric cloud forecasts for air-
 1886 craft science flights during SOLVE/THESEO 2000, *Weather and Forecasting*, *21*, 42–68.
- 1887 Eckermann, S., L. Hoffmann, M. H. and D.L. Wu, and M. Alexander (2009), Antarctic
 1888 NAT PSC belt of June 2003: Observational validation of the mountain wave seeding
 1889 hypothesis, *Geophys. Res. Lett.*, *36*(L02807), doi:10.1029/2008GL036,629.
- 1890 Einaudi, F., A. Bedard, and J. Finnigan (1989), A Climatology of Gravity Waves and
 1891 Other Coherent Disturbances at the Boulder Atmospheric Observatory during Mar-

- 1892 chApril 1984, *jas*, 46, 303–329.
- 1893 Eom, J. (1975), Analysis of the internal gravity wave occurrence of 19 April 1970 in the
1894 Midwest, *Mon. Weath. Rev.*, 103, 217–226.
- 1895 Ern, M., and P. Preusse (2011), Implications for atmospheric dynamics derived from
1896 global observations of gravity wave momentum flux in stratosphere and mesosphere, *J.*
1897 *Geophys. Res.*, 116(D19107), doi:10.1029/2011JD015,821.
- 1898 Ern, M., P. Preusse, M. Alexander, and C. Warner (2004), Absolute values of grav-
1899 ity wave momentum flux derived from satellite data, *J. Geophys. Res.*, 109(D20103),
1900 doi:10.1029/2006JD007,327.
- 1901 Esler, J., and L. Polvani (2004), Kelvin-Helmholtz instability of potential vorticity layers:
1902 a route to mixing, *J. Atmos. Sci.*, 61, 1392–1405.
- 1903 Eyring, V., and Co-Authors (2007), Multimodel projections of stratospheric ozone in the
1904 21st century, *J. Geophys. Res.*, 112(D16303), doi:10.1029/2006JD008,332.
- 1905 Fetzer, E., and J. Gille (1994), Gravity wave variance in LIMS temperatures. Part I:
1906 Variability and comparison with background winds, *J. Atmos. Sci.*, 51, 2461–2483.
- 1907 Ford, R. (1994a), The response of a rotating ellipse of uniform potential vorticity to gravity
1908 wave radiation, *Phys. Fluids*, 6(11), 3694–3704.
- 1909 Ford, R. (1994b), The instability of an axisymmetric vortex with monotonic potential
1910 vorticity in rotating shallow water, *J. Fluid Mech.*, 280, 303–334.
- 1911 Ford, R. (1994c), Gravity wave radiation from vortex trains in rotating shallow water, *J.*
1912 *Fluid Mech.*, 281, 81–118.
- 1913 Ford, R., M. E. McIntyre, and W. A. Norton (2000), Balance and the slow quasimanifold:
1914 some explicit results, *J. Atmos. Sci.*, 57, 1236–1254.

- 1915 Ford, R., M. E. McIntyre, and W. A. Norton (2002), Reply, *J. Atmos. Sci.*, *59*, 2878–2882.
- 1916 Fovell, R., D. Durran, and J. Holton (1992), Numerical simulations of convectively gen-
1917 erated stratospheric gravity waves, *J. Atmos. Sci.*, *49*, 1427–1442.
- 1918 Fritts, D. (1980), Simple stability limits for vertically propagating unstable modes in a
1919 $\tanh(z)$ velocity profile with a rigid lower boundary, *J. Atmos. Sci.*, *37*, 1642–1648.
- 1920 Fritts, D. (1982), Shear excitation of atmospheric gravity waves, *J. Atmos. Sci.*, *39*, 1936–
1921 1952.
- 1922 Fritts, D. (1984), Shear excitation of atmospheric gravity waves. 2: Nonlinear radiation
1923 from a free shear-layer, *J. Atmos. Sci.*, *41*, 524–537.
- 1924 Fritts, D., and M. Alexander (2003), Gravity wave dynamics and effects in the middle
1925 atmosphere, *Reviews of Geophysics*, *41*(1), 1003.
- 1926 Fritts, D., and G. Nastrom (1992), Sources of mesoscale variability of gravity waves. Part
1927 II: Frontal, convective, and jet stream excitation, *J. Atmos. Sci.*, *49*(2), 111–127.
- 1928 Fritts, D., C. Bizon, J. Werne, and C. Meyer (2003), Layering accompanying turbu-
1929 lence generation due to shear instability and gravity-wave breaking, *J. Geophys. Res.*,
1930 *108*(D8), 8452.
- 1931 Fritts, D., B. Williams, C. She, J. Vance, M. Rapp, F.-J. Lübken, A. Müllemann,
1932 F. Schmidlin, and R. Goldberg (2004), Observations of extreme temperature and wind
1933 gradients near the summer mesopause with the MaCWAVE/MIDAS rocket campaign,
1934 *Geophys. Res. Lett.*, *31*(L24S06), doi:10.1029/2003GL019,389.
- 1935 Fritts, D. C., and Z. Luo (1992), Gravity wave excitation by geostrophic adjustment of
1936 the jet stream. Part I: Two-dimensional forcing, *J. Atmos. Sci.*, *49*(8), 681–697.

- 1937 Fueglistaler, S., A. E. Dessler, T. J. Dunkerton, I. Folkins, Q. Fu, and P. W. Mote (2009),
1938 Tropical Tropopause Layer, *Rev. Geophys.*, *47*(RG1004), doi:10.1029/2008RG000,267.
- 1939 Gall, R., R. Williams, and T. Clark (1987), On the minimum scale of surface fronts, *J.*
1940 *Atmos. Sci.*, *44*, 2562–2574.
- 1941 Gall, R., R. Williams, and T. Clark (1988), Gravity waves generated during frontogenesis,
1942 *J. Atmos. Sci.*, *45*(15), 2204–2219.
- 1943 Garner, S. (1989), Fully Lagrangian numerical solutions of unbalanced frontogenesis and
1944 frontal collapse, *J. Atmos. Sci.*, *46*(6), 717–739.
- 1945 Geller, M., and J. Gong (2010), Gravity wave kinetic, potential, and vertical fluctu-
1946 ation energies as indicators of different frequency gravity waves, *J. Geophys. Res.*,
1947 *115*(D11111), doi:10.1029/2009JD012,266.
- 1948 Gertz, A., and D. Straub (2009), Near-Inertial Oscillations and the Damping of Midlati-
1949 tude Gyres: A Modeling Study, *J. Phys. Oceanogr.*, *39*, 23382350.
- 1950 Gettelman, A., P. Hoor, L. Pan, W. Randel, M. Hegglin, and T. Birner (2011), The
1951 extratropical upper troposphere and lower stratosphere, *Rev. Geophys.*, *49*(RG3003),
1952 2011RG000,355.
- 1953 Gill, A. E. (1982), *Atmosphere-ocean dynamics*, 662p pp., Academic Press.
- 1954 Glendening, J. (1993), Nonlinear displacement of the geostrophic velocity jet created by
1955 mass imbalance, *J. Atmos. Sci.*, *50*, 1617–1628.
- 1956 Gong, J., and M. Geller (2010), Vertical fluctuation energy in United States high ver-
1957 tical resolution radiosonde data as an indicator of convective gravity wave sources, *J.*
1958 *Geophys. Res.*, *115*(D11110), doi:10.1029/2009JD012,265.

- 1959 Gong, J., M. Geller, and L. Wang (2008), Source spectra information de-
 1960 rived from U.S. high-resolution radiosonde data, *J. Geophys. Res.*, *113*(D10106),
 1961 doi:10.1029/2007JD009,252.
- 1962 Gossard, E., and W. Hooke (1975), *Waves in the atmosphere. Developments in atmo-*
 1963 *spheric science II*, 456pp pp., Elsevier Scientific Publishing Company.
- 1964 Griffiths, M., and M. J. Reeder (1996), Stratospheric inertia-gravity waves generated in
 1965 a numerical model of frontogenesis. I: Model solutions, *Q.J.R. Meteorol. Soc.*, *122*,
 1966 1153–1174.
- 1967 Grivet-Talocia, S., F. Einaudi, W. Clark, R. Dennett, G. Nastrom, and T. VanZandt
 1968 (1999), A 4-yr Climatology of Pressure Disturbances Using a Barometer Network in
 1969 Central Illinois, *J. Atmos. Sci.*, *127*(7), 1613–1629.
- 1970 Guest, F., M. Reeder, C. Marks, and D. Karoly (2000), Inertia-gravity waves observed in
 1971 the lower stratosphere over Macquarie Island, *J. Atmos. Sci.*, *57*, 737–752.
- 1972 Gula, J., R. Plougonven, and V. Zeitlin (2009a), Ageostrophic instabilities of fronts in a
 1973 channel in a stratified rotating fluid, *J. Fluid Mech.*, *627*, 485–507.
- 1974 Gula, J., V. Zeitlin, and R. Plougonven (2009b), Instabilities of two-layer shallow-water
 1975 flows with vertical shear in the rotating annulus, *J. Fluid Mech.*, *638*, 27–47.
- 1976 Hart, J. (1972), A laboratory study of baroclinic instability, *Geophys. Astrophys. Fluid*
 1977 *Dyn.*, *3*, 181–209.
- 1978 Haynes, P. (2005), Stratospheric dynamics, *Ann. Rev. Fluid Mech.*, *37*, 263–293.
- 1979 Haynes, P., and J. Anglade (1997), The vertical-scale cascade in atmospheric tracers due
 1980 to large-scale differential advection, *J. Atmos. Sci.*, *54*, 1121–1136.

- 1981 Hertzog, A., and F. Vial (2001), A study of the dynamics of the equatorial lower strato-
1982 sphere by use of ultra-long-duration balloons 2. Gravity waves, *J. Geophys. Res.*, *106*,
1983 22,745–22,761.
- 1984 Hertzog, A., C. Souprayen, and A. Hauchecorne (2001), Observation and backward tra-
1985 jectory of an inertia-gravity wave in the lower stratosphere, *Annales Geophysicae*, *19*,
1986 1141–1155.
- 1987 Hertzog, A., F. Vial, A. Dörnbrack, S. Eckermann, B. Knudsen, and J.-P. Pom-
1988 mereau (2002a), In situ observations of gravity waves and comparisons with nu-
1989 merical simulations during the SOLVE/THESEO 2000 campaign, *J. Geophys. Res.*,
1990 *D20*(doi:10.1029/2001JD001025), 8292.
- 1991 Hertzog, A., F. Vial, C. Mechoso, C. Basdevant, and P. Cocquerez (2002b), Quasi-
1992 Lagrangian measurements in the lower stratosphere reveal an energy peak associated
1993 with near-inertial waves, *Geophys. Res. Let.*, *29*(8), 70.
- 1994 Hertzog, A., G. Boccara, R. Vincent, F. Vial, and P. Coquerez (2008), Estimation of
1995 gravity-wave momentum fluxes and phase speeds from long-duration stratospheric bal-
1996 loon flights. 2. Results from the Vorcore campaign in Antarctica, *J. Atmos. Sci.*, *65*,
1997 3056–3070.
- 1998 Hertzog, A., M. Alexander, and R. Plougonven (2012), On the probability density func-
1999 tions of gravity waves momentum flux in the stratosphere, *J. Atmosph. Sci.*
- 2000 Hertzog, A., et al. (2007), Stratéole/Vorcore - Long duration, superpressure balloons to
2001 study the antarctic stratosphere during the 2005 winter, *J. Ocean. Atmos. Tech.*, *24*,
2002 2048–2061.

- 2003 Hines, C. (1968), A possible source of waves in noctilucent clouds, *J. Atmos. Sci.*, *25*,
 2004 937–942.
- 2005 Hirota, I., and T. Niki (1985), A statistical study of inertia-gravity waves in the middle
 2006 atmosphere, *J. Meteor. Soc. Japan*, *63*, 1055–1065.
- 2007 Hitchman, M., M. Buker, G. Tripoli, E. Browell, W. Grant, T. McGee, and J. Burris
 2008 (2003), Nonorographic generation of Arctic polar stratospheric clouds during December
 2009 1999, *J. Geophys. Res.*, *108*(D5), 8325.
- 2010 Holton, J. R. (1992), *An introduction to dynamic meteorology*, third ed., Academic Press,
 2011 London.
- 2012 Hoskins, B., M. McIntyre, and A. Robertson (1985), On the use and significance of isen-
 2013 tropic potential vorticity maps, *Q.J.R. Meteorol. Soc.*, *111*(470), 877–946.
- 2014 Hoskins, B. J. (1982), The mathematical theory of frontogenesis, *Ann. Rev. Fluid Mech.*,
 2015 *14*, 131–151.
- 2016 Hoskins, B. J., and F. P. Bretherton (1972), Atmospheric frontogenesis models: mathe-
 2017 matical formulation and solution, *J. Atmos. Sci.*, *29*, 11–37.
- 2018 Jacoby, T., P. Read, P. Williams, and R. Young (2011), Generation of inertia-gravity
 2019 waves in the rotating thermal annulus by a localized boundary layer instability, *Geophys.*
 2020 *Astrophys. Fluid Dyn.*, *iFirst: 11 March 2011*(10.1080/03091929.2011.560151), 1–21.
- 2021 Jensen, E., and L. Pfister (2004), Transport and freeze-drying in the tropical tropopause
 2022 layer, *J. Geophys. Res.*, *109*(D02207), doi:10.1029/2003JD004,022.
- 2023 Jensen, E., O. Toon, L. Pfister, and H. Selkirk (1996), Dehydration of the upper tropo-
 2024 sphere and lower stratosphere by subvisible cirrus clouds near the tropical tropopause,
 2025 *Geophys. Res. Lett.*, *23*(8), 825–828.

- 2026 Jewett, B., M. Ramamurthy, and R. Rauber (2003), Origin, evolution, and finescale struc-
2027 ture of the St. Valentine's Day mesoscale gravity wave observed during STORM-FEST.
2028 Part III: Gravity wave genesis and the role of evaporation, *Mon. Weath. Rev.*, *131*(4),
2029 617–633.
- 2030 Jin, Y. (1997), *A numerical model study of the role of mesoscale gravity waves in rainband*
2031 *dynamics in the central United States during STORM-FEST (Ph.D. Dissertation)*, 318
2032 pp. pp., North Carolina State University.
- 2033 Joly, A., et al. (1997), The Fronts and Atlantic Stormtracks Experiment (FASTEX):
2034 scientific objectives and experimental design, *Bull. Amer. Meteorol. Soc.*, *78*(9), 1917–
2035 1940.
- 2036 Juckes, M. (1994), Quasi-geostrophic dynamics of the tropopause, *J. Atmos. Sci.*, *51*,
2037 2756–2768.
- 2038 Kalashnik, M. V. (1998), Forming of frontal zones during geostrophic adjustment in a
2039 continuously stratified fluid, *Izvetiya, Atmospheric and Oceanic Physics*, *34*(6), 785–
2040 792.
- 2041 Kalashnik, M. V. (2000), Geostrophic adjustment and frontogenesis in a continuously
2042 stratified fluid, *Izvetiya, Atmospheric and Oceanic Physics*, *36*(3), 386–395.
- 2043 Kalnay, E. (2003), *Atmospheric modeling, data assimilation and predictability*, 341pp pp.,
2044 Cambridge University Press.
- 2045 Kaplan, M., S. Koch, Y.-L. Lin, R. Weglarz, and R. Rozumalski (1997), Numerical Sim-
2046 ulations of a Gravity Wave Event over CCOPE. Part I: The Role of Geostrophic Ad-
2047 justment in Mesoscale Jetlet Formation, *Mon. Weath. Rev.*, *125*, 1185–1211.

- 2048 Kim, S.-Y., and H.-Y. Chun (2011), Statistics and Possible Sources of Aviation Turbulence
 2049 over South Korea, *J. App. Meteor. Clim.*, *50*, 311–324.
- 2050 Kim, Y.-J., S. Eckermann, and H.-Y. Chun (2003), An overview of the past, present
 2051 and future of gravity-wave drag parametrization for numerical climate and weather
 2052 prediction models, *Atmosphere-Ocean*, *41*, 65–98.
- 2053 Klein, P., and S. L. Smith (2001), Horizontal dispersion of near-inertial oscillations in a
 2054 turbulent mesoscale eddy field, *J. Mar. Res.*, *59*, 697–723.
- 2055 Klein, P., B. Hua, G. Lapeyre, X. Capet, S. LeGentil, and H. Sasaki (2008), Upper ocean
 2056 turbulence from high 3D resolution simulations, *J. Phys. Oceanogr.*, *38*, 1748.
- 2057 Knox, J., D. McCann, and P. Williams (2008), Application of the lighthill-ford theory of
 2058 spontaneous imbalance to Clear-Air Turbulence forecasting, *J. Atmos. Sci.*, *65*, 3292–
 2059 3304.
- 2060 Knox, J., D. McCann, and P. Williams (2009), Reply, *J. Atmos. Sci.*, *66*, 2511–2516.
- 2061 Knupp, K. (2006), Observational Analysis of a Gust Front to Bore to Solitary Wave
 2062 Transition within an Evolving Nocturnal Boundary Layer, *J. Atmos. Sci.*, *63*(8), 2016–
 2063 2035.
- 2064 Koch, S., and R. Golus (1988), A mesoscale gravity-wave event observed during CCOPE.
 2065 1. Multiscale statistical analysis of wave characteristics, *Mon. Weath. Rev.*, *116*(12),
 2066 2527–2544.
- 2067 Koch, S., and C. O’Handley (1997), Operational Forecasting and Detection of Mesoscale
 2068 Gravity Waves, *Wea. Forecasting*, *12*, 253–281.
- 2069 Koch, S., and S. Saleeby (2001), An automated system for the analysis of gravity waves
 2070 and other mesoscale phenomena, *Weather and Forecasting*, *16*, 661–679.

- 2071 Koch, S., R. Golus, and P. Dorian (1988), A mesoscale gravity wave event observed
2072 during CCOPE. Part II: Interactions between mesoscale convective systems and the
2073 antecedent waves, *Mon. Weath. Rev.*, *116*, 2545–2569.
- 2074 Koch, S., F. Einaudi, P. Dorian, S. Lang, and G. Heymsfield (1993), A Mesoscale Gravity-
2075 Wave Event Observed during CCOPE. Part IV: Stability Analysis and Doppler-derived
2076 Wave Vertical Structure, *Mon. Weath. Rev.*, *121*, 2483–2510.
- 2077 Koch, S., F. Zhang, M. Kaplan, Y. Lin, R. Weglarz, and C. Trexler (2001), Numerical
2078 simulations of a gravity wave event over CCOPE. Part III: The role of a mountain-plains
2079 solenoid in the generation of the second wave episode, *Mon. Weath. Rev.*, *129*(5), 909–
2080 933.
- 2081 Koch, S., et al. (2005), Turbulence and gravity waves within an upper-level front, *J.*
2082 *Atmos. Sci.*, *62*, 3885–3908.
- 2083 Koch, S. E., and P. B. Dorian (1988), A mesoscale gravity wave event observed during
2084 CCOPE. Part III: wave environment and possible source mechanisms, *Mon. Wea. Rev.*,
2085 *116*, 2570–2591.
- 2086 Koppel, L., L. Bosart, and D. Keyser (2000), A 25-yr Climatology of Large-Amplitude
2087 Hourly Surface Pressure Changes over the Conterminous United States, *Mon. Weath.*
2088 *Rev.*, *128*(1), 51–68.
- 2089 Kuo, A. C., and L. M. Polvani (1997), Time-dependent fully nonlinear geostrophic ad-
2090 justment, *J. Phys. Oceanogr.*, *27*, 1614–1634.
- 2091 Kuo, A. C., and L. M. Polvani (2000), Nonlinear geostrophic adjustment, cy-
2092 clone/anticyclone asymmetry, and potential vorticity rearrangement, *Phys. Fluids*,
2093 *12*(5), 1087–1100.

- 2094 Kuo, H. (1997), A new perspective of geostrophic adjustment, *Dyn. Atmos. Ocean*, *27*,
 2095 413–437.
- 2096 Kushner, P., M. McIntyre, and T. Shepherd (1998), Coupled kelvin-wave and mirage wave
 2097 instabilities in semi-geostrophic dynamics, *J. Phys. Oceanogr.*, *28*, 513–518.
- 2098 Lalas, D., and F. Einaudi (1976), On the characteristics of waves generated by shear
 2099 layers, *J. Atmos. Sci.*, *33*, 1248–1259.
- 2100 Lalas, D., F. Einaudi, and D. Fua (1976), The destabilizing effect of the ground on Kelvin-
 2101 Helmholtz waves in the atmosphere, *J. Atmos. Sci.*, *33*, 59–69.
- 2102 Lane, T., J. Doyle, R. Plougonven, R. Sharman, and M. Shapiro (2004), Numerical mod-
 2103 eling of gravity waves and shearing instabilities above an observed jet, *J. Atmos. Sci.*,
 2104 *61*, 2692–2706.
- 2105 Lapeyre, G., B. Hua, and P. Klein (1999), Does the tracer gradient vector align with the
 2106 strain eigenvectors in 2d turbulence?, *Phys. Fluids*, *11*, 3729–3737.
- 2107 Lawrence, G., F. Browand, and L. Redekopp (1991), The stability of a sheared density
 2108 interface, *Phys. Fluids*, *3*, 2360–2370.
- 2109 LeDizès, S., and P. Billant (2009), Radiative instability in stratified vortices, *Phys. Fluids*,
 2110 *21*(096602), doi:10.1063/1.3241,995.
- 2111 Leith, C. (1980), Nonlinear normal mode initialization and quasi-geostrophic theory, *J.*
 2112 *Atmos. Sci.*, *37*, 958–968.
- 2113 Ley, B., and W. Peltier (1978), Wave generation and frontal collapse, *J. Atmos. Sci.*,
 2114 *35*(1), 3–17.
- 2115 Li, Q., J. Xu, J. Yue, W. Yuan, and X. Liu (2011), Statistical characteristics of gravity
 2116 wave activities observed by an OH airglow imager at Xinglong, in northern China, *ANN*.

- 2117 *Geophys.*, 29(doi:10.5194/angeo-29-1401-2011), 1401–1410.
- 2118 Lighthill, J. M. (1952), On sound generated aerodynamically, I. General theory, *Proc.*
2119 *Roy. Soc. London*, 211(A), 564–587.
- 2120 Lighthill, J. M. (1978), *Waves in Fluids*, Cambridge University Press.
- 2121 Lin, Y., and F. Zhang (2008), Tracking gravity waves in baroclinic jet-front systems, *J.*
2122 *Atmos. Sci.*, 65, 2402–2415.
- 2123 Lin, Y.-L., and R. Goff (1988), A case study of solitary wave in the atmosphere originating
2124 near a region of deep convection, *J. Atmos. Sci.*, 45, 194–205.
- 2125 Lindzen, R., and M. Fox-Rabinowitz (1989), Consistent vertical and horizontal resolution,
2126 *Mon. Weath. Rev.*, 117, 2575–2583.
- 2127 Lindzen, R., and K.-K. Tung (1976), Banded convective activity and ducted gravity waves,
2128 *Mon. Weath. Rev.*, 104, 1602–1617.
- 2129 Liu, A., and G. Swenson (2003), A modeling study of O₂ and OH airglow perturbations
2130 induced by atmospheric gravity waves, *J. Geophys. Res.*, 108(D4), 4151.
- 2131 Lorenz, E. (1980), Attractor sets and quasi-geostrophic equilibrium, *J. Atmos. Sci.*, 37,
2132 1685–1699.
- 2133 Lorenz, E. (1986), On the existence of a slow manifold, *J. Atmos. Sci.*, 43, 1547–1557.
- 2134 Lorenz, E., and V. Krishnamurty (1987), On the nonexistence of a slow manifold, *J.*
2135 *Atmos. Sci.*, 44, 2940–2950.
- 2136 Lott, F. (1997), The transient emission of propagating gravity waves by a stably stratified
2137 shear layer, *Q.J.R. Meteorol. Soc.*, 123, 1603–1619.
- 2138 Lott, F., H. Kelder, and H. Teitelbaum (1992), A transition from Kelvin-Helmholtz in-
2139 stabilities to propagating wave instabilities, *Phys. Fluids*, 4(9), 1990–1997.

- 2140 Lott, F., R. Plougonven, and J. Vanneste (2010), Gravity waves generated by sheared
 2141 potential vorticity anomalies, *J. Atmos. Sci.*, *67*(DOI:10.1175/2009JAS3134.1), 157–
 2142 170.
- 2143 Lott, F., L. Guez, and P. Maury (2012a), A stochastic parameterization of non-orographic
 2144 gravity waves: Formalism and impact on the equatorial stratosphere, *Geophys. Res.*
 2145 *Lett.*, *39*(L06807), 10.1029/2012GL051,001.
- 2146 Lott, F., R. Plougonven, and J. Vanneste (2012b), Gravity waves generated by sheared
 2147 three-dimensional potential vorticity anomalies, *J. Atmos. Sci.*
- 2148 Lovegrove, A., P. Read, and C. Richards (2000), Generation of inertia-gravity waves in a
 2149 baroclinically unstable fluid, *Q.J.R. Meteorol. Soc.*, *126*, 3233–3254.
- 2150 Luo, Z., and D. Fritts (1993), Gravity wave excitation by geostrophic adjustment of the
 2151 jet stream. Part II: Three dimensional forcing, *J. Atmos. Sci.*, *50*(1), 104–115.
- 2152 Lynch, P. (2002), *Geometric Methods and Models, Vol. II, Large-Scale Atmosphere-Ocean*
 2153 *Dynamics*, chap. The swinging spring: A simple model for atmospheric balance., pp.
 2154 64–108, Cambridge University Press.
- 2155 Machenhauer, B. (1977), On the dynamics of gravity oscillations in a shallow water model,
 2156 with applications to normal mode initialization, *Contrib. Atmos. Phys.*, *50*, 253–271.
- 2157 MacKay, R. (2004), *Energy localisation and transfer*, chap. Slow manifolds, pp. 149–192,
 2158 World Sci.
- 2159 Mahalov, A., M. Moustou, B. Nicolaenko, and K. Tse (2007), Computational studies of
 2160 inertia-gravity waves radiated from upper tropospheric jets, *Theoretical and Computa-*
 2161 *tional Fluid Dynamics*, *21*(6), 399–422.

- 2162 Mamatsashvili, G., V. Avsarkisov, G. Chagelishvili, R. Chanishvili, and M. Kalashnik
2163 (2010), Transient Dynamics of Nonsymmetric Perturbations versus Symmetric Instabil-
2164 ity in Baroclinic Zonal Shear Flows, *J. Atmos. Sci.*, *67*(9), 2972–2989.
- 2165 Mann, G., K. S. Carslaw, M. P. Chipperfield, and S. Davies (2005), Large nitric acid trihy-
2166 date particles and denitrification caused by mountain waves in the Arctic stratosphere,
2167 *J. Geophys. Res.*, *110*(D08202), doi:10.1029/2004JD005271.
- 2168 Mastrantonio, G., F. Einaudi, D. Fua, and D. Lalas (1976), Generation of gravity waves
2169 by jet streams in the atmosphere, *J. Atmos. Sci.*, *33*, 1730–1738.
- 2170 Matsumoto, S. (1961), A note on geostrophic adjustment and gravity waves in the atmo-
2171 sphere, *J. Meteor. Soc. Japan*, *39*, 18–28.
- 2172 McDonald, A., S. George, and R. Woollands (2009), Can gravity waves significantly im-
2173 pact PSC occurrence in the Antarctic?, *Atmos. Chem. Phys.*, *9*, 8825–8840.
- 2174 McIntyre, M. (2001), Global effects of gravity waves in the middle atmosphere: a theo-
2175 retical perspective, *Adv. Space Res.*, *27*(10), 1723–1736.
- 2176 McIntyre, M. (2009), Spontaneous imbalance and hybrid vortex-gravity wave structures,
2177 *J. Atmos. Sci.*, *66*, 1315–1326.
- 2178 McIntyre, M., and M. Weissman (1978), On radiating instabilities and resonant overreflec-
2179 tion, *J. Atmos. Sci.*, *35*, 1190–1196.
- 2180 McWilliams, J. (2003), *Nonlinear Processes in Geophysical Fluid Dynamics*, chap. Diag-
2181 nostic force balance and its limits, pp. 287–304, Kluwer.
- 2182 McWilliams, J., and I. Yavneh (1998), Fluctuation growth and instability associated with
2183 a singularity of the balance equations, *Phys. Fluids*, *10*(10), 2587–2596.

- 2184 McWilliams, J., M. Molemaker, and I. Yavneh (2001), From stirring to mixing of mo-
 2185 mentum: cascades from balanced flows to dissipation in the oceanic interior, in *Aha*
 2186 *Huliko'a Proceedings*, pp. 59–66.
- 2187 McWilliams, J. C., and P. R. Gent (1980), Intermediate models of planetary circulations
 2188 in the atmosphere and ocean, *J. Atmos. Sci.*, *37*(8), 1657–1678.
- 2189 Miller, J. (1948), On the concept of frontogenesis, *J. Meteorology*, *5*, 169–171.
- 2190 Molemaker, M., J. McWilliams, and I. Yavneh (2001), Instability and equilibration of
 2191 centrifugally stable stratified Taylor-Couette flow, *Phys. Rev. Lett.*, *86*(23), 5270–5273.
- 2192 Molemaker, M., J. McWilliams, and I. Yavneh (2005), Baroclinic instability and loss of
 2193 balance, *J. Phys. Oceanogr.*, *35*, 1505–1517.
- 2194 Molemaker, M., J. McWilliams, and X. Capet (2010), Balanced and unbalanced routes to
 2195 dissipation in an equilibrated Eady flow, *J. Fluid Mech.*, *654*, 35–63.
- 2196 Moustou, M., H. Teitelbaum, P. van Velthoven, and H. Kelder (1999), Analysis of
 2197 gravity waves during the POLINAT experiment and some consequences for stratosphere-
 2198 troposphere exchange, *J. Atmos. Sci.*, *56*, 1019–1030.
- 2199 Müller, P., J. McWilliams, and M. Molemaker (2005), *Marine Turbulence: Theories,*
 2200 *Observations and Models*, chap. Routes to dissipation in the ocean: the 2d/3d turbulence
 2201 conundrum., pp. 397–405, Cambridge University Press, Cambridge.
- 2202 Muraki, D., and C. Snyder (2007), Vortex dipoles for surface quasigeostrophic models, *J.*
 2203 *Atmos. Sci.*, *64*, 2961–2967.
- 2204 Murayama, Y., T. Tsuda, R. Wilson, H. Nakane, S. Hayashida, N. Sugimoto, I. Mat-
 2205 sui, and Y. Sasano (1994), Gravity wave activity in the upper stratosphere and lower
 2206 mesosphere observed with the Rayleigh lidar at Tsukuba, Japan, *Geophys. Res. Lett.*,

2207 21(14), 1539–1542.

2208 Nakamura, N. (1988), Scale selection of baroclinic instability - effects of stratification and
2209 nongeostrophy, *J. Atmos. Sci.*, 45(21), 3253–3267.

2210 Nastrom, G., and D. Fritts (1992), Sources of mesoscale variability of gravity waves. part
2211 i: topographic excitation, *J. Atmos. Sci.*, 49(2), 101–109.

2212 Nicholls, M., R. Pielke, and W. Cotton (1991), Thermally forced gravity waves in an
2213 atmosphere at rest, *J. Atmos. Sci.*, 48(16), 1869–1884.

2214 Obukhov, A. (1949), On the question of geostrophic wind (in Russian), *Izv. Akad. Nauk.*
2215 *SSSR Ser. Geograf.-Geofiz.*, 13(4), 281–306.

2216 Olafsdottir, E., A. O. Daalhuis, and J. Vanneste (2008), Inertia-gravity-wave generation
2217 by a sheared vortex, *J. Fluid Mech.*, 569, 169–189.

2218 O’Sullivan, D., and T. Dunkerton (1995), Generation of inertia-gravity waves in a simu-
2219 lated life cycle of baroclinic instability, *J. Atmos. Sci.*, 52(21), 3695–3716.

2220 Ou, H. W. (1984), Geostrophic adjustment: a mechanism for frontogenesis, *J. Phys.*
2221 *Oceanogr.*, 14, 994–1000.

2222 Paegle, J. (1978), The transient mass-flow adjustment of heated atmospheric circulations,
2223 *J. Atmos. Sci.*, 35, 1678–1688.

2224 Palmer, T. (2001), A nonlinear dynamical perspective on model error: A proposal for
2225 non-local stochastic-dynamic parametrization in weather and climate prediction models,
2226 *Q.J.R. Meteorol. Soc.*, 127(572), 279–304.

2227 Pan, L., et al. (2010), The stratosphere-troposphere analyses of regional transport 2008
2228 experiment, *Bull. Amer. Meteorol. Soc.*, 91(DOI:10.1175/2009BAMS2865.1), 327–342.

- 2229 Parsons, D., and P. Hobbs (1983), The mesoscale and microscale structure and organi-
 2230 zation of clouds and precipitation in mid-latitude cyclones. 11. Comparisons between
 2231 observational and theoretical aspects of rainbands, *J. Atmos. Sci.*, *40*(10), 2377–2397.
- 2232 Pavelin, E., and J. Whiteway (2002), Gravity wave interactions around the jet stream,
 2233 *Geophys. Res. Lett.*, *29*(21), 2024–2027.
- 2234 Pavelin, E., J. Whiteway, and G. Vaughan (2001), Observation of gravity wave generation
 2235 and breaking in the lowermost stratosphere, *J. Geophys. Res.*, *106*(D6), 5173–5179.
- 2236 Pawson, S., et al. (2000), The GCM-Reality Intercomparison Project for SPARC (GRIPS):
 2237 scientific issues and initial results, *Bull. Amer. Meteor. Soc.*, *81*, 781–796.
- 2238 Peters, D., P. Hoffmann, and M. Alpers (2003), On the appearance of inertia-gravity
 2239 waves on the North-Easterly side of an anticyclone, *Meteo. Zeitschrift*, *12*(1), 25–35.
- 2240 Pierce, R., and T. Fairlie (1993), Chaotic advection in the stratosphere: implications
 2241 for the dispersal of chemically perturbed air from the polar vortex, *J. Geophys. Res.*,
 2242 *98*(D10), 18,589–18,595.
- 2243 Pierce, R., and W. Grant (1998), Seasonal evolution of Rossby and gravity wave induced
 2244 laminae in ozonesonde data obtained from Wallops Island, Virginia, *Geophys. Res. Lett.*,
 2245 *25*, 1859–1862.
- 2246 Plougonven, R., and C. Snyder (2005), Gravity waves excited by jets: propagation versus
 2247 generation, *Geoph. Res. Lett.*, *32*(L18892), doi:10.1029/2005GL023,730.
- 2248 Plougonven, R., and C. Snyder (2007), Inertia-gravity waves spontaneously generated by
 2249 jets and fronts. Part I: Different baroclinic life cycles, *J. Atmos. Sci.*, *64*, 2502–2520.
- 2250 Plougonven, R., and H. Teitelbaum (2003), Comparison of a large-scale inertia-gravity
 2251 wave as seen in the ECMWF and from radiosondes, *Geophys. Res. Lett.*, *30*(18), 1954.

- 2252 Plougonven, R., and V. Zeitlin (2002), Internal gravity wave emission from a pancake
2253 vortex: an example of wave-vortex interaction in strongly stratified flows, *Phys. of*
2254 *Fluids*, *14*(3), 1259–1268.
- 2255 Plougonven, R., and V. Zeitlin (2005), Lagrangian approach to the geostrophic adjustment
2256 of frontal anomalies in a stratified fluid, *Geophys. Astr. Fluid Dyn.*, *99*(2), 101–135.
- 2257 Plougonven, R., and F. Zhang (2007), On the forcing of inertia-gravity waves by synoptic-
2258 scale flows, *J. Atmos. Sci.*, *64*, 1737–1742.
- 2259 Plougonven, R., H. Teitelbaum, and V. Zeitlin (2003), Inertia-gravity wave generation by
2260 the tropospheric mid-latitude jet as given by the fastex radiosoundings, *J. Geophys.*
2261 *Res.*, *108*(D21), 4686.
- 2262 Plougonven, R., D. Muraki, and C. Snyder (2005), A baroclinic instability that couples
2263 balanced motions and gravity waves, *J. Atmos. Sci.*, *62*, 1545–1559.
- 2264 Plougonven, R., A. Hertzog, and H. Teitelbaum (2008), Observations and simulations of
2265 a large-amplitude mountain wave breaking above the Antarctic Peninsula, *J. Geophys.*
2266 *Res.*, *113*(D16113), doi:10.1029/2007JD009,739.
- 2267 Plougonven, R., C. Snyder, and F. Zhang (2009), Comments on 'application of the
2268 Lighthill-Ford theory of spontaneous imbalance to clear-air turbulence forecasting',
2269 *J. Atmos. Sci.*, *66*, 2506–2510.
- 2270 Plougonven, R., A. Arzac, A. Hertzog, L. Guez, and F. Vial (2010), Mesoscale simulations
2271 of the gravity wave field above antarctica during vorcore, *Quart. J. Roy. Meteorolog.*
2272 *Soc.*, *136*(650), 1371–1377.
- 2273 Plougonven, R., A. Hertzog, and L. Guez (2012), Gravity waves over Antarctica and
2274 the Southern Ocean: consistent momentum fluxes in mesoscale simulations and strato-

- 2275 spheric balloon observations, *in preparation for Quart. J. Roy. Meteorolog. Soc.*
- 2276 Pokrandt, P., G. Tripoli, and D. Houghton (1996), Processes leading to the formation of
 2277 mesoscale waves in the Midwest cyclone of 15 December 1987, *Mon. Weath. Rev.*, *124*,
 2278 2726–2752.
- 2279 Polzin, K. (2008), Mesoscale EddyInternal Wave Coupling. Part I: Symmetry, Wave Cap-
 2280 ture, and Results from the Mid-Ocean Dynamics Experiment, *J. Phys. Oceanogr.*, *38*,
 2281 2556–2574.
- 2282 Polzin, K. (2010), Mesoscale EddyInternal Wave Coupling. Part II: Energetics and Results
 2283 from PolyMode, *J. Phys. Oceanogr.*, *340*, 789–801.
- 2284 Potter, B., and J. Holton (1995), The role of monsoon convection in the dehydration of
 2285 the lower tropical stratosphere, *J. Atmos. Sci.*, *52*(8), 1034–1050.
- 2286 Powers, J. (1997), Numerical model simulation of a mesoscale gravity-wave event: sensi-
 2287 tivity tests and spectral analyses, *Mon. Weath. Rev.*, *125*, 1838–1869.
- 2288 Powers, J., and R. Reed (1993), Numerical simulation of the large-amplitude mesoscale
 2289 gravity wave event of 15 December 1987 in the Central United States, *Mon. Weath.*
 2290 *Rev.*, *121*, 2285–2308.
- 2291 Preusse, P., S. Eckermann, and M. Ern (2008), Transparency of the atmosphere to short
 2292 horizontal wavelength gravity waves, *J. Geophys. Res.*, *113*(10.1029/2007JD009682),
 2293 D24,104.
- 2294 Queney, P. (1948), The problem of air flow over mountains: A summary of theoretical
 2295 studies, *Bull. Am. Meteorol. Soc.*, *29*, 16–26.
- 2296 Ralph, F., M. Crochet, and S. Venkateswaran (1993), Observations of a mesoscale ducted
 2297 gravity wave, *J. Atmos. Sci.*, *50*(19), 3277–3291.

- 2298 Ralph, F., P. Neiman, and T. Keller (1999), Deep-tropospheric gravity waves created by
2299 leeside cold fronts, *J. Atmos. Sci.*, *56*, 2986–3009.
- 2300 Ramamurthy, M., R. Rauber, B. Collins, and N. Malhotra (1993), A comparative study
2301 of large amplitude gravity wave events, *Mon. Weath. Rev.*, *121*(11), 2951–2974.
- 2302 Rauber, R., M. Yang, M. Ramamurthy, and B. Jewett (2001), Origin, Evolution, and
2303 Finescale Structure of the St. Valentines Day Mesoscale Gravity Wave Observed during
2304 STORM-FEST. Part I: Origin and Evolution, *Mon. Weath. Rev.*, *129*(2), 198–217.
- 2305 Read, P. (1992), Applications of singular systems analysis to baroclinic chaos, *Physica D*,
2306 *58*, 455–468.
- 2307 Reeder, M. J., and M. Griffiths (1996), Stratospheric inertia-gravity waves generated in
2308 a numerical model of frontogenesis. Part II: Wave sources, generation mechanisms and
2309 momentum fluxes., *Q.J.R. Meteorol. Soc.*, *122*, 1175–1195.
- 2310 Reznik, G., V. Zeitlin, and M. B. Jelloul (2001), Nonlinear theory of geostrophic adjust-
2311 ment. Part 1. Rotating shallow-water model, *J. Fluid Mech.*, *445*, 93–120.
- 2312 Rhines, P., E. Lindahl, and A. Mendez (2006), Optical altimetry: A new method for
2313 observing rotating fluids with application to Rossby waves on a polar beta-plane., *J.*
2314 *Fluid Mech.*, *572*, 389–412.
- 2315 Richiardone, R., and M. Manfrin (2003), A rain episode related to a mesoscale gravity
2316 wave, *Bull. Amer. Meteorol. Soc.*, *84*(10.1175/BAMS-84-11-1494), 1494–1498.
- 2317 Richter, J., M. Geller, R. Garcia, H.-L. Liu, and F. Zhang (2007), Report on the gravity
2318 wave retreat, *SPARC Newsletter*, *28*, 26–27.
- 2319 Richter, J., F. Sassi, and R. Garcia (2010), Toward a physically based grav-
2320 ity wave source parameterization in a general circulation model, *J. Atmos. Sci.*,

- 2321 67([doi:10.1175/2009JAS3112.1](https://doi.org/10.1175/2009JAS3112.1)), 136–156.
- 2322 Riedinger, X., S. LeDizès, and P. Meunier (2010a), Viscous stability properties of a Lamb-
 2323 Oseen vortex in a stratified fluid, *J. Fluid Mech.*, *655*, 255–278.
- 2324 Riedinger, X., P. Meunier, and S. LeDizès (2010b), Instability of a columnar vortex in a
 2325 stratified fluid, *Exp. Fluids*, *49*, 673–681.
- 2326 Riedinger, X., S. LeDizès, and P. Meunier (2011), Radiative instability of the flow around
 2327 a rotating cylinder in a stratified fluid, *J. Fluid Mech.*, *672*, 130–146.
- 2328 Rind, D., R. Suozzo, N. Balachandran, A. Lacis, and G. Russell (1988), The GISS global
 2329 climate-middle atmosphere model. Part I: model structure and climatology, *J. Atmos.*
 2330 *Sci.*, *45*(3), 329–370.
- 2331 Rossby, C. (1938), On the mutual adjustment of pressure and velocity distributions in
 2332 certain simple current systems II, *J. Mar. Res.*, *1*, 239–263.
- 2333 Sakai, S. (1989), Rossby-kelvin instability: a new type of ageostrophic instability caused
 2334 by a resonance between rossby waves and gravity waves, *J. Fluid Mech.*, *202*, 149–176.
- 2335 Sato, K. (1994), A statistical study of the structure, saturation and sources of inertio-
 2336 gravity waves in the lower stratosphere observed with the MU radar, *J. Atmos. Terr.*
 2337 *Phys.*, *56*(6), 755–774.
- 2338 Sato, K., and M. Yoshiki (2008), Gravity wave generation around the polar vortex in the
 2339 stratosphere revealed by 3-hourly radiosonde observations at Syowa Station, *J. Atmos.*
 2340 *Sci.*, *65*, 3719–3735.
- 2341 Sato, K., S. Watanabe, Y. Kawatani, Y. Tomikawa, K. Miyazaki, and M. Takayashi
 2342 (2009), On the origins of mesospheric gravity waves, *Geophys. Res. Lett.*, *36*(L19801),
 2343 [doi:10.1029/2009GL039,908](https://doi.org/10.1029/2009GL039,908).

- 2344 Schechter, D. (2008), The spontaneous imbalance of an atmospheric vortex at high Rossby
2345 number, *J. Atmos. Sci.*, *65*, 2498–2521.
- 2346 Schechter, D., and M. Montgomery (2006), Conditions that inhibit the spontaneous radi-
2347 ation of spiral inertia-gravity waves from an intense mesoscale cyclone, *J. Atmos. Sci.*,
2348 *63*, 435–456.
- 2349 Schmidt, J., and W. Cotton (1990), Interactions between upper and lower tropospheric
2350 gravity waves on squall line structure and maintenance, *J. Atmos. Sci.*, *47*, 1205–1222.
- 2351 Schneider, R. (1990), Large-Amplitude Mesoscale Wave Disturbances Within the Intense
2352 Midwest Extratropical Cyclone of 15 December 1987, *Weather and forecasting*, *5*, 533–
2353 558.
- 2354 Schroeder, S., P. Preusse, M. Ern, and M. Riese (2009), Gravity waves resolved
2355 in ECMWF and measured by SABER, *Geophys. Res. Lett.*, *36*(L10805), doi:
2356 10.1029/2008GL037,054.
- 2357 Schubert, W., J. Hack, P. S. Dias, and S. Fulton (1980), Geostrophic adjustment in an
2358 axisymmetric vortex, *J. Atmos. Sci.*, *37*, 1464–1484.
- 2359 Scinocca, J., and R. Ford (2000), The nonlinear forcing of large-scale internal gravity
2360 waves by stratified shear instability, *J. Atmos. Sci.*, *57*, 653–672.
- 2361 Scolan, H., J.-B. Flor, and J. Gula (2011), Frontal instabilities and waves in a differentially
2362 rotating fluid, *J. Fluid Mech.*, *685*, 532–542.
- 2363 Sharman, R., C. Tebaldi, G. Wiener, and J. Wolff (2006), An integrated approach to mid-
2364 and upper-level turbulence forecasting, *Weather and Forecasting*, pp. 268–287.
- 2365 Sharman, R., S. Trier, T. Lane, and J. Doyle (2012), Source and dynamics of turbu-
2366 lence in the upper troposphere and lower stratosphere: a review, *Geophys. Res. Lett.*,

- 2367 39(L12803), doi:10.1029/2012GL051,996.
- 2368 Shibata, T., K. Sato, H. Kobayashi, M. Yabuki, and M. Shiobara (2003), Antarctic polar
 2369 stratospheric clouds under temperature perturbations by nonorographic inertia-gravity
 2370 waves observed by micropulse lidar at Syowa Station, *J. Geophys. Res.*, 108(D3), 4105.
- 2371 Shutts, G., and S. Vosper (2011), Stratospheric gravity waves revealed in NWP forecast
 2372 models, *Q.J.R. Meteorol. Soc.*, 137(655), 303–317.
- 2373 Simmons, A., and B. Hoskins (1978), The life cycles of some nonlinear baroclinic waves,
 2374 *J. Atmos. Sci.*, 35, 414–432.
- 2375 Snyder, C. (1995), Stability of steady fronts with uniform potential vorticity, *J. Atmos.*
 2376 *Sci.*, 52(6), 724–736.
- 2377 Snyder, C., W. Skamarock, and R. Rotunno (1993), Frontal dynamics near and following
 2378 frontal collapse, *J. Atmos. Sci.*, 50(18), 3194–3211.
- 2379 Snyder, C., D. Muraki, R. Plougonven, and F. Zhang (2007), Inertia-gravity waves gen-
 2380 erated within a dipole vortex, *J. Atmos. Sci.*, 64, 4417–4431.
- 2381 Snyder, C., R. Plougonven, and D. Muraki (2009), Forced linear inertia-gravity waves on
 2382 a basic-state dipole vortex, *J. Atmos. Sci.*, 66(11), 3464–3478.
- 2383 Song, I.-S., and H.-Y. Chun (2005), Momentum flux spectrum of convectively forced
 2384 internal gravity waves and its application to gravity wave drag parameterization. Part
 2385 I: Theory, *J. Atmos. Sci.*, 62, 107–124.
- 2386 Spiga, A., H. Teitelbaum, and V. Zeitlin (2008), Identification and separation of the
 2387 sources of inertia-gravity waves in the Andes Cordillera region, *Ann. Geophys.*, 26,
 2388 2551–2568.

- 2389 Stone, P. (1970), On non-geostrophic baroclinic instability: Part II, *J. Atmos. Sci.*, *27*,
2390 721–726.
- 2391 Sugimoto, N., and K. Ishii (2012), Spontaneous gravity wave radiation in a shallow water
2392 system on a rotating sphere, *J. Meteor. Soc. Jap.*, *90*, 101–125.
- 2393 Sugimoto, N., K. Ishioka, and K. Ishii (2008), Parameter sweep experiments on sponta-
2394 neous gravity wave radiation from unsteady rotational flow in an f-plane shallow water
2395 system, *J. Atmos. Sci.*, *65*, 235–249.
- 2396 Sutherland, B. (2006), Rayleigh wave - internal wave coupling and internal wave genera-
2397 tion above a model jet stream, *J. Atmos. Sci.*, *63*, 1042–1055.
- 2398 Sutherland, B., and W. Peltier (1995), Internal gravity wave emission into the middle
2399 atmosphere from a model tropospheric jet, *J. Atmos. Sci.*, *52*, 3214–3235.
- 2400 Sutherland, B., C. Caulfield, and W. Peltier (1994), Internal gravity wave generation and
2401 hydrodynamic instability, *J. Atmos. Sci.*, *51*, 3261–3280.
- 2402 Sutyrin, G. (2007), Ageostrophic instabilities in a horizontally uniform baroclinic flow
2403 along a slope, *J. Fluid Mech.*, *588*(DOI: 10.1017/S0022112007006829), 463–473.
- 2404 Sutyrin, G. (2008), Lack of balance in continuously stratified rotating flows, *J. Fluid*
2405 *Mech.*, *615*(DOI: 10.1017/S0022112008004059), 93–100.
- 2406 Taylor, M., and M. Bishop (1995), All-sky measurements of short-period waves imaged in
2407 the OI (557.7 nm), Na(589.2 nm) and near-infrared OH and O₂(0,1) nightglow emissions
2408 during the Aloha-93 campaign, *Geophys. Res. Lett.*, *22*(20), 2833–2836.
- 2409 Teitelbaum, H., M. Moustouai, J. Ovarlez, and H. Kelder (1996), The role of atmospheric
2410 waves in the laminated structure of ozone profiles, *Tellus*, *48A*, 442–455.

- 2411 Tepper, M. (1951), On the dessication of a cloud bank by a propgating pressure wave,
 2412 *Mon. Weath. Rev.*, *79*, 61–70.
- 2413 Thomas, L., R. Worthington, and A. McDonald (1999), Inertia-gravity waves in the tro-
 2414 posphere and lower stratosphere associated with a jet stream exit region, *Ann. Geo-*
 2415 *physicae*, *17*, 115–121.
- 2416 Thorncroft, C., B. Hoskins, and M. McIntyre (1993), Two paradigms of baroclinic-wave
 2417 life-cycle behaviour, *Q.J.R. Meteorol. Soc.*, *119*, 17–55.
- 2418 Tokioka, T. (1970), Non-geostrophic and non-hydrostatic stability of a baroclinic fluid, *J.*
 2419 *Meteorol. Soc. Japan*, *48*, 503–520.
- 2420 Tomikawa, Y., K. Sato, K. Kita, M. Fujiwara, M. Yamamori, and T. Sano (2002), Forma-
 2421 tion of an ozone lamina due to differential advection revealed by intensive observations,
 2422 *J. Geophys. Res.*, *107*(D10(4092)), 10.1029/2001JD000,386.
- 2423 Trexler, M., and S. Koch (2000), The Life Cycle of a Mesoscale Gravity Wave as Observed
 2424 by a Network of Doppler Wind Profilers, *Mon. Weath. Rev.*, *128*, 2423–2446.
- 2425 Trier, S., R. Sharman, and T. Lane (2012), Influences of moist convection on a cold season
 2426 outbreak of Clear-Air Turbulence (CAT), *in press for Mon. Wea. Rev.*
- 2427 Tse, K., A. Mahalov, B. Nicolaenko, and H. Fernando (2003), Quasi-equilibrium dynamics
 2428 of shear-stratified turbulence in a model tropospheric jet, *J. Fluid Mech.*, *496*, 73–103.
- 2429 Tuyl, A. V., and J. Young (1982), Numerical simulation of nonlinear jet streak adjustment,
 2430 *Mon. Wea. Rev.*, *110*, 2038–2054.
- 2431 Uccellini, L., and S. Koch (1987), The synoptic setting and possible energy sources for
 2432 mesoscale wave disturbances, *Mon. Wea. Rev.*, *115*, 721–729.

- 2433 Vallis, G. (1992), Mechanisms and parameterization of geostrophic adjustment and a
2434 variational approach to balanced flow, *J. Atmos. Sci.*, *49*, 1144–1160.
- 2435 Vallis, G. (2006), *Atmospheric and oceanic fluid dynamics*, 745 pp., Cambridge University
2436 Press.
- 2437 Vanneste, J. (2004), Inertia-gravity wave generation by balanced motion: revisiting the
2438 Lorenz-Krishnamurty model, *J. Atmos. Sci.*, *61*, 224–234.
- 2439 Vanneste, J. (2006), Wave radiation by balanced motion in a simple model, *SIAM J. Appl.*
2440 *Dynam. Syst.*, *5*, 783–807.
- 2441 Vanneste, J. (2008), Exponential smallness of inertia-gravity-wave generation at small
2442 rossby number, *J. Atmos. Sci.*, *65*, 1622–1637.
- 2443 Vanneste, J. (2013), Balance and spontaneous wave generation in geophysical flows, *Ann.*
2444 *Rev. Fluid Mech.*, *45*, 147–172.
- 2445 Vanneste, J., and I. Yavneh (2004), Exponentially small inertia-gravity waves and the
2446 breakdown of quasi-geostrophic balance, *J. Atmos. Sci.*, *61*, 211–223.
- 2447 Vanneste, J., and I. Yavneh (2007), Unbalanced instabilities of rapidly rotating stratified
2448 shear flows, *J. Fluid Mech.*, *584*, 373–396.
- 2449 Vaughan, G., and R. Worthington (2007), Inertia-gravity waves observed by the UK MST
2450 radar, *qjrms*, *133*(S2), 179–188.
- 2451 Vautard, R., and B. Legras (1986), Invariant manifolds, quasi-geostrophy and initializa-
2452 tion, *J. Atmos. Sci.*, *43*(4), 565–584.
- 2453 Vincent, R., A. Hertzog, G. Boccara, and F. Vial (2007), Quasi-Lagrangian superpressure
2454 balloon measurements of gravity-wave momentum fluxes in the polar stratosphere of
2455 both hemispheres, *Geophys. Res. Lett.*, *34*(L19804), doi:10.1029/2007GL031,072.

- 2456 Viudez, A. (2007), The origin of the stationary frontal wave packet spontaneously gener-
 2457 ated in rotating stratified vortex dipoles, *J. Fluid Mech.*, *593*, 359–383.
- 2458 Viudez, A. (2008), The stationary frontal wave packet spontaneously generated in
 2459 mesoscale dipoles, *jpo*, *38*, 243–256.
- 2460 Viúdez, A., and D. Dritschel (2003), An explicit potential vorticity conserving approach
 2461 to modelling nonlinear internal gravity waves, *J. Fluid Mech.*
- 2462 Viúdez, A., and D. Dritschel (2006), Spontaneous generation of inertia-gravity wave pack-
 2463 ets by geophysical balanced flows, *J. Fluid Mech.*, *553*, 107–117.
- 2464 Waite, M. L., and C. Snyder (2009), The mesoscale kinetic energy spectrum of a baroclinic
 2465 life cycle, *J. Atmos. Sci.*, *66*(4), 883–901.
- 2466 Waite, M. L., and C. Snyder (2012), Mesoscale energy spectra of moist baroclinic waves,
 2467 *submitted to J. Atm. Sci.*
- 2468 Walterscheid, R., J. Hecht, R. Vincent, I. Reid, J. Woithe, and M. Hickey (1999), Analysis
 2469 and interpretation of airglow and radar observations of quasi-monochromatic gravity
 2470 waves in the upper mesosphere and lower thermosphere over Adelaide, Australia (35S,
 2471 138E), *J. Atmos. Solar-Terr. Phys.*, *61*(6), 461–478.
- 2472 Wang, L., and M. Geller (2003), Morphology of gravity-wave energy as observed from 4
 2473 years (1998-2001) of high vertical resolution U.S. radiosonde data, *J. Geophys. Res.*,
 2474 *108*(doi:10.1029/2002JD002786), 4489.
- 2475 Wang, S. (2008), *Gravity Waves from Vortex Dipoles and Jets (Ph.D. Dissertation)*, Texas
 2476 A&M University.
- 2477 Wang, S., and F. Zhang (2006), Sensitivity of mesoscale gravity waves to the baroclinicity
 2478 of jet-front systems, *in press for Mon. Wea. Rev.*

- 2479 Wang, S., and F. Zhang (2007), Sensitivity of mesoscale gravity waves to the baroclinicity
2480 of jet-front systems, *Mon. Weath. Rev.*, *135*, 670–688.
- 2481 Wang, S., and F. Zhang (2010), Source of gravity waves within a vortex-dipole jet revealed
2482 by a linear model, *J. Atm. Sci.*, *67*, 1438–1455.
- 2483 Wang, S., F. Zhang, and C. Snyder (2009), Generation and propagation of inertia-gravity
2484 waves from vortex dipoles and jets, *J. Atm. Sci.*, *66*, 1294–1314.
- 2485 Wang, S., F. Zhang, and C. Epifanio (2010), Forced gravity wave response near the jet
2486 exit region in a linear model, *Q.J.R. Meteorol. Soc.*, *136*, 1773–1787.
- 2487 Warn, T., and R. Ménard (1986), Nonlinear balance and gravity-inertial wave saturation
2488 in a simple atmospheric model, *Tellus*, *38A*, 285–294.
- 2489 Warn, T., O. Bokhove, T. Shepherd, and G. Vallis (1995), Rossby number expansions,
2490 slaving principles, and balance dynamics, *Q.J.R. Meteorol. Soc.*, *121*, 723–739.
- 2491 Weglarz, R., and Y.-L. Lin (1997), Nonlinear adjustment of a rotating homogeneous
2492 atmosphere to zonal momentum forcing, *Tellus*, *50A*, 616–636.
- 2493 Williams, P., T. Haine, and P. Read (2005), On the generation mechanisms of short-scale
2494 unbalanced modes in rotating two-layer flows with vertical shear, *J. Fluid Mech.*, *528*,
2495 1–22.
- 2496 Williams, P., D. Fritts, C. She, and R. Goldberg (2006), Gravity wave propagation
2497 through a large semidiurnal tide and instabilities in the mesosphere and lower ther-
2498 mosphere during the winter 2003 MaCWAVE rocket campaign, *Annales Geophysicae*,
2499 *24*(doi:10.5194/angeo-24-1199-2006), 1199–1208.
- 2500 Williams, P., T. Haine, and P. Read (2008), Inertiagravity waves emitted from balanced
2501 flow: Observations, properties, and consequences, *J. Atmos. Sci.*, *65*(11), 3543–3556.

- 2502 Wilson, R., M.-L. Chanin, and A. Hauchecorne (1991), Gravity waves in the middle
 2503 atmosphere observed Rayleigh by Lidar: 2. Climatology, *J. Geophys. Res.*, *96*, 5169–
 2504 5183.
- 2505 Wu, D., and S. Eckermann (2008), Global gravity wave variances from Aura MLS: char-
 2506 acteristics and interpretation, *J. Atmos. Sci.*, *65*(12), 3695–3718.
- 2507 Wu, D., and J. Waters (1996), Satellite observations of atmospheric variances: A possible
 2508 indication of gravity waves, *Geophys. Res. Lett.*, *23*, 36313634.
- 2509 Wu, D., P. Preusse, S. Eckermann, J. Jiang, M. de la Torre Juarez, L. Coy, and D. Wang
 2510 (2006), Remote sounding of atmospheric gravity waves with satellite limb and nadir
 2511 techniques, *Adv. Space Res.*, *37*, 22692277.
- 2512 Wu, D. L., and F. Zhang (2004), A study of mesoscale gravity waves over the
 2513 North Atlantic with satellite observations and a mesoscale model , *J. Geophys. Res.*,
 2514 *109*(D22104), doi:10.1029/2004JD005,090.
- 2515 Wu, R., and W. Blumen (1995), Geostrophic adjustment of a zero potential vorticity flow
 2516 initiated by a mass imbalance, *J. Phys. Oceanogr.*, *25*, 439–445.
- 2517 Wunsch, C., and R. Ferrari (2004), Vertical mixing energy and the general circulation of
 2518 the oceans, *Annu. Review Fluid Mech.*, *36*, 281–314.
- 2519 Yamazaki, Y., and W. Peltier (2001a), The existence of subsynoptic-scale baroclinic insta-
 2520 bility and the nonlinear evolution of shallow disturbances, *J. Atmos. Sci.*, *58*, 657–683.
- 2521 Yamazaki, Y., and W. Peltier (2001b), Baroclinic instability in an Euler equations-based
 2522 column model: the coexistence of a deep synoptic scale mode and shallow subsynoptic
 2523 scale modes, *J. Atmos. Sci.*, *58*, 780–792.

- 2524 Yan, X., N. Arnold, and J. Remedios (2010), Global observations of gravity waves
2525 from High Resolution Dynamics Limb Sounder temperature measurements: A year-
2526 long record of temperature amplitude and vertical wavelength, *J. Geophys. Res.*,
2527 *115*(D10113), doi:10.1029/2008JD011,511.
- 2528 Yavneh, I., J. McWilliams, and M. Molemaker (2001), Non-axisymmetric instability of
2529 centrifugally stable stratified taylor-couette flow, *J. Fluid Mech.*, *448*, 1–21.
- 2530 Yeh, T. (1949), On energy dissipation in the atmosphere, *J. Meteor.*, *6*, 1–16.
- 2531 Young, W., and M. B. Jelloul (1997), Propagation of near-inertial oscillations through a
2532 geostrophic flow, *J. Mar. Res.*, *55*, 735–766.
- 2533 Zeitlin, V. (2008), Decoupling of balanced and unbalanced motions and inertia-gravity
2534 wave emission: Small versus large rossby numbers, *J. Atmos. Sci.*, *65*(11), 3528–3542.
- 2535 Zeitlin, V., S. Medvedev, and R. Plougonven (2003), Frontal geostrophic adjustment, slow
2536 manifold and nonlinear wave phenomena in one-dimensional rotating shallow-water.
2537 Part 1: Theory, *J. of Fluid Mech.*, *481*, 269–290.
- 2538 Zhang, D.-L., and J. Fritsch (1988), Numerical simulation of the meso-b scale structure
2539 and evolution of the 1977 Johnstown flood. Part III: Internal gravity waves and the
2540 squall line, *J. Atmos. Sci.*, *45*, 12521268.
- 2541 Zhang, F. (2004), Generation of mesoscale gravity waves in upper-tropospheric jet-front
2542 systems, *J. Atmos. Sci.*, *61*(4), 440–457.
- 2543 Zhang, F., and S. Koch (2000), Numerical simulations of a gravity wave event over
2544 CCOPE. Part II: Waves generated by an orographic density current, *Mon. Weath.*
2545 *Rev.*, *128*(8), 2777–2796.

- 2546 Zhang, F., S. Koch, C. Davis, and M. Kaplan (2000), A survey of unbalanced flow diag-
 2547 nostics and their application, *Adv. Atmos. Sci.*, *17*(2), 165–183.
- 2548 Zhang, F., S. Koch, C. Davis, and M. Kaplan (2001), Wavelet analysis and the governing
 2549 dynamics of a large amplitude mesoscale gravity wave event along the east coast of the
 2550 united states, *Q.J.R. Meteorol. Soc.*, *127*, 2209–2245.
- 2551 Zhang, F., S. Koch, and M. Kaplan (2003), Numerical simulations of a large-amplitude
 2552 gravity wave event, *Meteo. Atmos. Phys.*, *84*, 199–216.
- 2553 Zhang, F., S. Wang, and R. Plougonven (2004), Potential uncertainties in using the hodo-
 2554 graph method to retrieve gravity wave characteristics from individual soundings, *Geo-
 2555 phys. Res. Lett.*, *31*(L11110), doi:10.1029/2004GL019,841.
- 2556 Zhang, F., N. Bei, R. Rotunno, and C. Snyder (2007), Mesoscale predictability of moist
 2557 baroclinic waves: Convection permitting experiments and multistage error growth dy-
 2558 namics, *J. Atmos. Sci.*, *64*, 3579–3594.
- 2559 Zhang, F., M. Zhang, K. Bowman, L. Pan, and E. Atlas (2009), Aircraft measurements
 2560 and numerical simulations of gravity waves in the extratropical utls region during the
 2561 start08 field campaign, in *The 13th Conference on Mesoscale Processes*.
- 2562 Zhang, F., J. Wei, and S. Wang (2011), Dynamics and impacts of gravity waves in the
 2563 baroclinic jet-front systems with moist convection, in *The 14th AMS conference on
 2564 mesoscale processes*.
- 2565 Zhang, S., and F. Yi (2005), A statistical study of gravity waves from radiosonde obser-
 2566 vations at Wuhan (30 degrees N, 114 degrees E) China, *Ann. Geophys.*, *23*, 665–673.
- 2567 Zhang, S., and F. Yi (2007), Latitudinal and seasonal variations of inertial gravity wave
 2568 activity in the lower atmosphere over central China, *J. Geophys. Res.*, *112*(D05109),

2569 doi:10.1029/2006JD007,487.

2570 Zhang, S., and F. Yi (2008), Intensive radiosonde observations of gravity waves in the
2571 lower atmosphere over Yichang (111 degrees 18 ' E, 30 degrees 42 ' N), China, *Ann.*
2572 *Geophys.*, *26*(7), 2005–2018.

2573 Zhu, X., and J. Holton (1987), Mean fields induced by local gravity-wave forcing in the
2574 middle atmosphere, *J. Atmos. Sci.*, *44*(3), 620–630.

2575 Zülicke, C., and D. Peters (2006), Simulation of inertia-gravity waves in a poleward break-
2576 ing Rossby wave, *J. Atmos. Sci.*, *63*, 3253–3276.

2577 Zülicke, C., and D. Peters (2008), Parameterization of strong stratospheric inertigravity
2578 waves forced by poleward-breaking rossby waves, *Mon. Wea. Rev.*, *136*, 98–119.

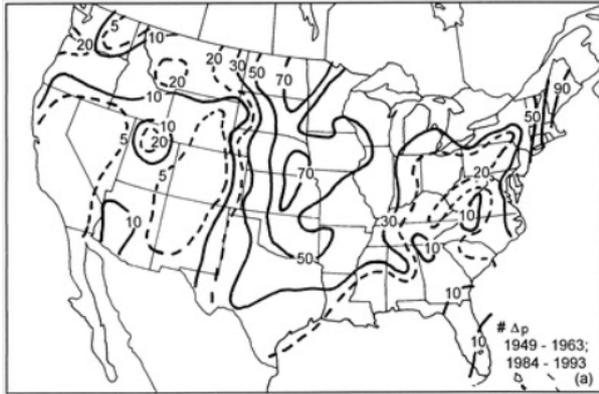


Figure 1. Distribution of large hourly surface pressure changes (defined to be greater than 4.25 hPa.), as diagnosed from the surface barograph network by *Koppel et al.* [2000]. The data covers 25 years (1949 – 1963 and 1984 – 1993).

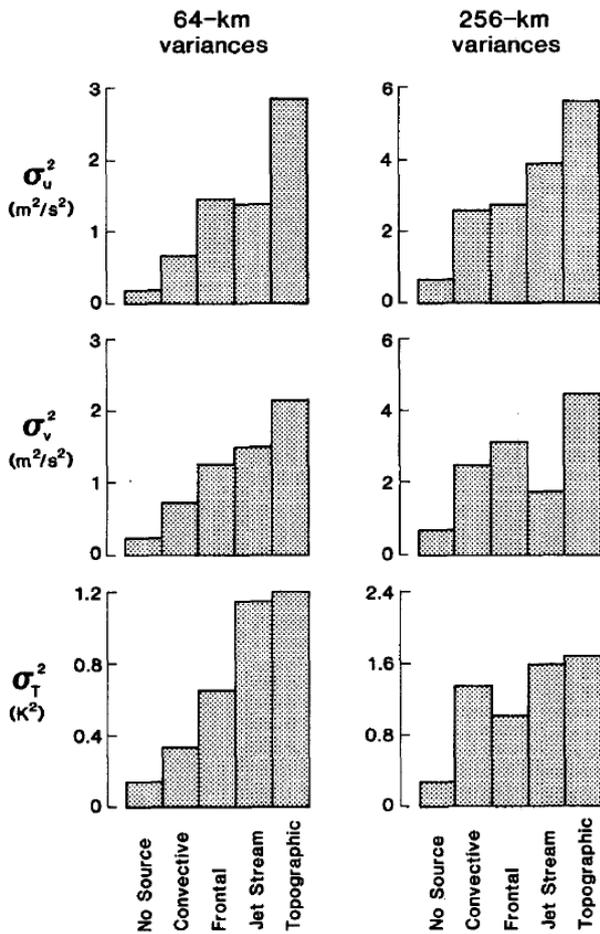


Figure 2. Average variances of the zonal (top) and meridional (middle) wind components, and of temperature (bottom), for flight segments of 64 (left) and 256 km (right). Inspection of the flow has allowed to categorize segments by the expected source of gravity waves. (Adapted from *Fritts and Nastrom* [1992]).

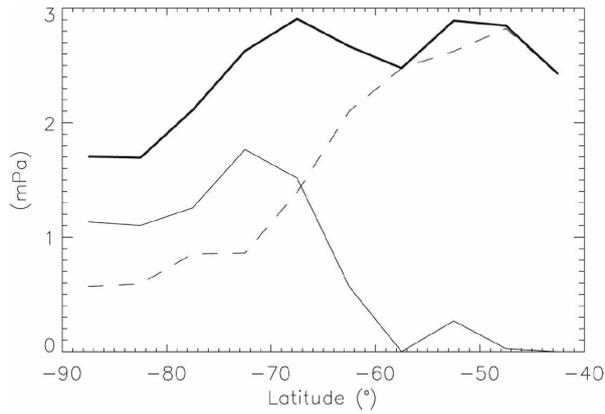


Figure 3. Latitudinal distribution of zonal mean, density weighted absolute momentum flux carried by waves over orographic regions (thin solid), by waves over non-orographic regions (thin dashed), and by both types of waves (thick solid), as estimated by [*Hertzog et al.*, 2008] from balloon observations.

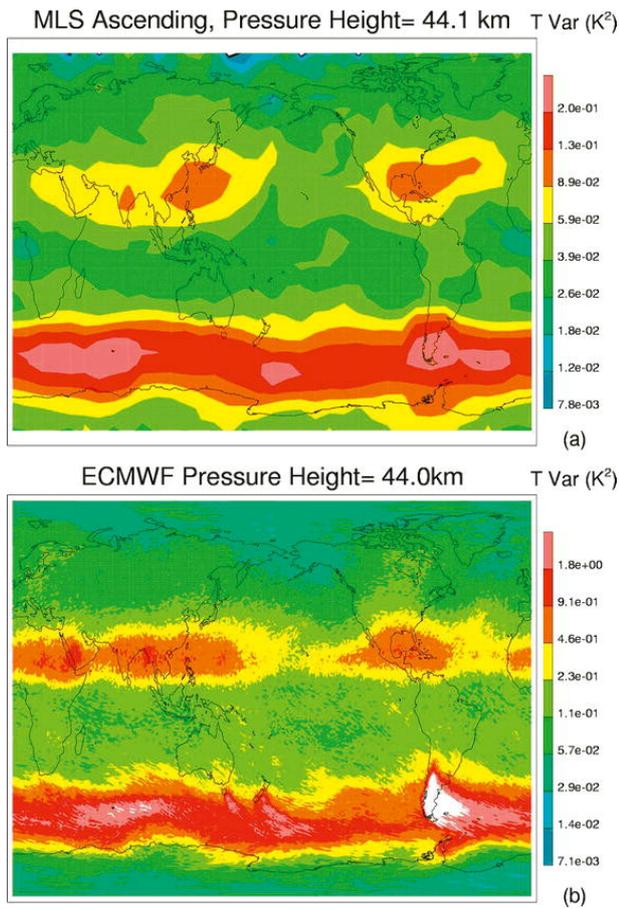


Figure 4. Monthly-mean temperature variances at 44-km pressure altitude from from (a) satellite observations from the Aura Microwave Limb Sounder, and (b) the ECMWF analyses at resolution TL799L91 for August 2006. For the latter, only horizontal wavelengths longer than 300 km were retained.

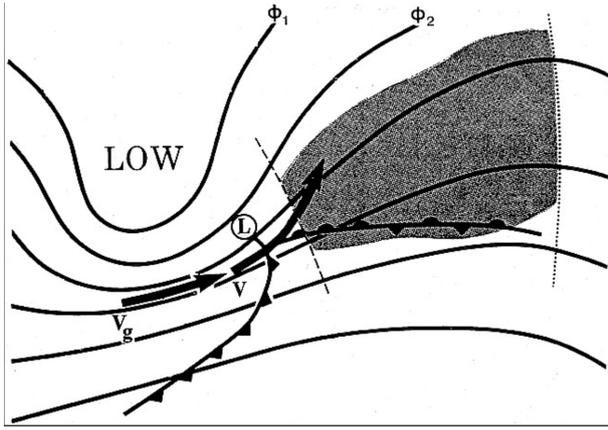


Figure 5. Flow configuration identified by *Uccellini and Koch* [1987] (UK87) as conducive to intense gravity waves: lines of geopotential in the mid-troposphere and surface fronts are indicated. Just downstream of the inflection axis (dashed line) the wind has a significant cross-stream ageostrophic component (wind vector crossing isolines of geopotential) and intense gravity waves are recurrently found (shaded region).

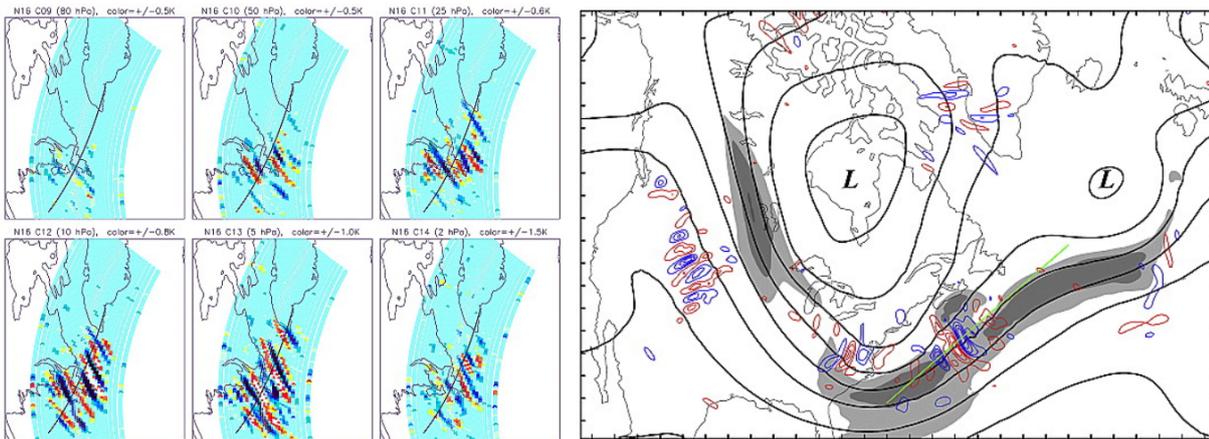


Figure 6. Comparison of gravity waves in satellite observations and in mesoscale simulations, from *Wu and Zhang* [2004]. Left panel: radiance perturbations from different channels of the NOAA 16 AMSU-A at 0630 UT on 20 January, showing gravity wave perturbations at different heights. Right panel: geopotential height (thick contours every 20 dam) and maxima of wind speed (shaded regions) at 300-hPa, and 80-hPa horizontal divergence (every $3 \times 10^{-5} \text{ s}^{-1}$; blue, positive; red, negative) from the MM5 simulations at 1800 UT on 19 January (starting on 19 January at 0000 UT). Simulated amplitudes of wind and temperature perturbations are 10 m s^{-1} and 5 K respectively.

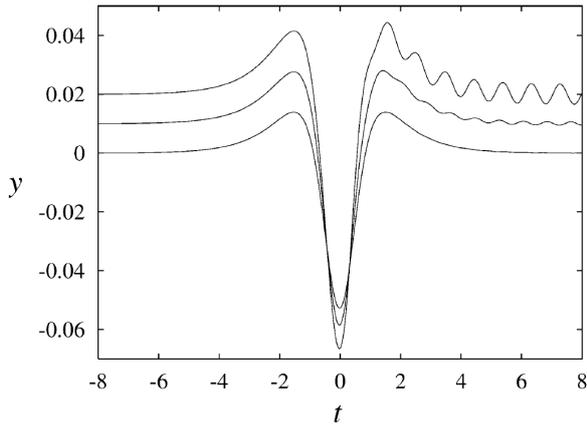


Figure 7. Evolution of $y(t)$, one of the 2 fast variables of the Lorenz-Krishnamurty model, as calculated by *Vanneste* [2004], for Rossby numbers $\epsilon = 0.15$ (upper curve, offset by 0.02), $\epsilon = 0.125$ (middle curve, offset by 0.01) and $\epsilon = 0.1$ (lower curve). The balanced evolution of the flow leads to temporary variations of y near $t = 0$. For $\epsilon = 0.15$, conspicuous fast oscillations are excited and remain thereafter. This emission is very sensitive to ϵ (exponential dependence).

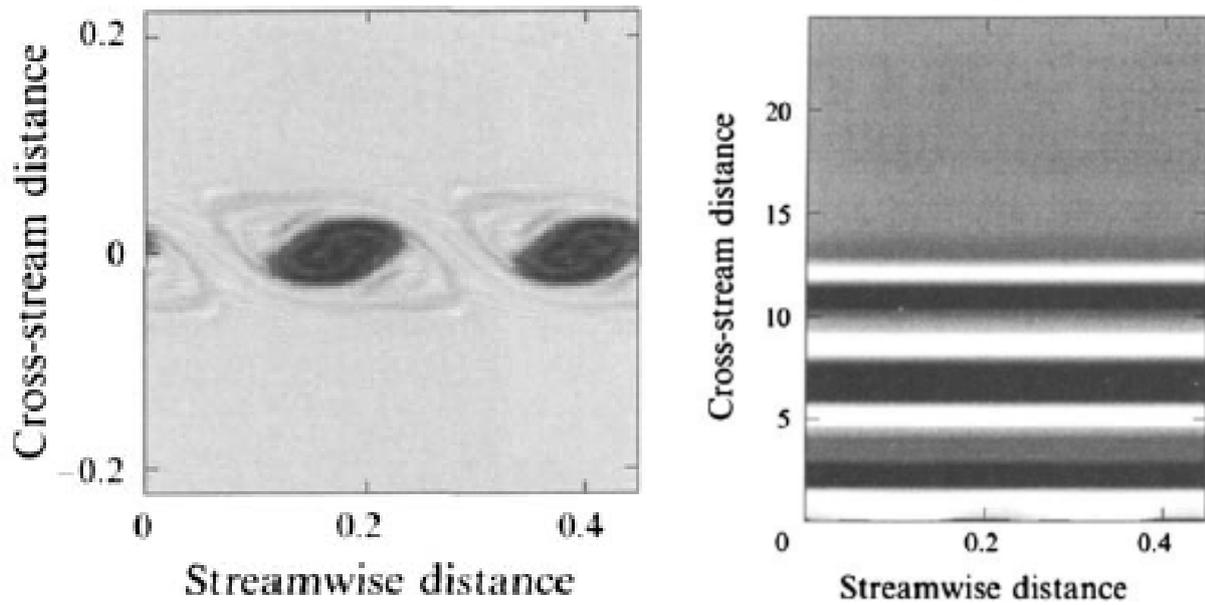


Figure 8. From *Ford* [1994c]: roll-up of the unstable potential vorticity strip (left) as seen from the potential vorticity distribution, and radiation of gravity waves in the far-field (right), as seen from the time derivative of the surface height. Note the large scale separation between the two phenomena.

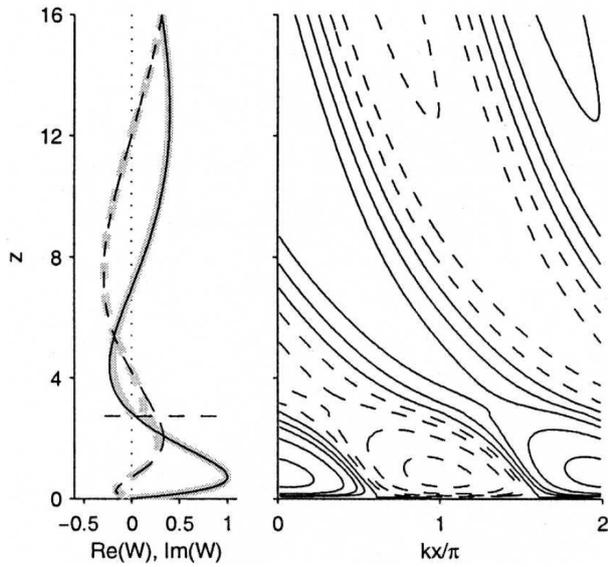


Figure 9. Vertical structure $W(z)$ for a normal mode of an unbalanced baroclinic instability in a vertical shear [Plougonven *et al.*, 2005]: the left panel shows the real (plain line) and imaginary parts of $W(z)$, with the horizontal dashed line indicating the inertial critical level. The right panel shows a vertical cross-section in the (x, z) plane, through one wavelength of the mode. Also shown in the left panel are asymptotic approximations of the balanced edge wave near the surface (below the ICL, obtained asymptotically in Rossby number), and a far field approximation of sheared gravity waves aloft (above the ICL).

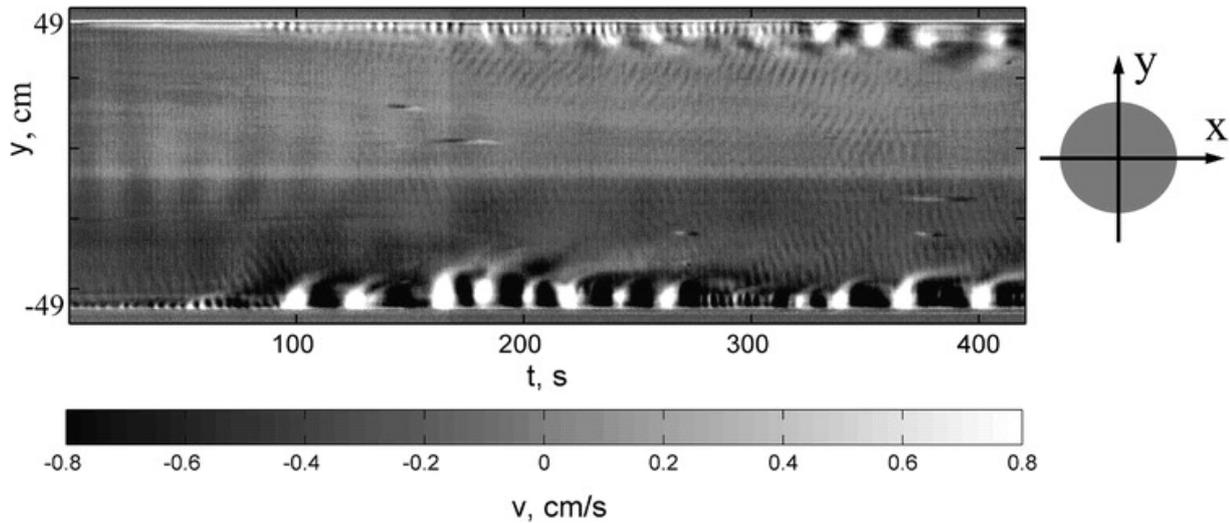


Figure 10. Hovmöller plot showing the x component of the gradient wind velocity along the y axis across the tank in the experiment of *Afanasyev et al.* [2008]. Features near the walls ($y = \pm 49\text{cm}$) describe the baroclinic instability of the coastal jet. The intentionally narrow grayscale range makes the short-scale inertial waves visible. They are emitted from the shorter-scale meanders of the coastal jet and propagate into the quiescent interior of the tank.

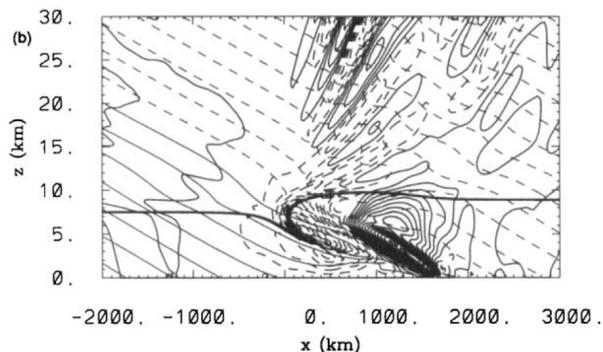


Figure 11. Isotachs of vertical velocity (thick lines, contour interval 5 m s^{-1}) in the two-dimensional simulation of frontogenesis of *Griffiths and Reeder* [1996] which produced the most stratospheric waves. Also shown is the tropopause (thick line) and the cross-front velocity (contour interval 5 m s^{-1}).

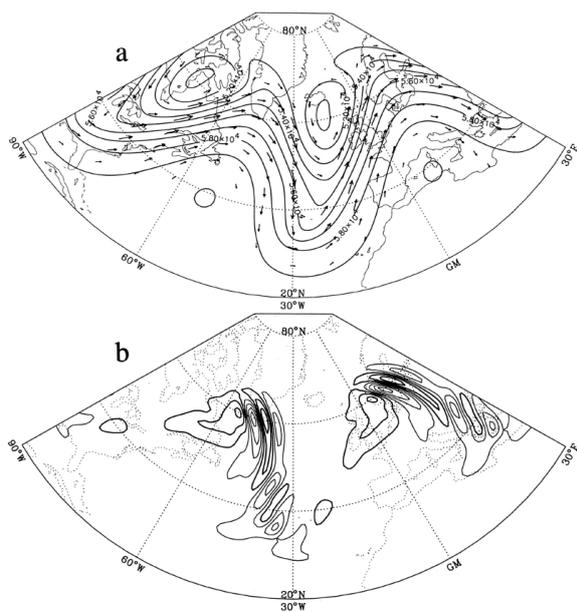


Figure 12. a) Geopotential height and wind at 503 hPa , at day 10 of the idealized baroclinic life cycle of *O’Sullivan and Dunkerton* [1995], and b) divergence of the horizontal wind at 130 hPa at the same time. Adapted from *O’Sullivan and Dunkerton* [1995].

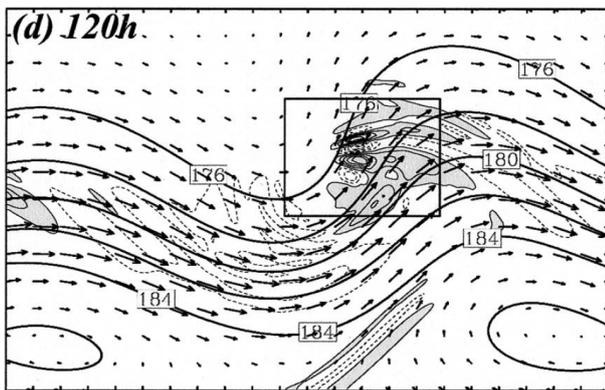


Figure 13. Pressure and divergence of horizontal wind, at $z = 13\text{km}$, in the baroclinic life cycle simulated by *Zhang* [2004]. Distance between the tick marks is 300 km.

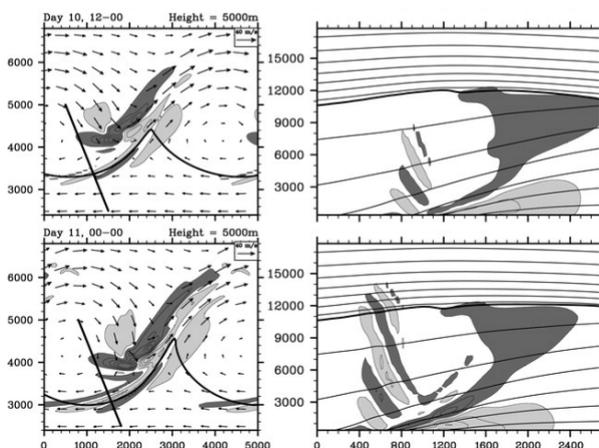


Figure 14. Inertia-gravity waves appearing ahead of cold surface fronts in a life cycle with enhanced surface anticyclonic shear. Left: horizontal maps of $\nabla \mathbf{u}_H$ at $z = 5\text{km}$, with one surface isentrope (thick line) to depict the surface fronts; right: vertical cross sections through the line segments indicated in the left panels (height in m, distance along section in km, southern end of the section to the left). The top and bottom panels are separated by 12 hours. Adapted from *Plougonven and Snyder* [2007].

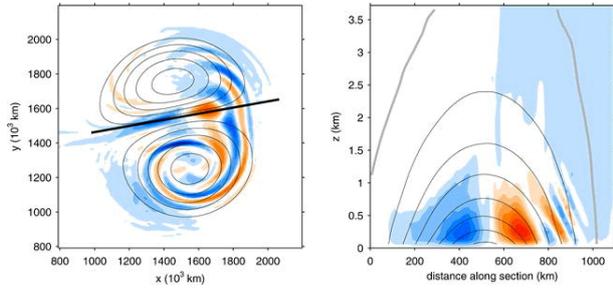


Figure 15. Horizontal (left) and vertical (right) cross-sections of the vertical velocity (colors) in a surface dipole, from *Snyder et al.* [2007]. Also shown are contours of potential temperature (left, at $z = 125m$) and of section-parallel horizontal flow. The horizontal cross-section of w corresponds to $z = 62.5m$.